

Multi-frequency, radio circular and linear polarization monitoring of OJ 287

Ioannis Myserlis, E. Angelakis, S. Komossa
Max-Planck-Institut für Radioastronomie



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OJ 287

short profile

$z = 0.306$

$\theta_{\text{obs}} \sim 3^\circ$

$M \sim 2 \times 10^{10} M_\odot$

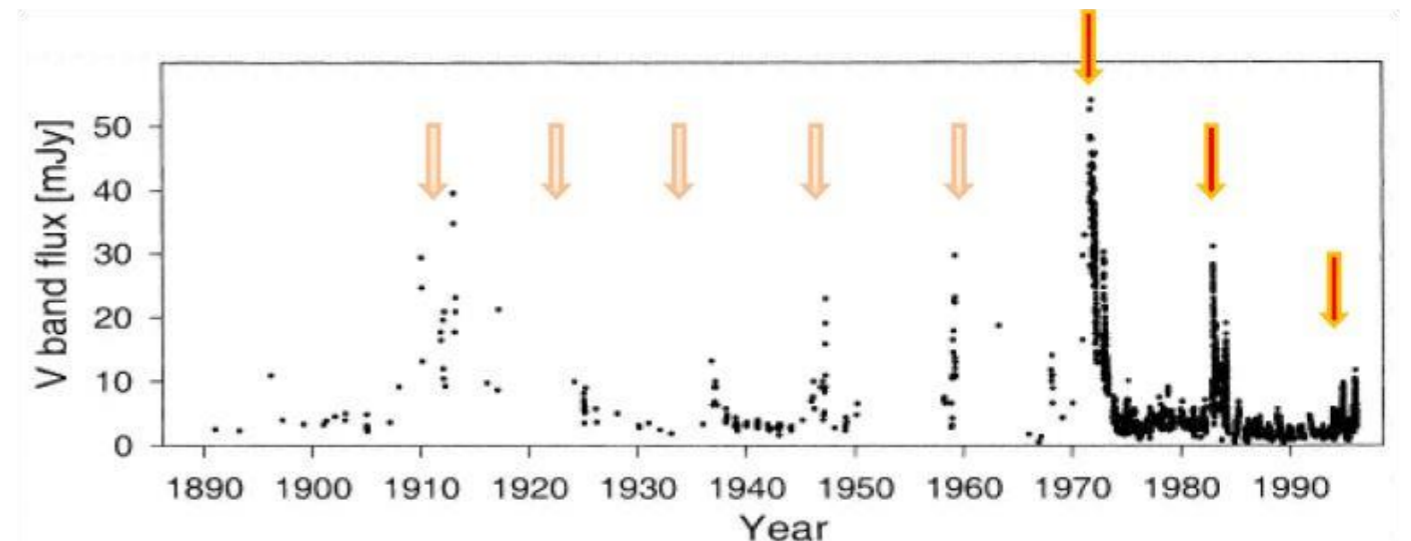
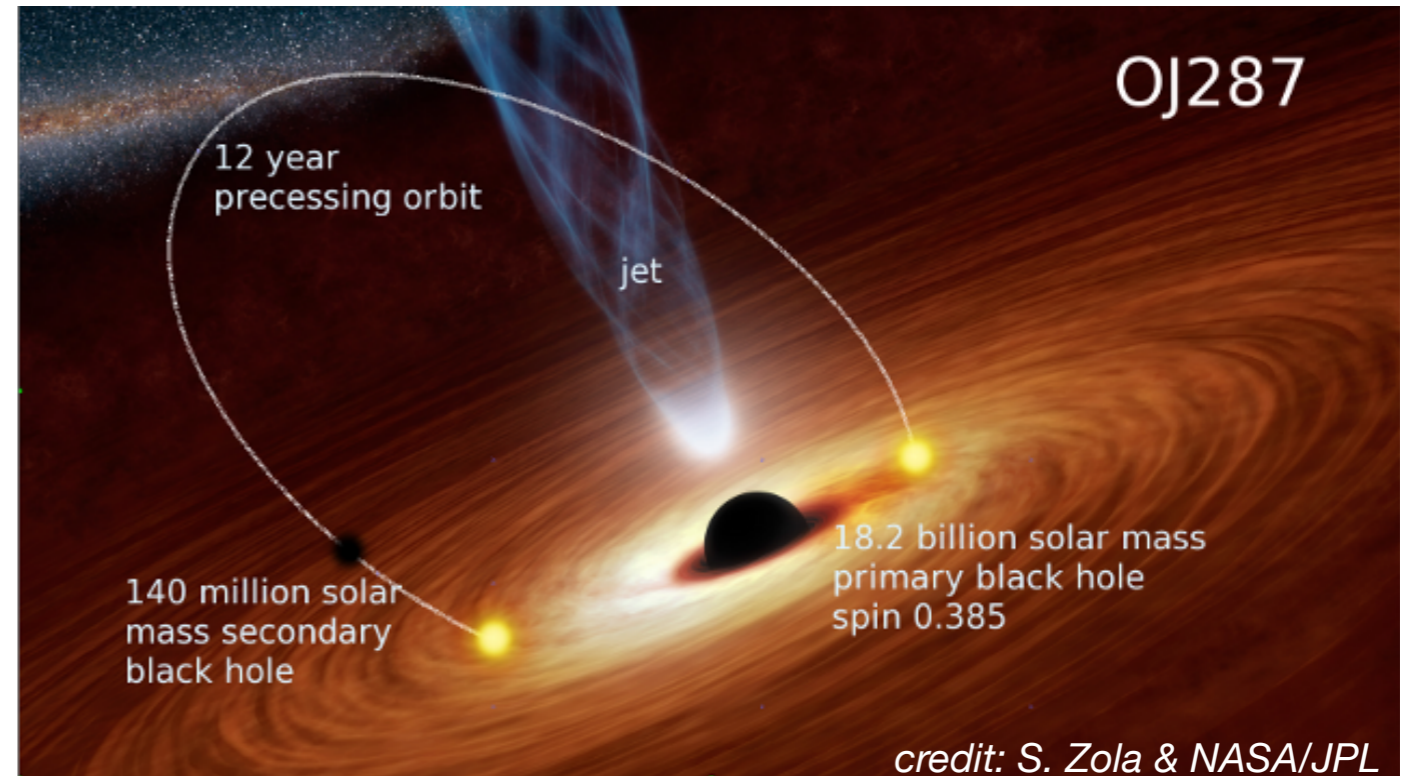
SMBBH candidate

- galaxy formation and evolution
 - BH demographics and growth
- Extremely close merger

Dense optical data

Sparse radio data

(esp. around opt. maxima)



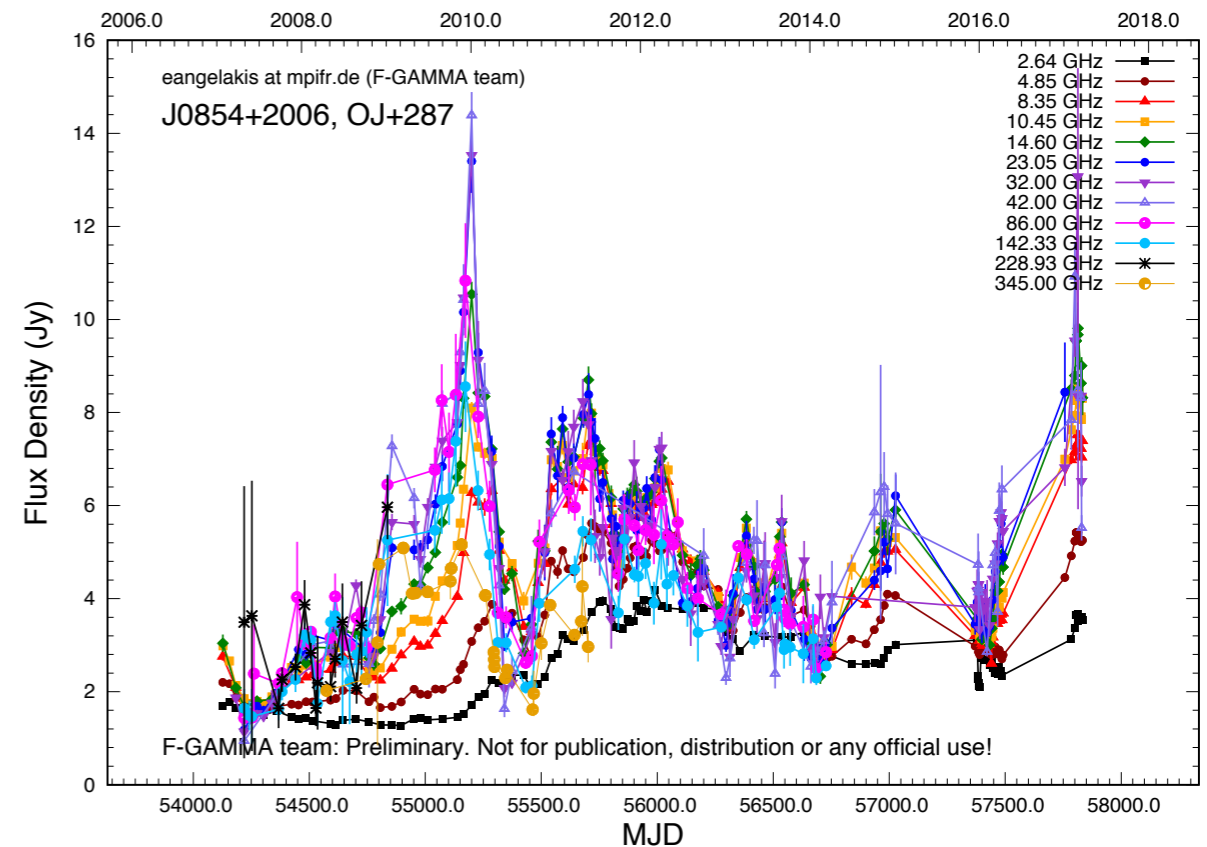
Ongoing, full-Stokes, high-cadence monitoring

