Proper motion and apparent contraction in the CSO J0650+6001

M. Orienti^{1,2}, D. Dallacasa^{1,2}

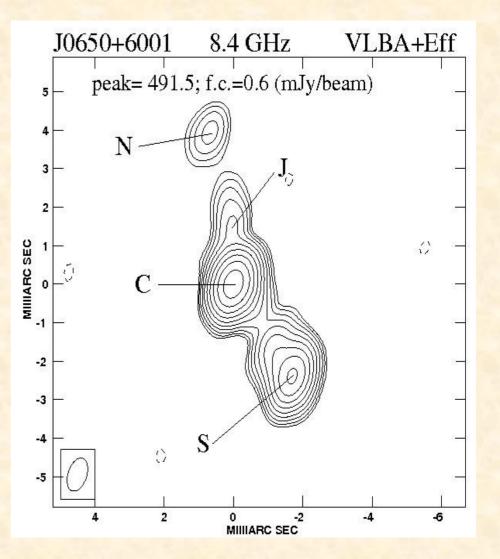
Astronomy Dept., University of Bologna, Italy
INAF - IRA Bologna, Italy

Compact Symmetric Objects

- Compact size (< 1 kpc);
- High radio luminosity ($P_{1.4 \text{ GHz}} > 10^{24} \text{ W/Hz}$);
- Significant fraction in radio source catalogues selected at 5 GHz (~15-30%);
- Radio synchrotron spectra with a turnover frequency ranging from ~ 100 MHz to a few GHz.

The radio source J0650+6001

- Quasar @ z=0.455;
- Spectral peak at 5.5 GHz;
- Total linear size LLS=40 pc;
- Very asymmetric triple radio moprhology;
- 70% of the total flux arises from the central component;
- $S_{\rm s}^{\rm S}/S_{\rm N}^{\rm s} \sim 30;$
- Misalignment between the outer components.



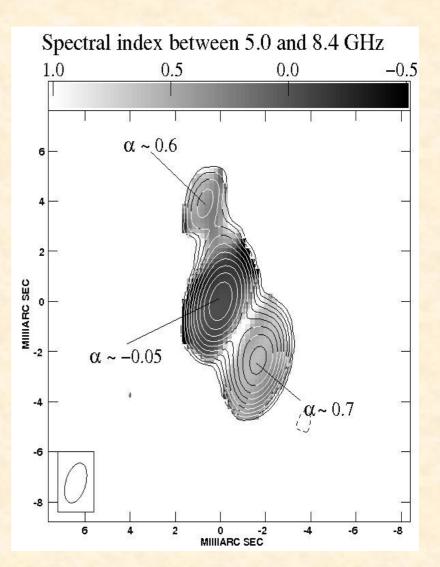
The radio source J0650+6001

• Central component with flat spectrum $\alpha = -0.05$

presence of the **source core**;

• Outer components with steep spectra $\alpha = 0.6 - 0.7$

hotspots and mini-lobes



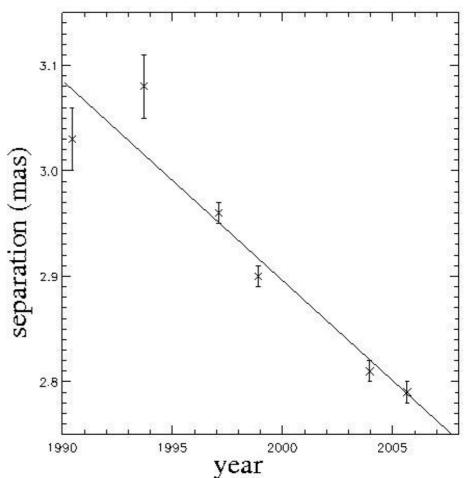
Proper motion & apparent contraction

 Components N and S are separating with
v_{is}~0.39c±0.18c

 $t_{kin} \sim 360 \pm 170 \text{ yr};$

 Component C and S are decreasing their distance with an apparent contraction velocity v_{a,c}~0.37c±0.02c





Beaming effects

From the source expansion:

 $eta_{\mathrm{s},\mathrm{s}} = rac{2eta_{\mathrm{s},\mathrm{i}}\mathrm{sin}m{ heta}}{1-eta_{\mathrm{s},\mathrm{i}}^2\mathrm{cos}^2m{ heta}}$

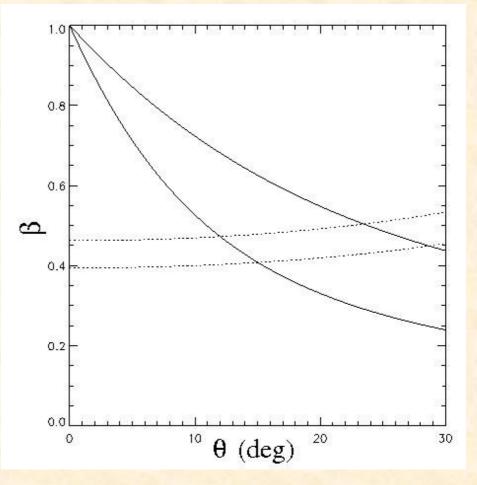
• From the flux density ratio:

$$\frac{S_{\rm S}}{S_{\rm N}} = R = \left(\frac{1+\beta{\rm cos}\theta}{1-\beta{\rm cos}\theta}\right)^{3+\alpha}$$

The possible $(\theta - \beta)$ combinations are:

$$v_{i,s} = 0.43c \pm 0.04c$$

 $12^{\circ} < \theta < 28^{\circ}$



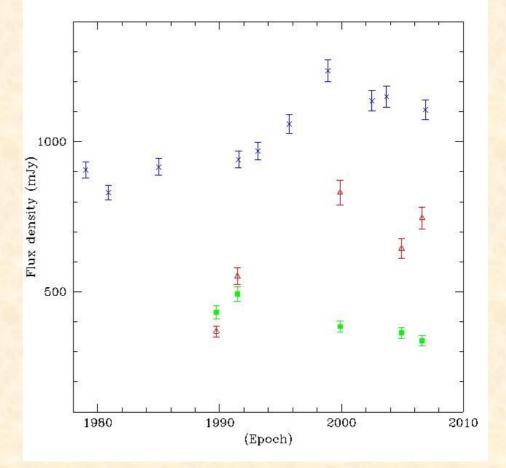
Flux density variability

• The 5 GHz lightcurve shows a steady increment of the flux densities S_{tot} and S_C

Source expansion;

From the beaming model, the distance between C and S decreases with:

v_{i,c}~0.66c±0.03c



• No variability in S_s



- J0650+6001 has a peculiar non-aligned asymmetric radio structure dominated by the central component;
- The outer components are increasing their distance with $v_{a,s}=0.39c\pm0.18c$, that corresponds to a source age $t_{kin}=360\pm170$ yr;
- Apparent contraction between components C and S, suggesting the presence of a knot in the jet that is moving from C towards S
- The total flux density and the flux density of component C are steadily increasing, in agreement with what expected in source expansion;
- The peculiar properties of this source may be explained in terms of Doppler beaming effects where the jets are separating at a mildly relativistic velocity, and oriented close to the line of sight.