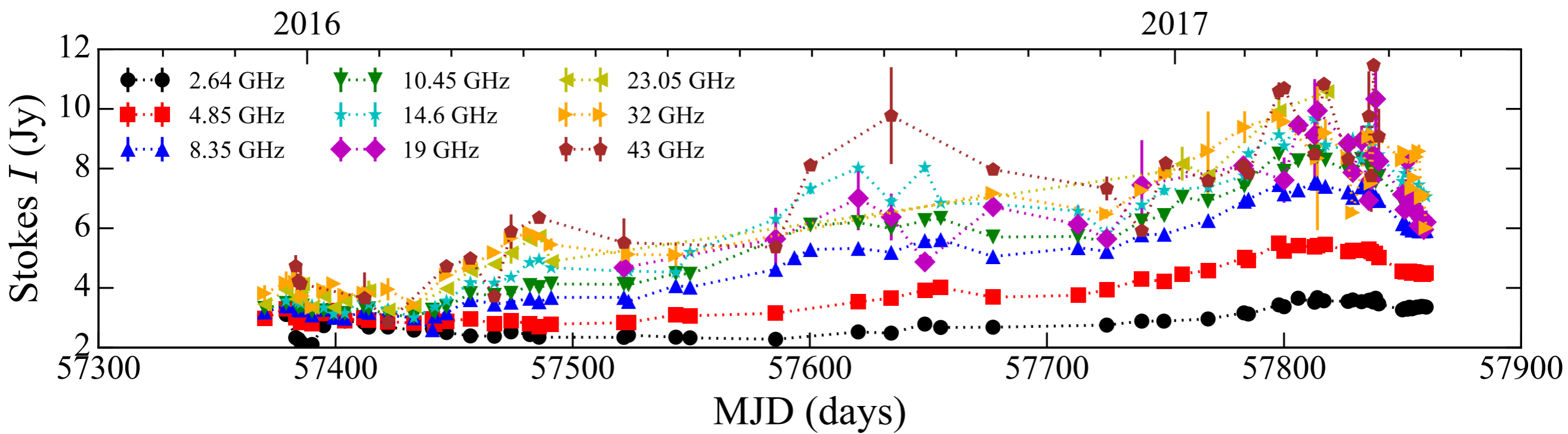
F-GAMMA

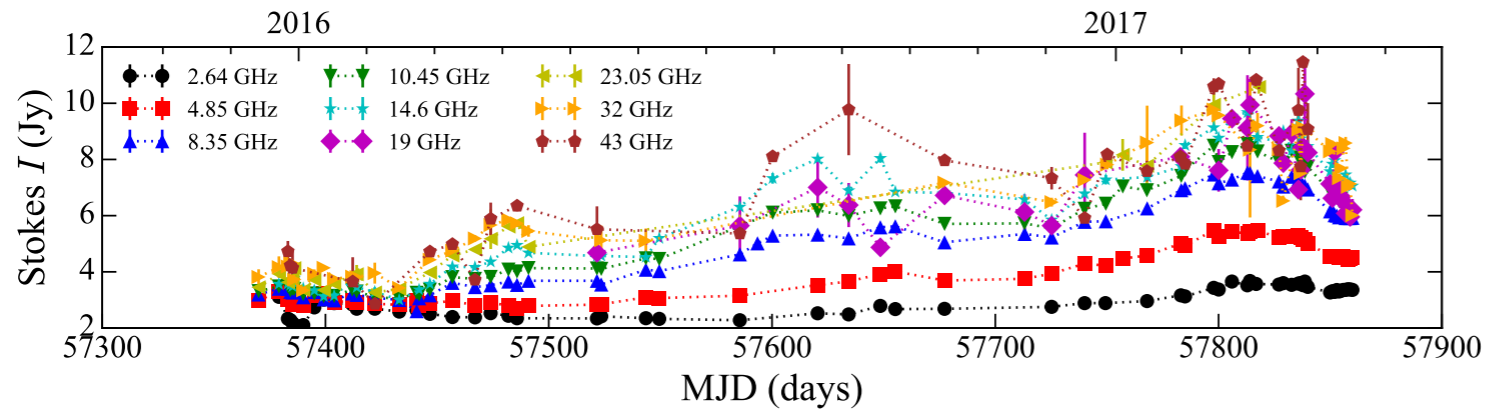
- Jan 2007 — Jan 2015
- mean cadence: 1.3 months
- 2.64–345 GHz at 12 frequency steps
- LP ($\sim 2.5\%$) and CP ($\sim 0.2\%$)

Dec 2015 - now (~ 500 days)

- mean cadence: 8 days
- 2.64–43 GHz at 8 frequency steps
- New full-Stokes data analysis pipeline
[Myserlis et al. 2017, A&A, in press](#)
- LP at 2.64, 4.85, 8.35 and 10.45 GHz
- CP at 4.85 and 8.35 GHz







Stokes *I*

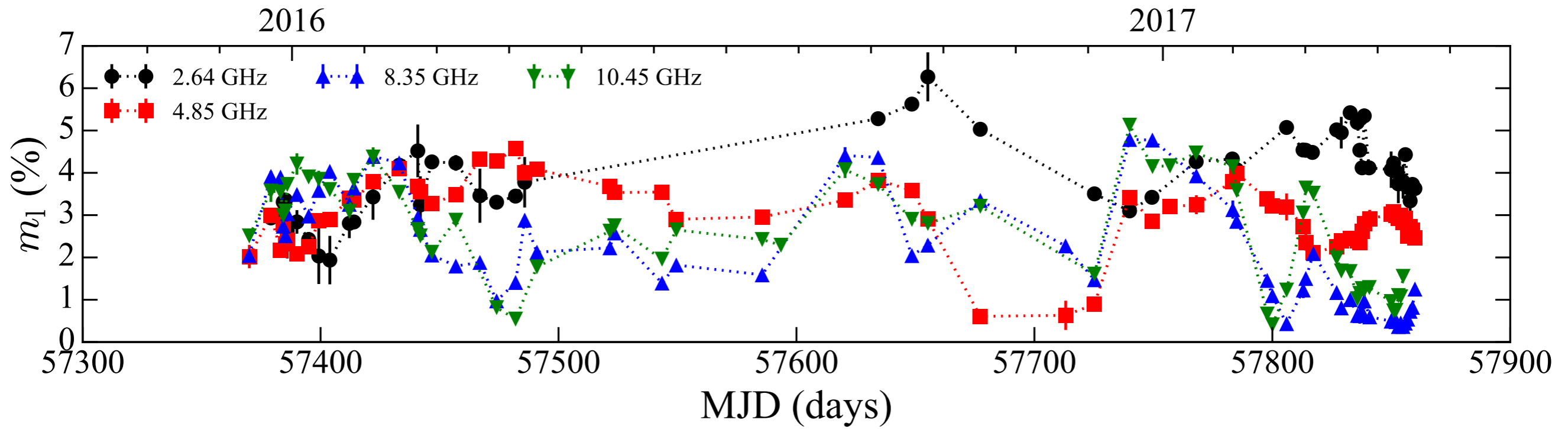
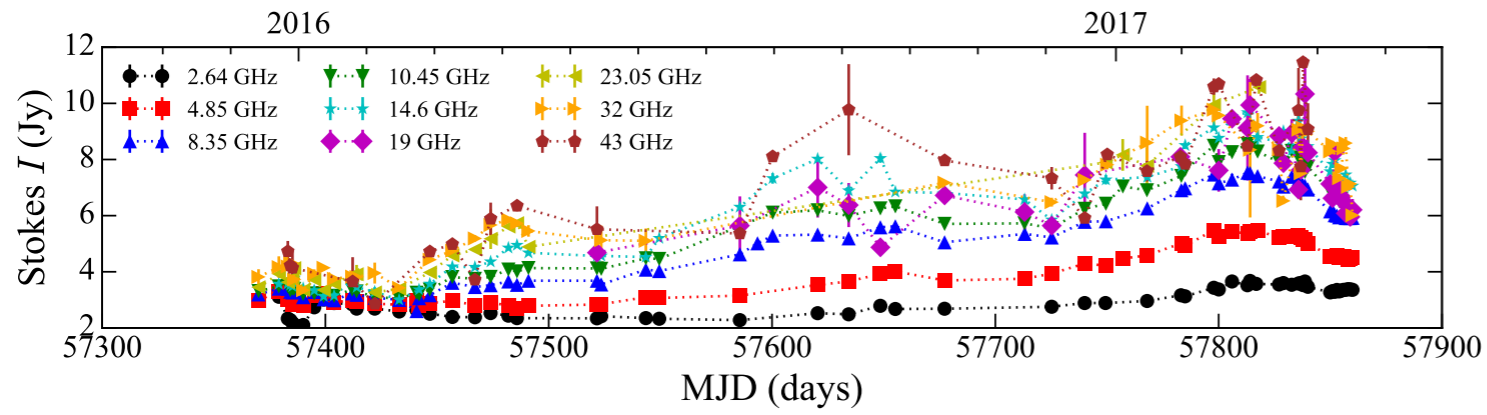
range: 2–11.5 Jy

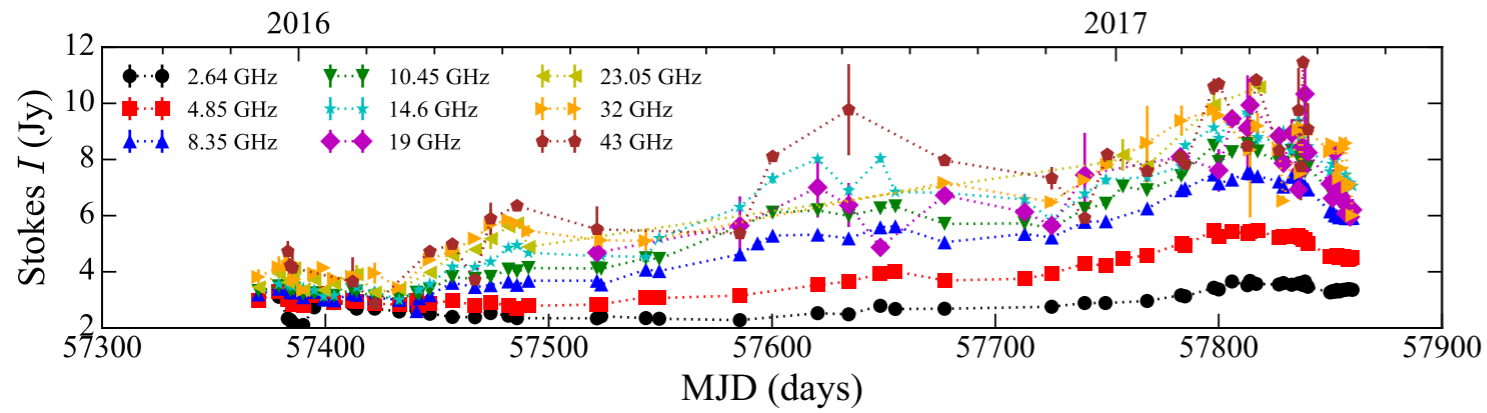
activity increased since mid-Feb 2016

- correlated with frequency

spectrum always inverted

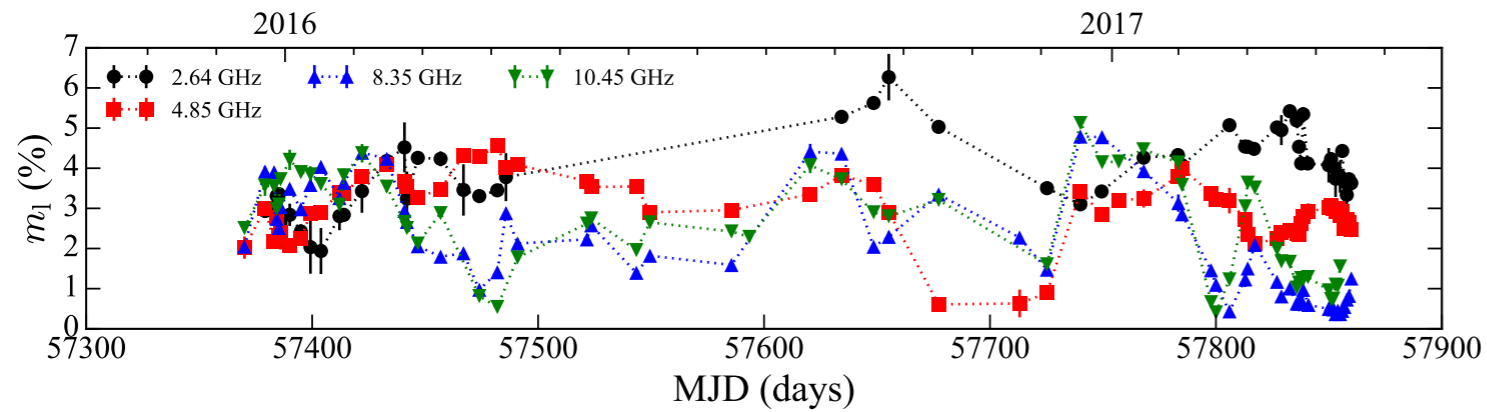
(highest) maximum: mid-Feb to mid-Mar 2017





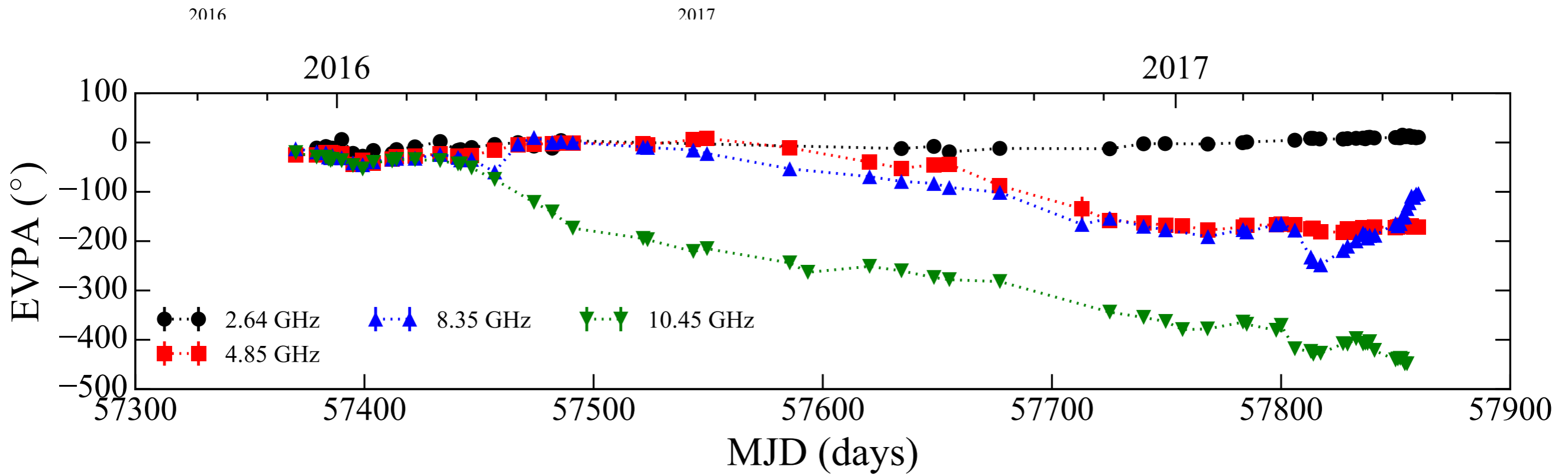
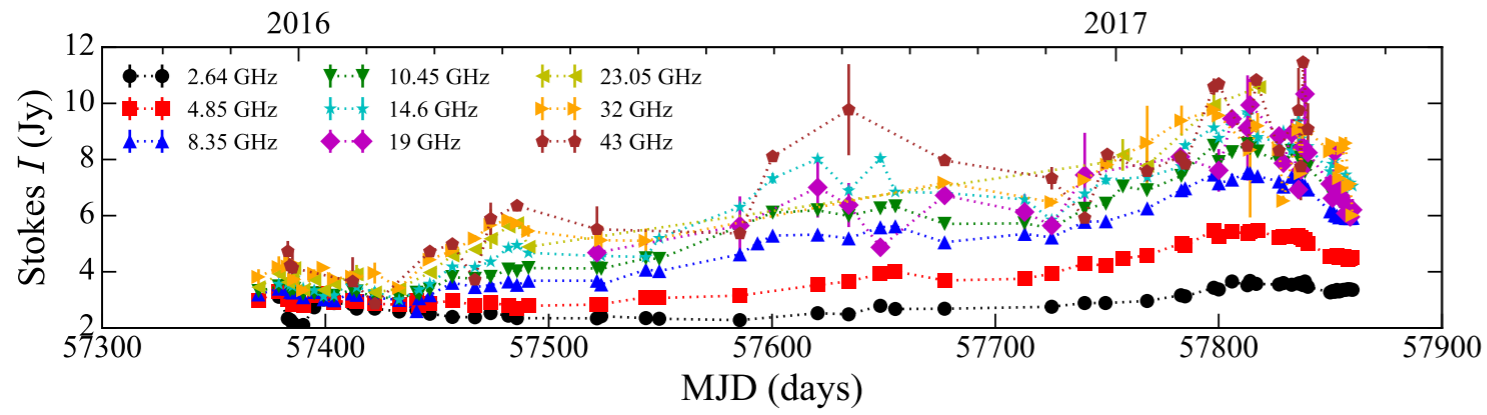
m_1

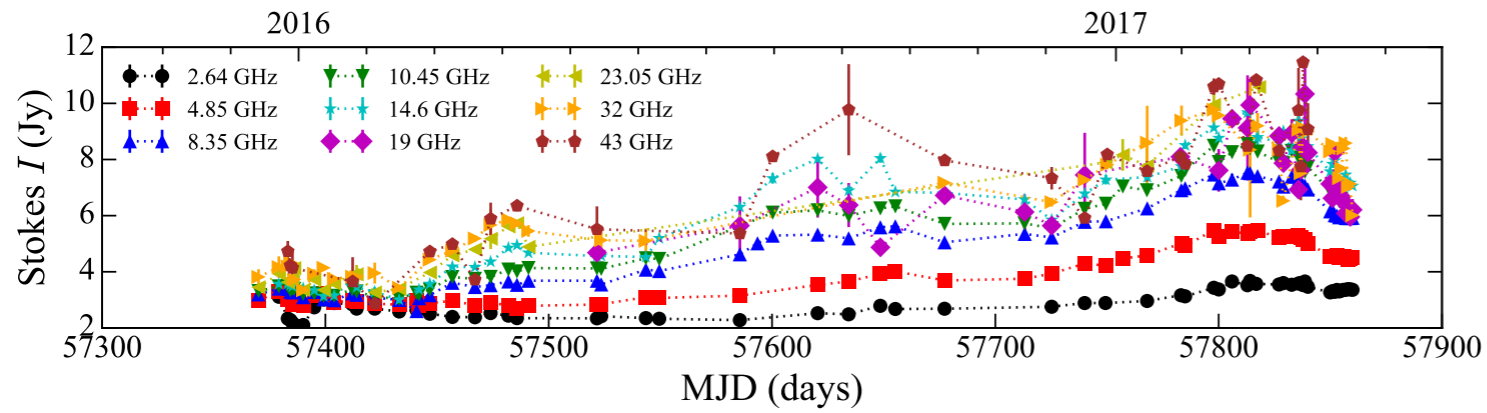
range: 0.4–6.3 %
median: 2.7 %



$\langle m_1 \rangle$ anti-correlated with freq.
variations correlated with freq.

simultaneous m_1 and I peaks
- except around MJD 57800





EVPA

CCW rotations

amplitude

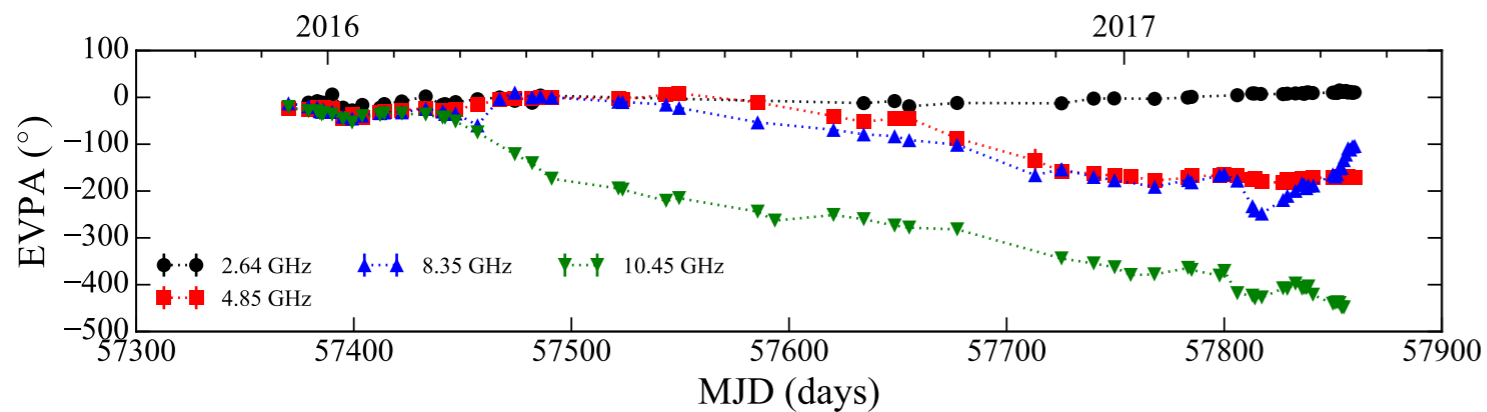
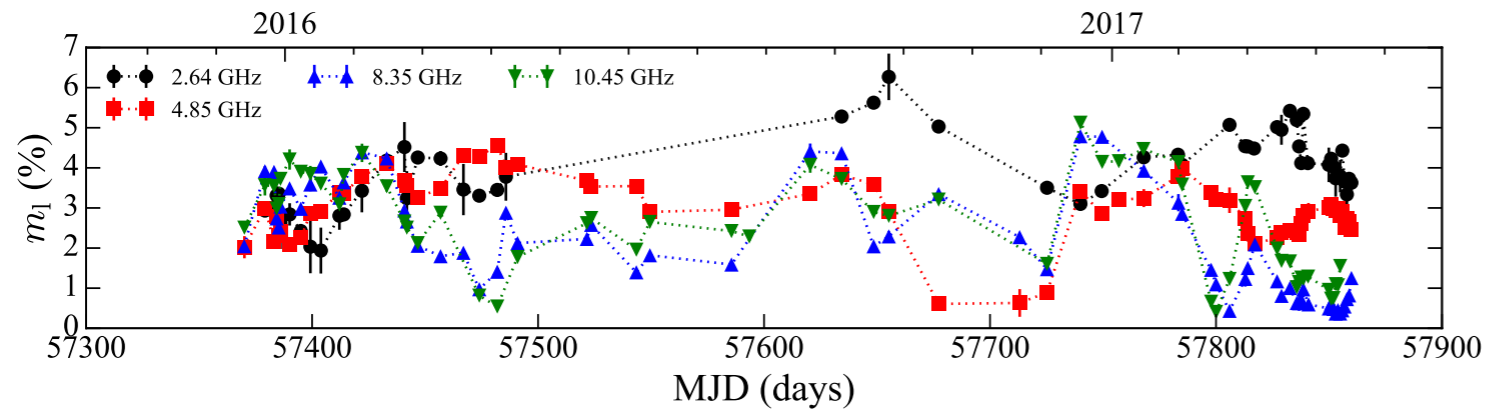
- correlated with freq.
- range: $\sim 190^\circ - 430^\circ$

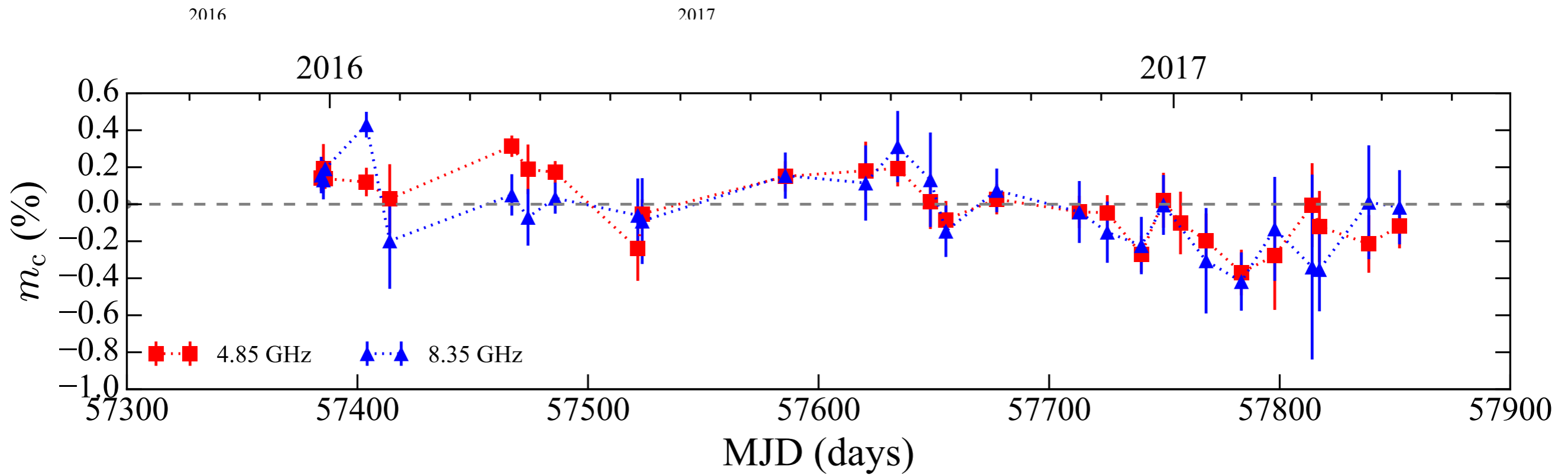
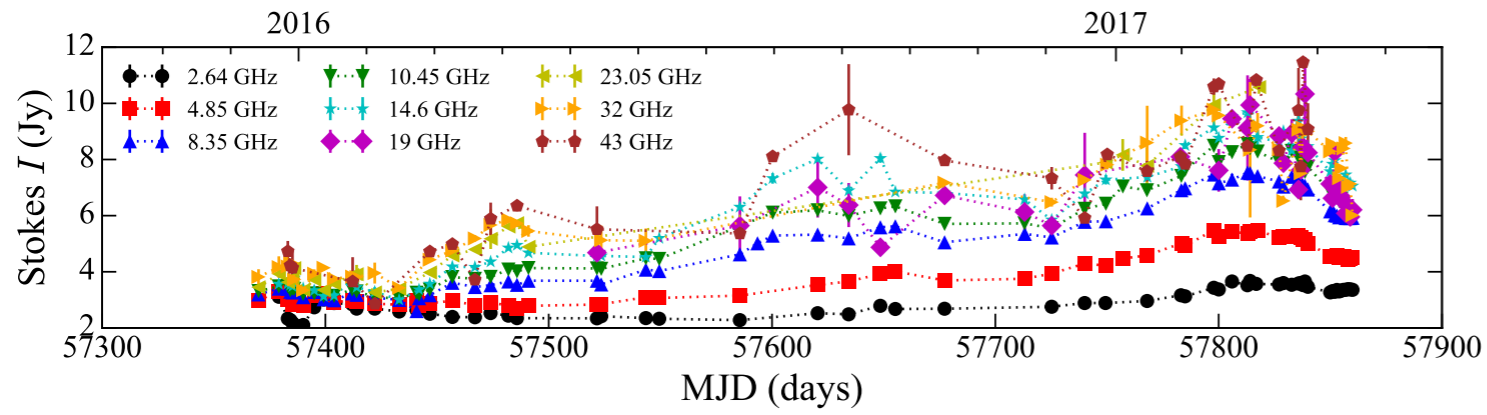
rate

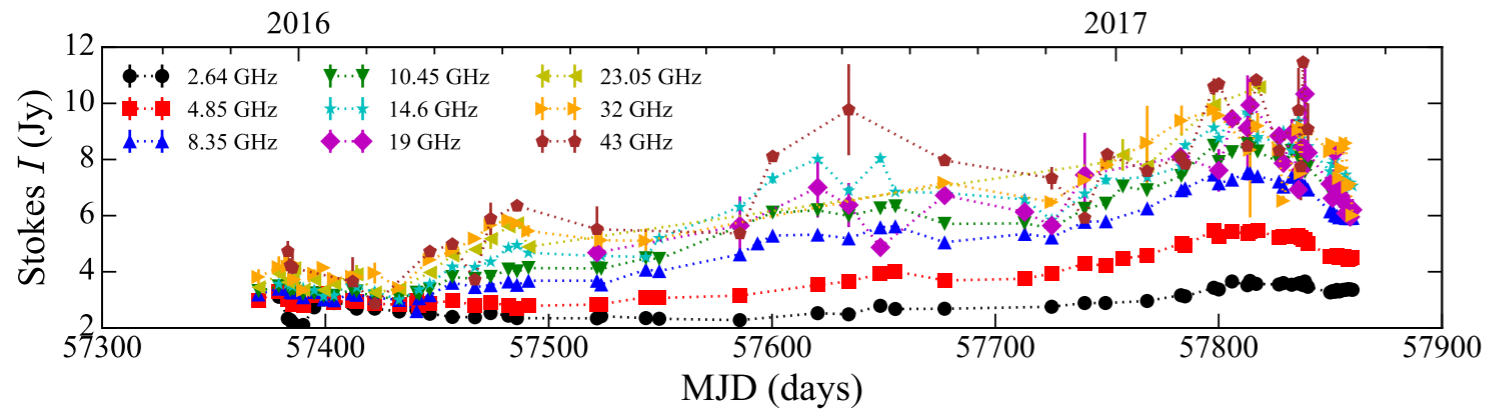
- correlated with freq.
- range: $0.5 - 1.2^\circ/\text{day}$

rotation evolution

- smooth until Feb 2017
- opposite senses btw 8.35 and 10.45 GHz afterwards

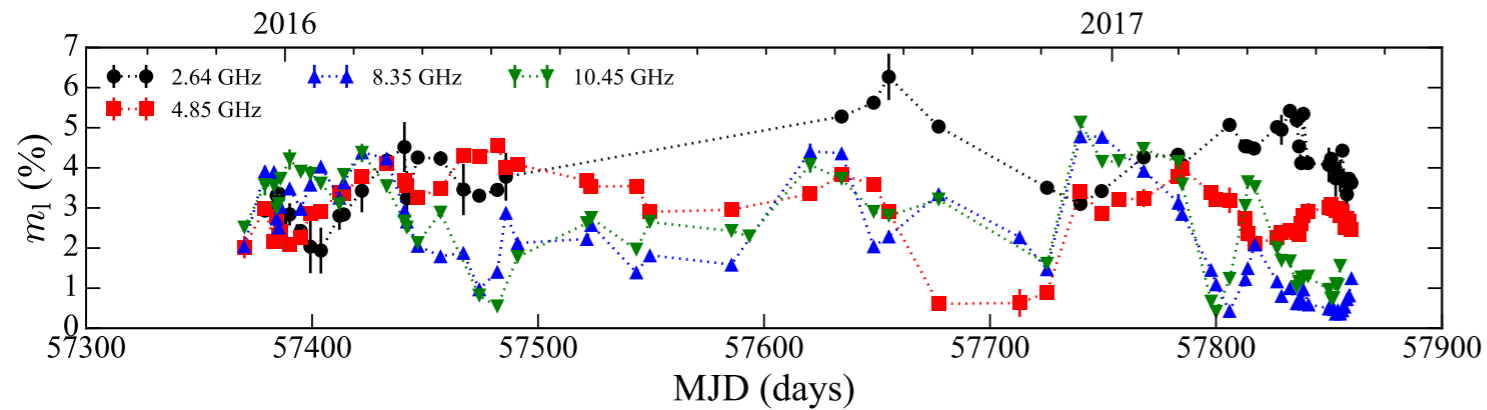






m_c

cadence: 17 days

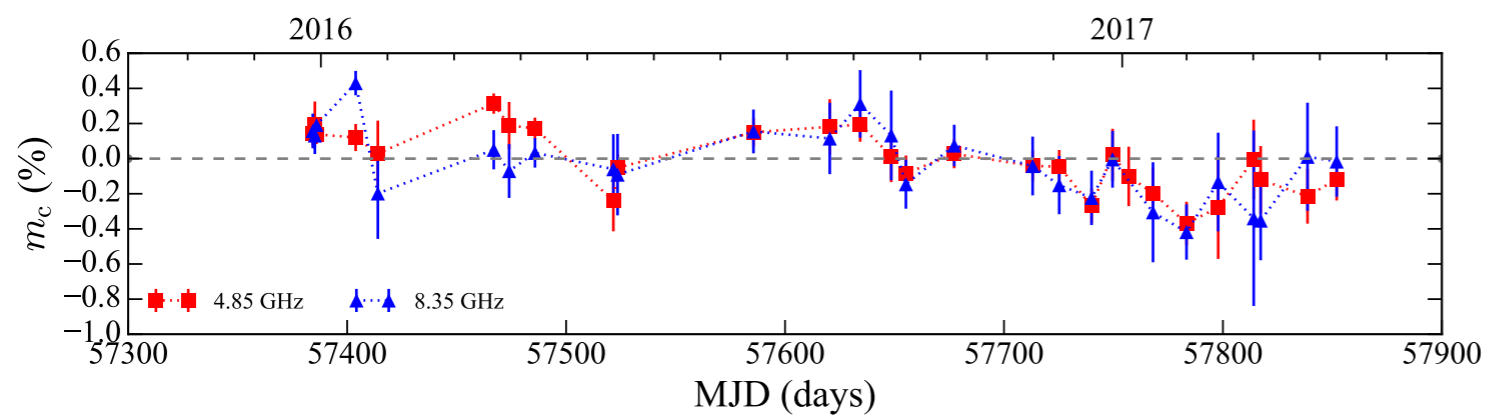
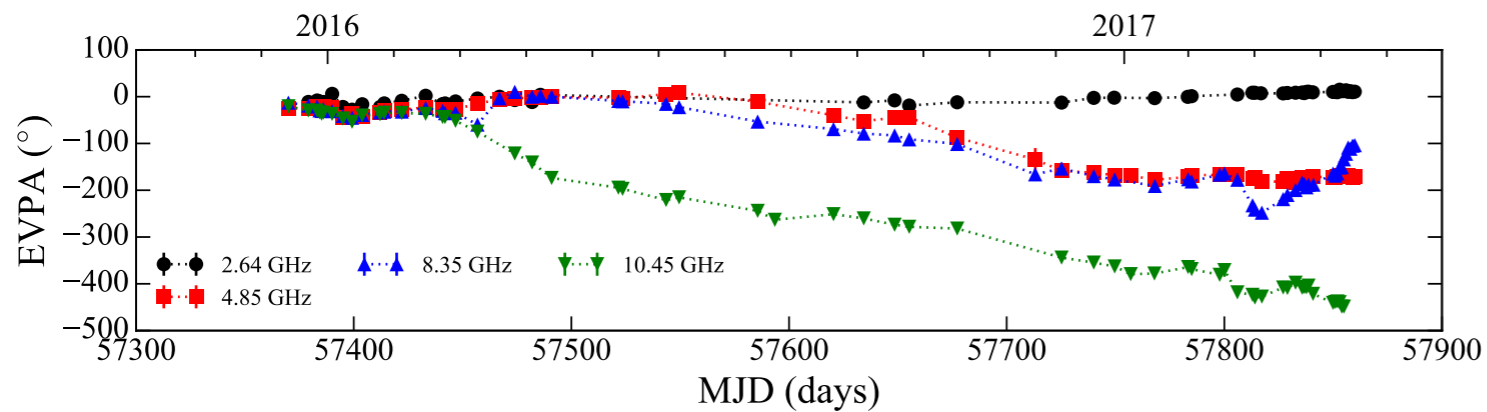


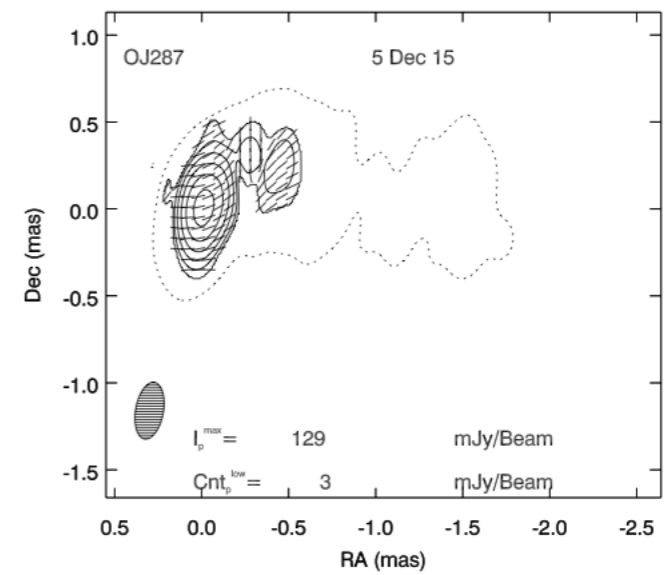
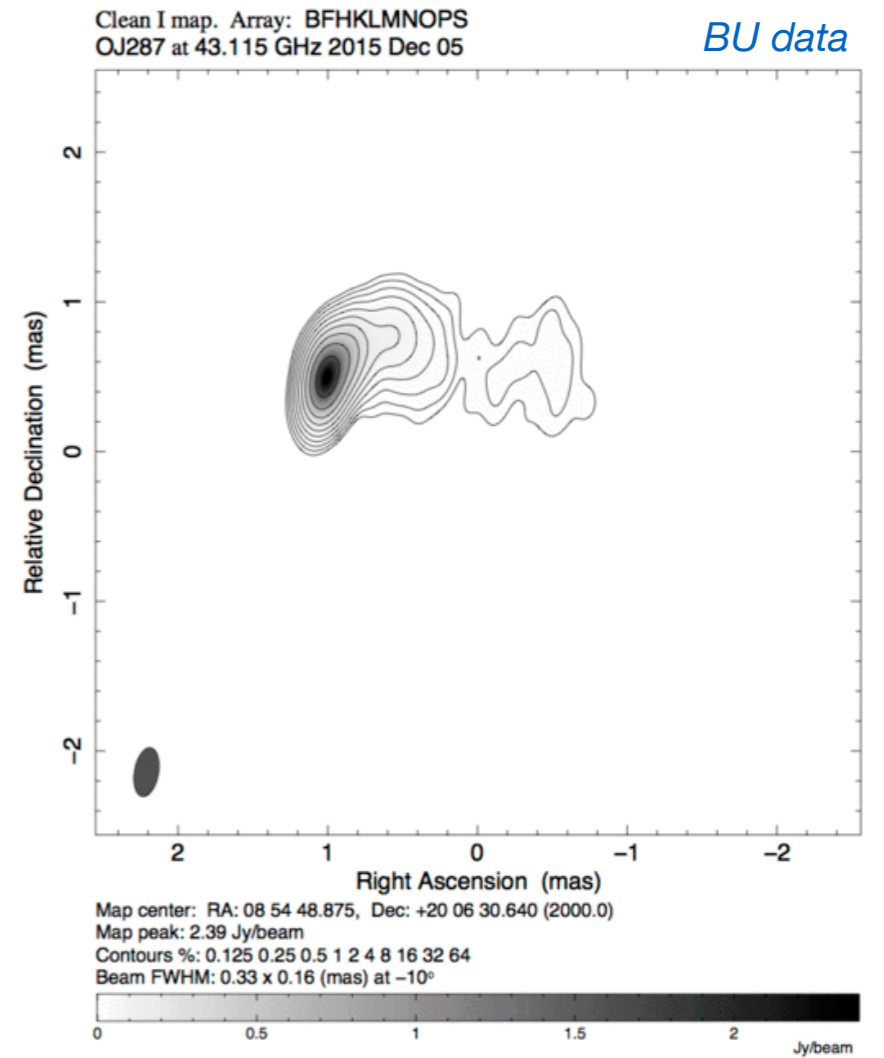
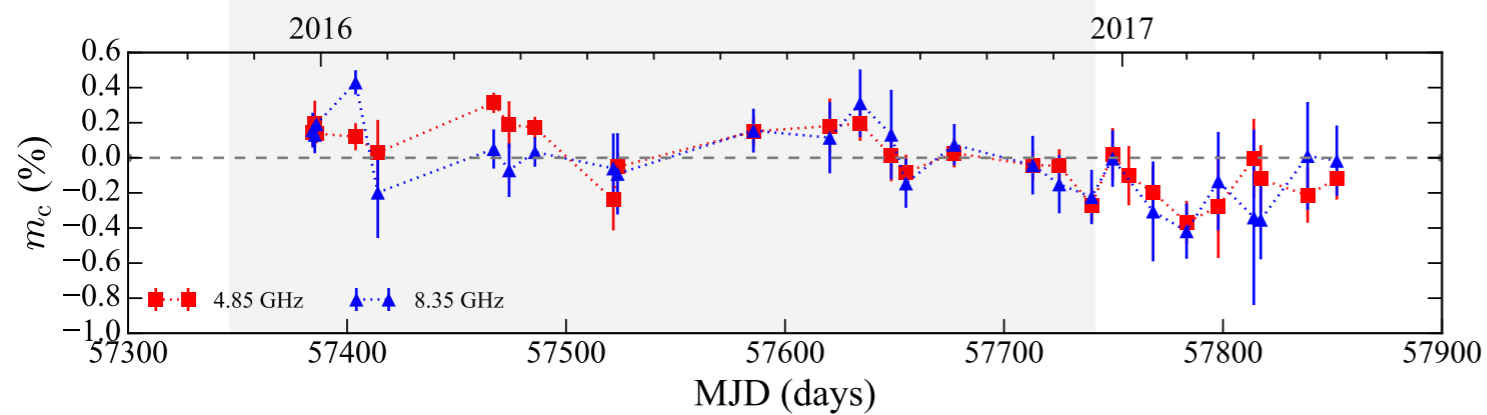
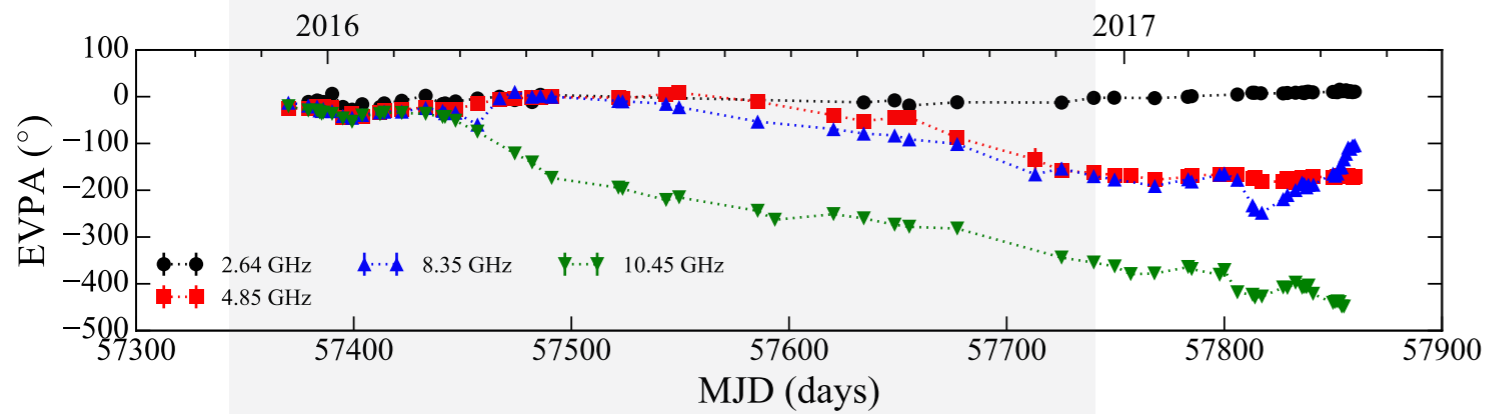
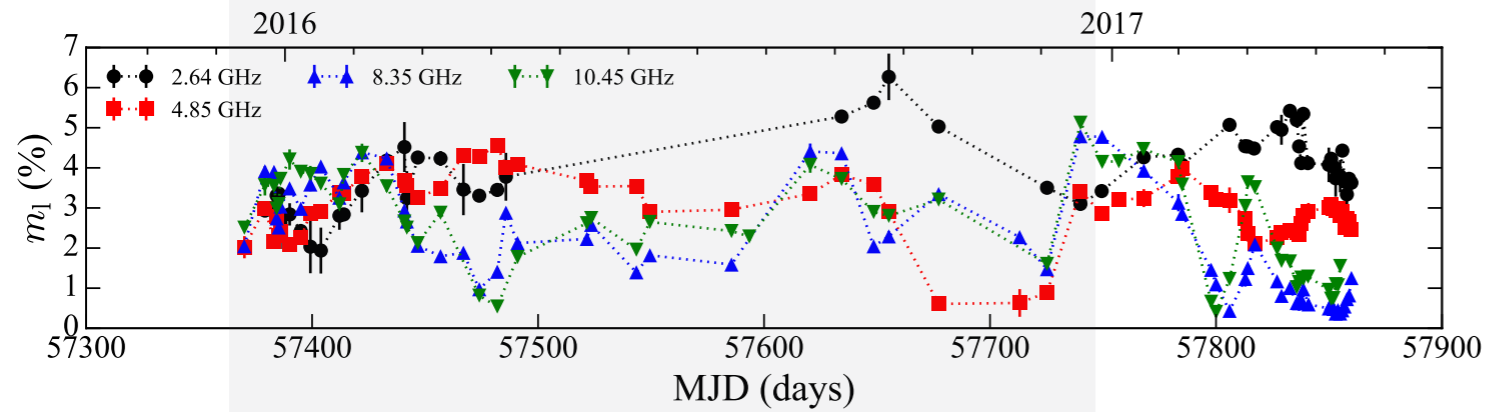
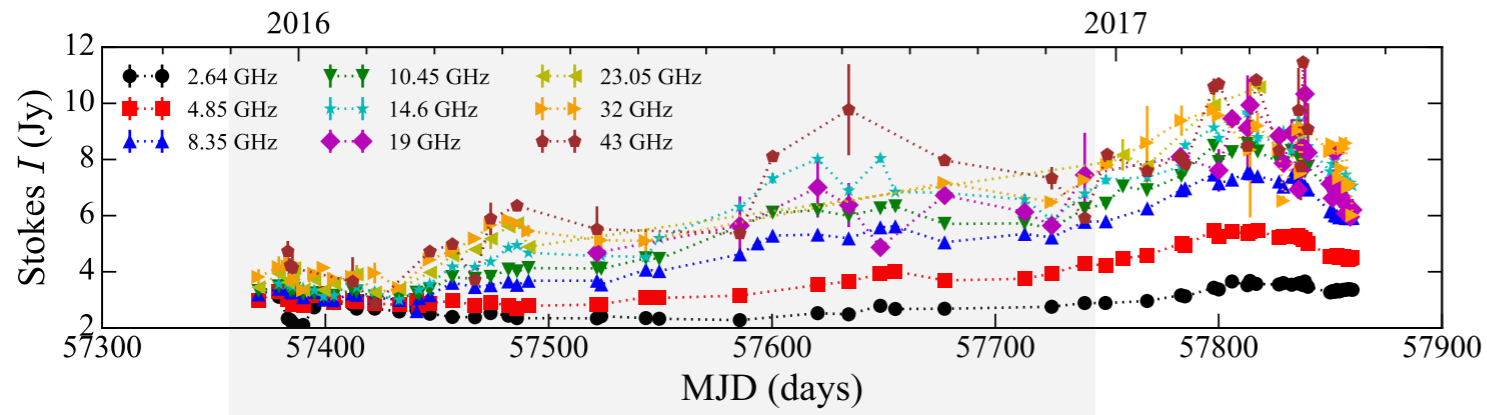
range: -0.4 – 0.3 %
 mean: 0 % (0.2 % $|m_c|$)

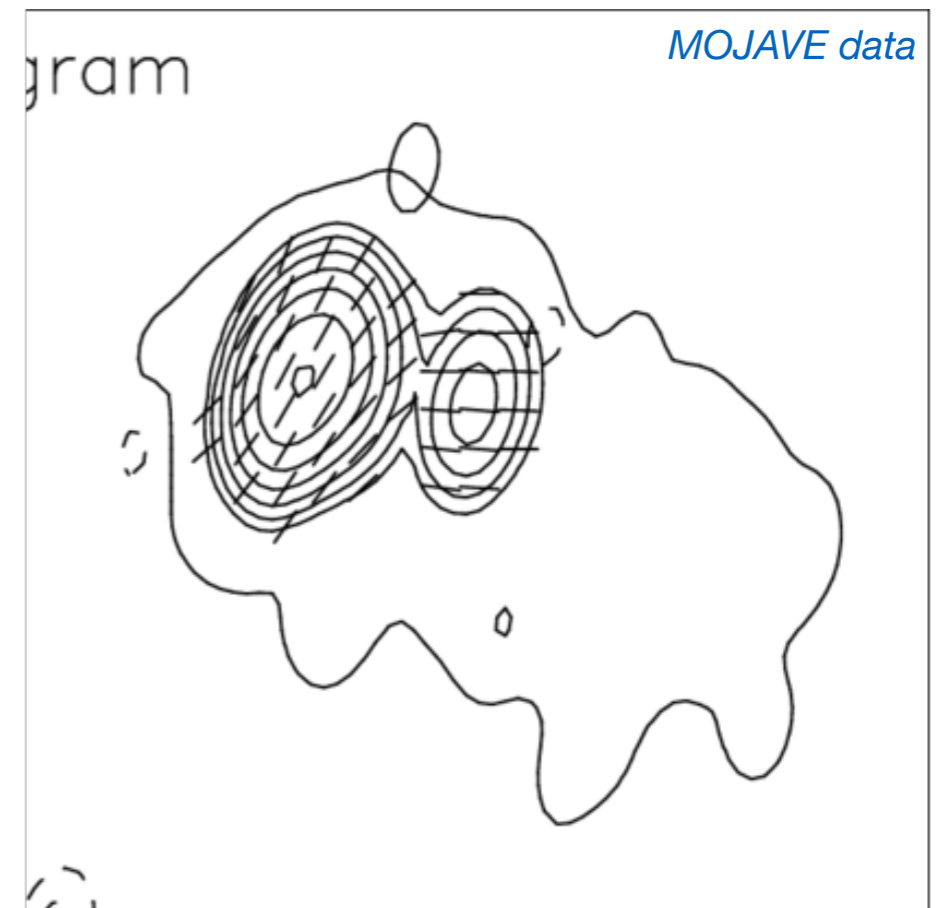
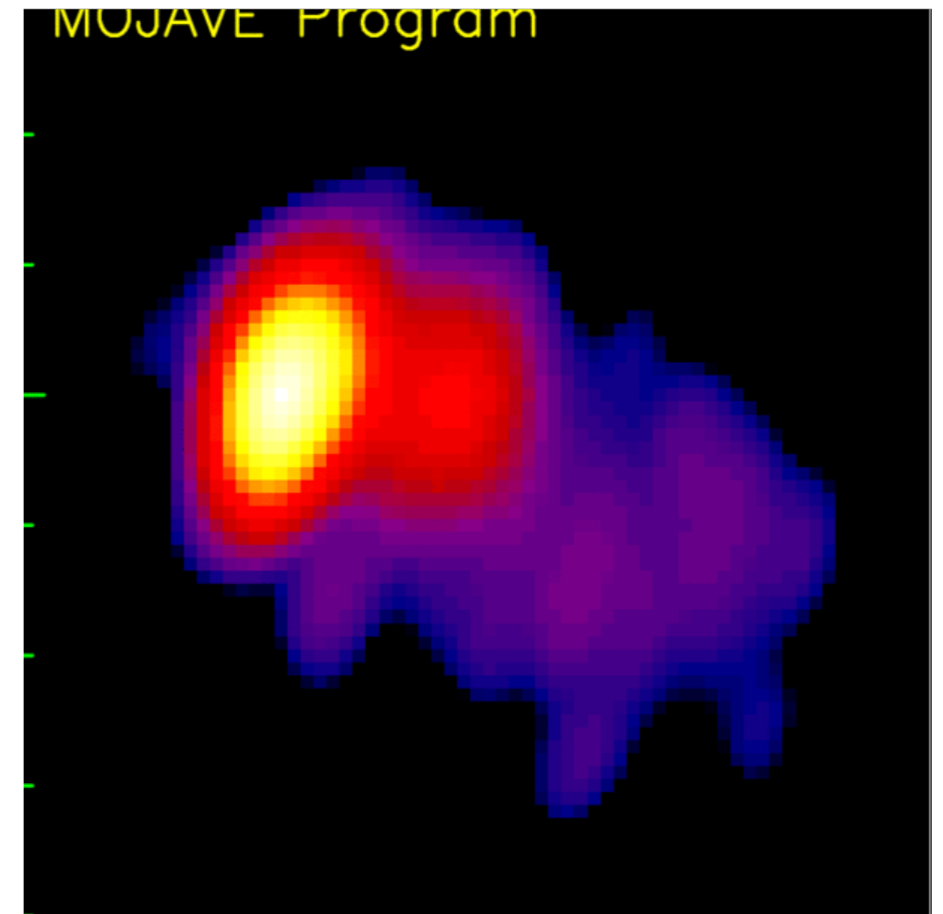
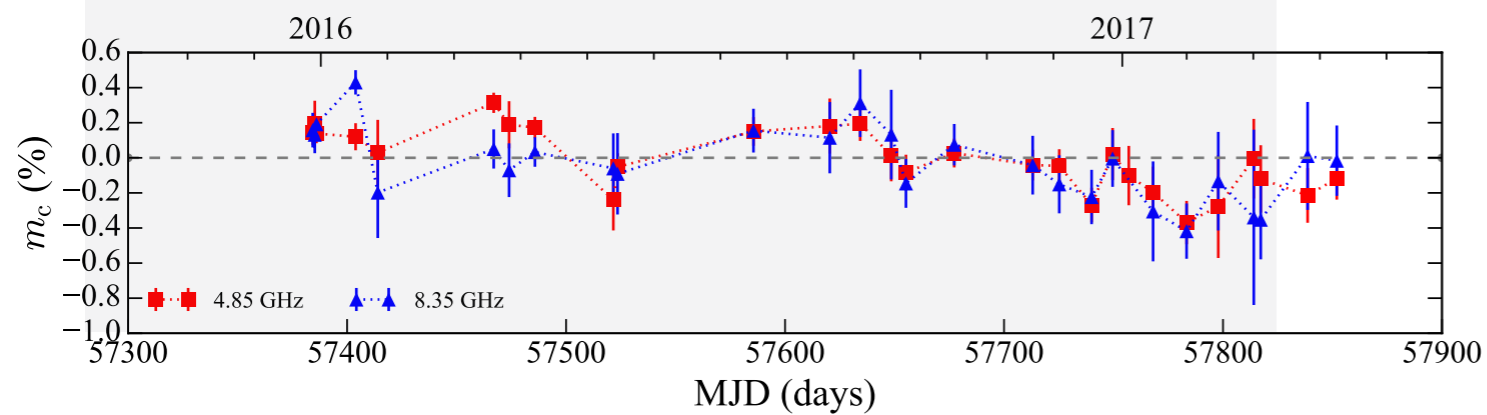
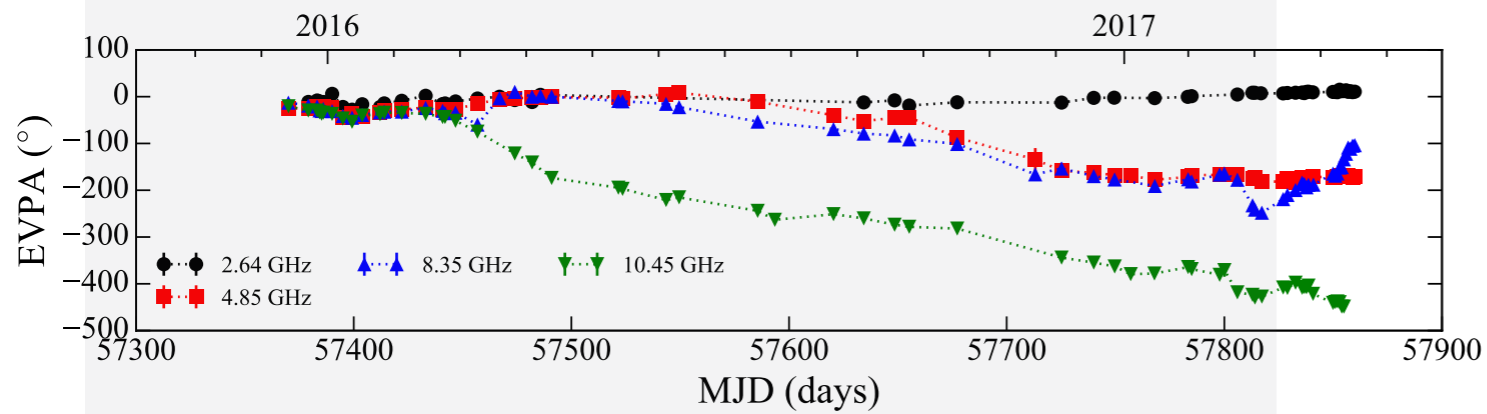
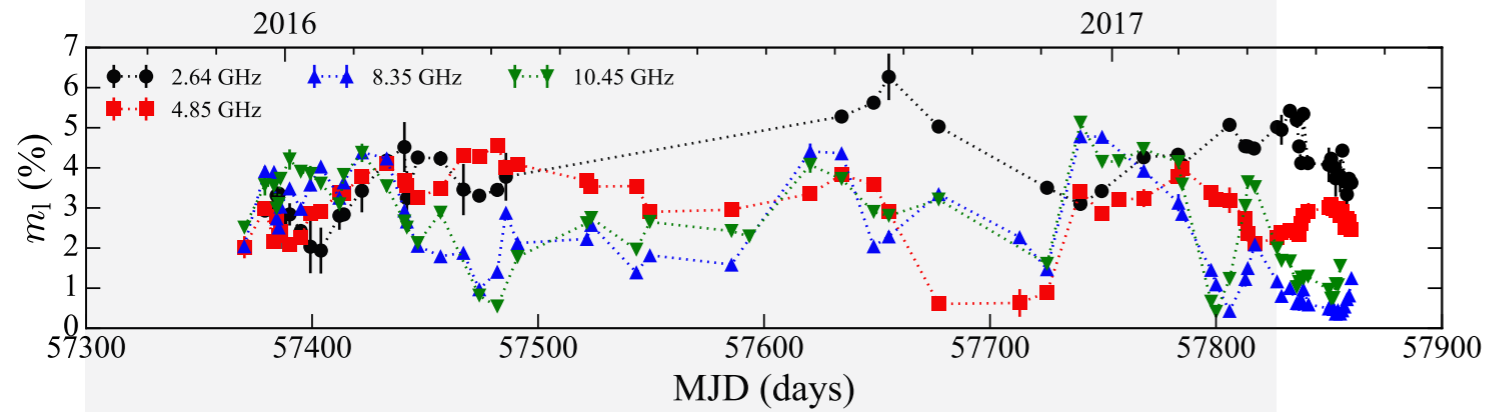
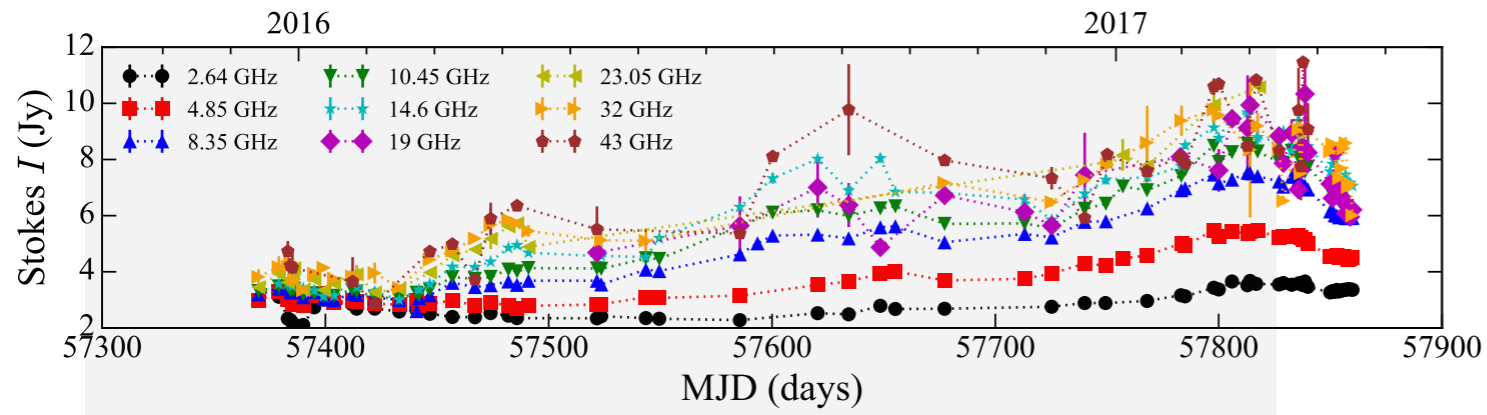
m_c peak at (highest) I max.

- 0.4 %

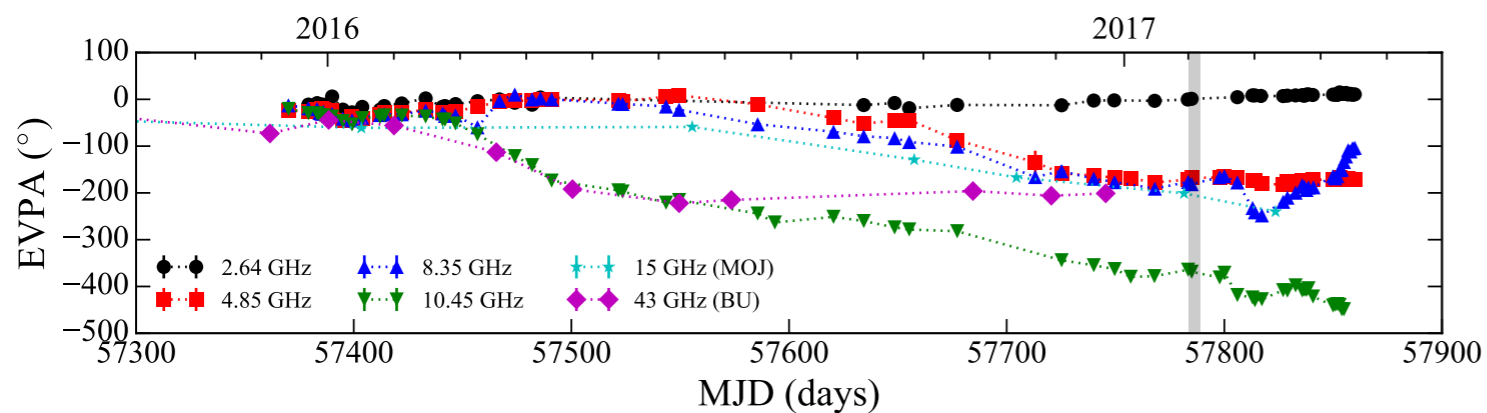
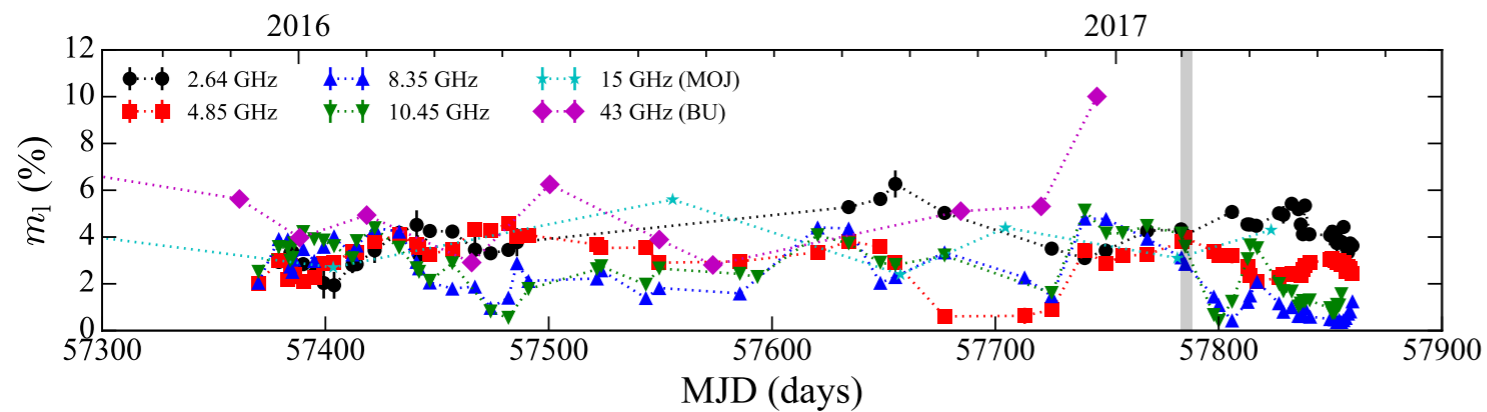
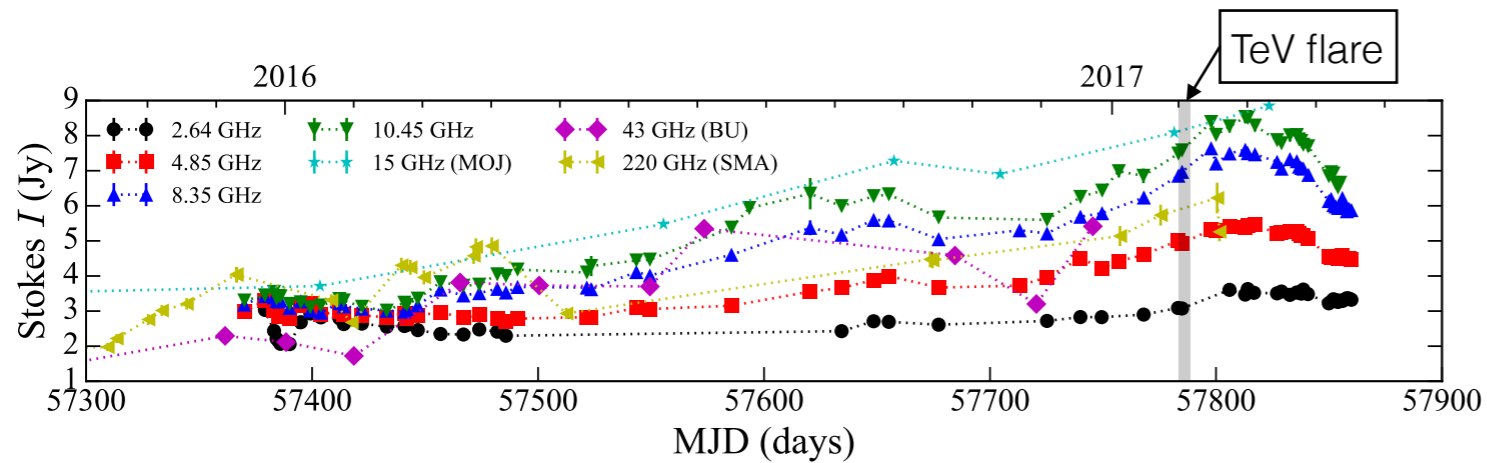
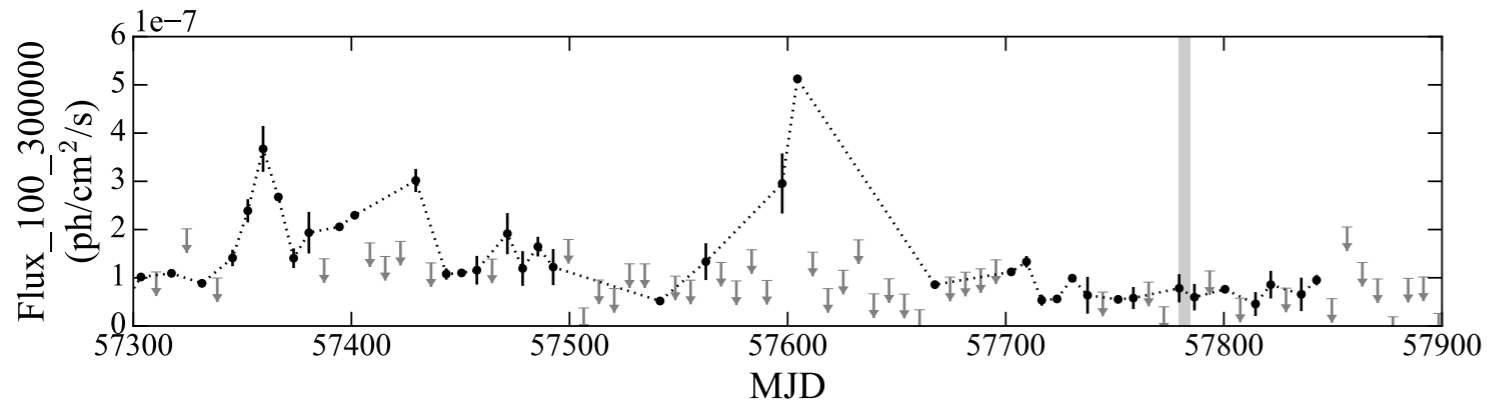
- LH







Other bands



220 GHz (1mm) from SMA

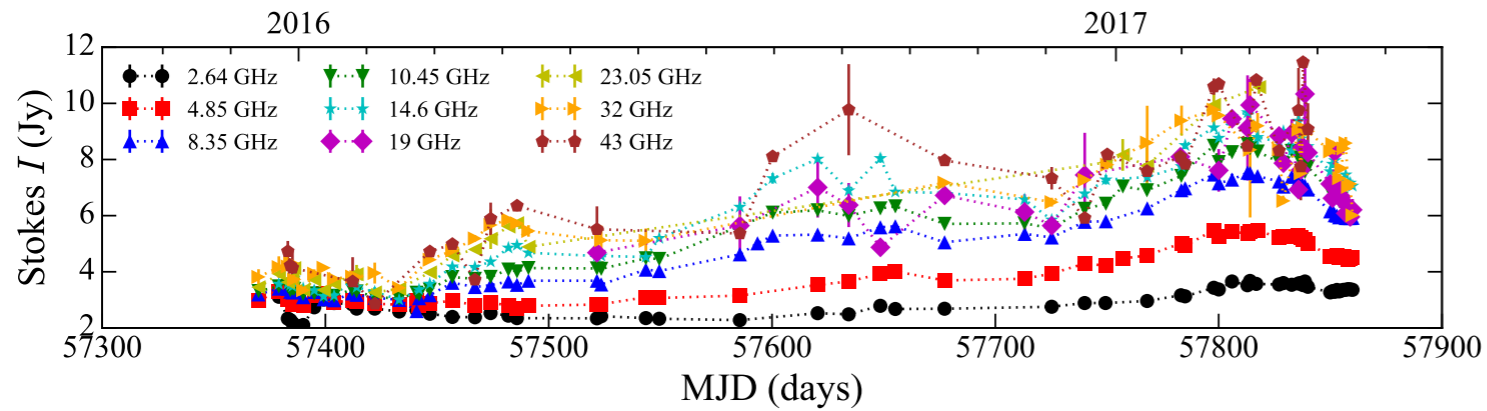
- similar evolution
- lower flux density
- steep spectrum at mm

Fermi

- flare at MJD ~57600, simultaneous with a local I maximum
- no flare at (highest) I maximum

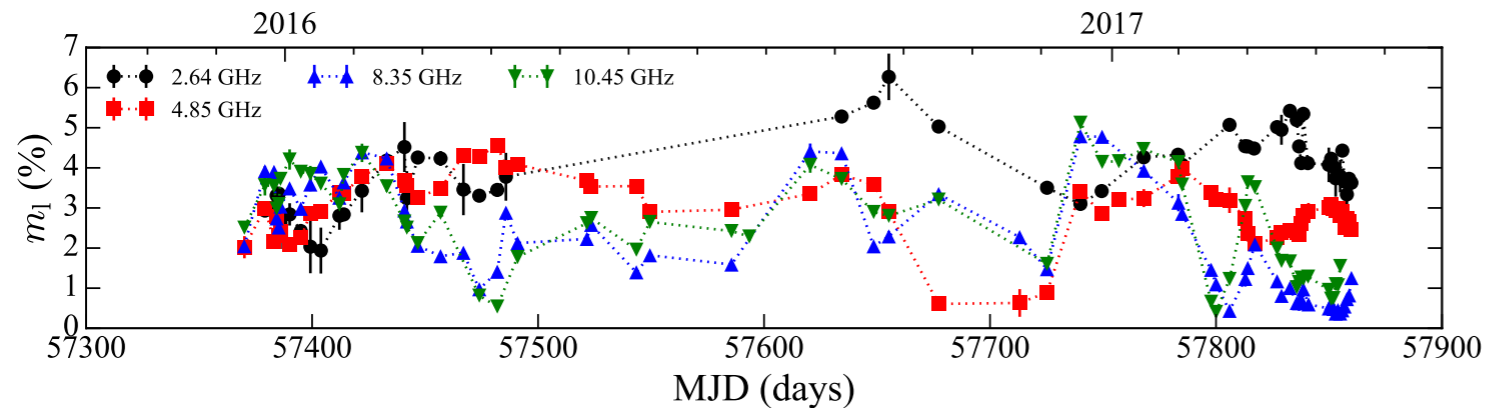
TeV

- flare at (highest) I maximum
- first detection



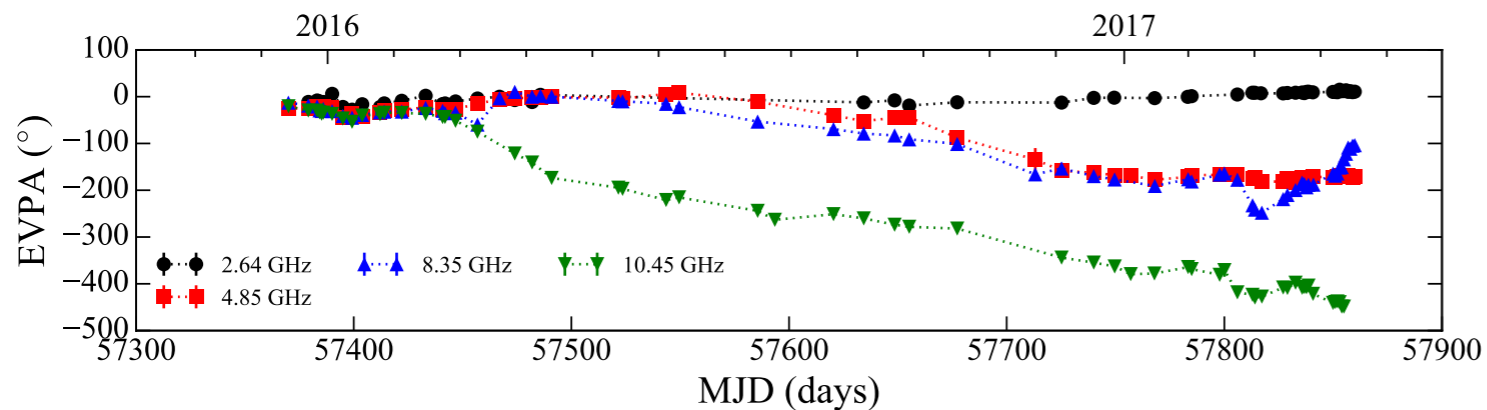
Interpretation (?)

Stokes I increase with flare at MJD ~ 57800



simultaneous LP and I peak(s)

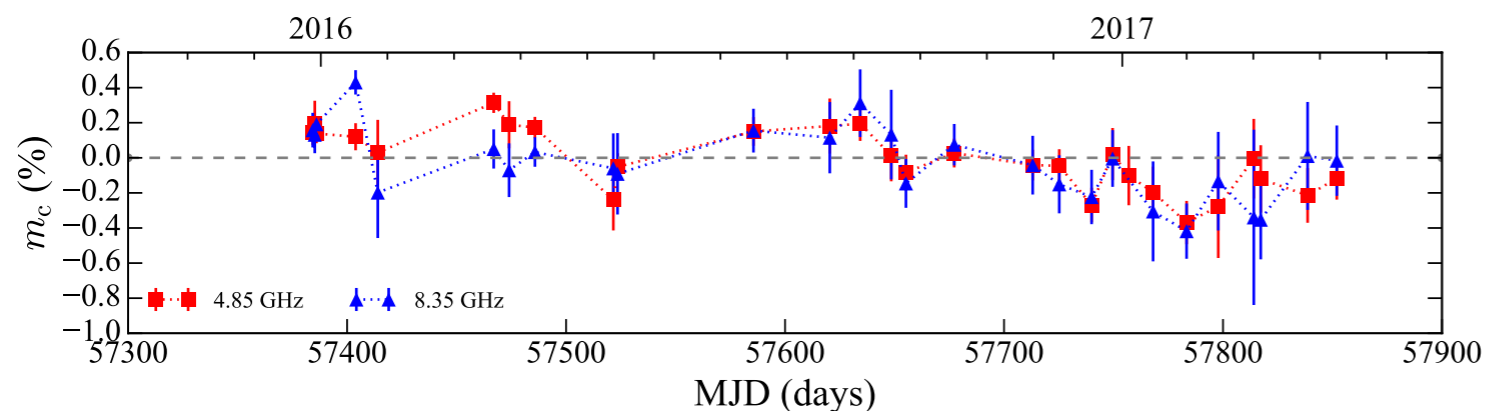
extremely long EVPA rotation



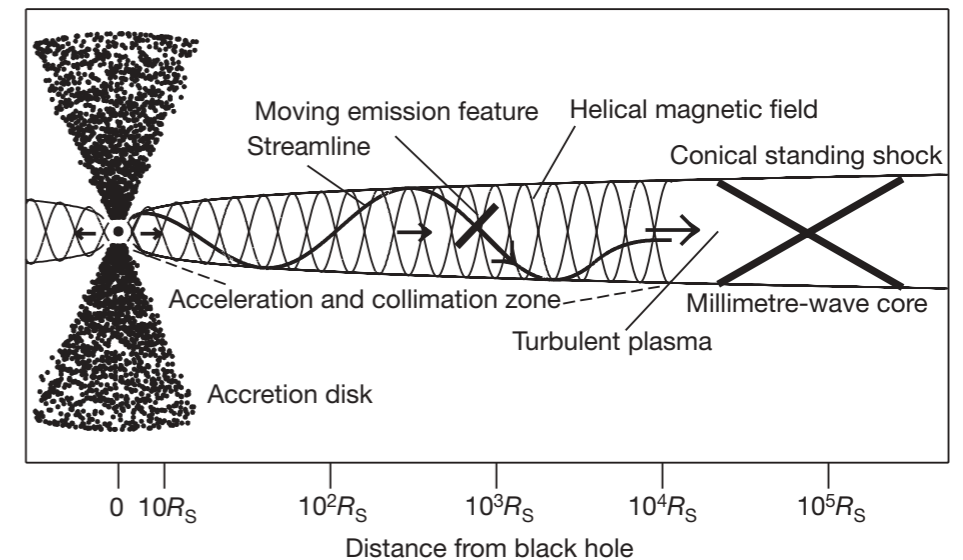
simultaneous CP and I peak

emerging component (?)

EVPA rotation of VLBI core



Marscher et al. 2008, Nature, 452, 966



Summary

High-cadence, full-Stokes monitoring with Effelsberg

- Dec 2015 – now
- mean cadence: 8 days
- LP at 4 frequencies: 2.64, 4.85, 8.35 and 10.45 GHz
- CP at 2 frequencies: 4.85 and 8.35 GHz

Stokes *I* flare in mid-Feb to mid-Mar 2017

Simultaneous LP and *I* peaks

opacity transition signature around MJD 57800

Long (up to 430°), CCW EVPA rotation

CP peak at Stokes *I* flare

BU & MOJAVE: new core component and EVPA rotation

Interpretation

- tracing the helical path or helical B-field component within the core
- opacity transition when the component emerges from the core