



The blazar CTA 102 behaviour during two giant outbursts

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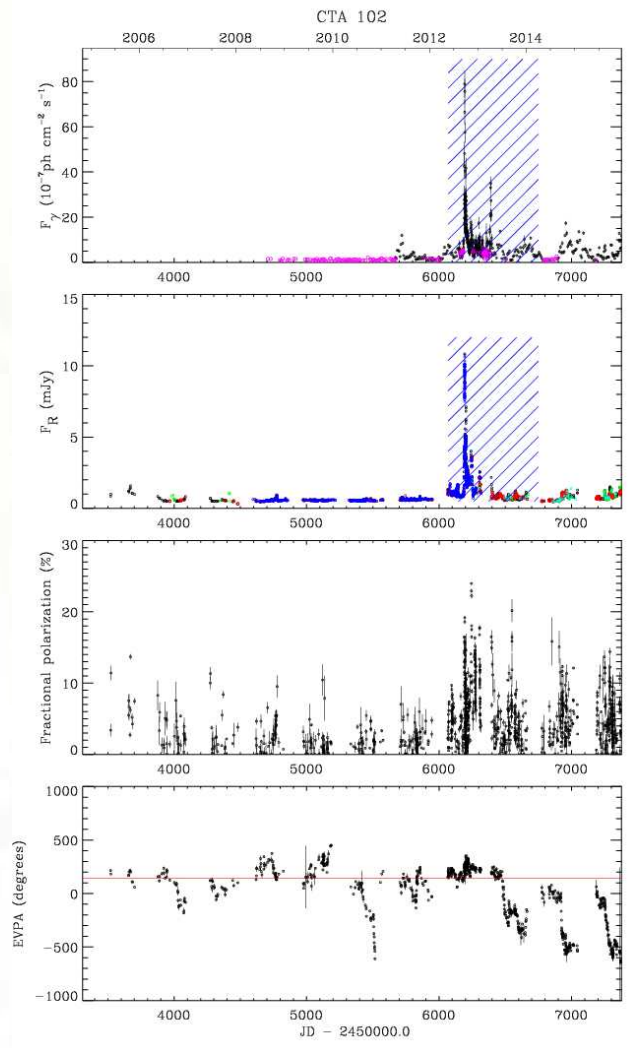
Svetlana Jorstad, Alan Marscher, Paul Smith, Iván Agudo, Daria Morozova, Sergey Savchenko, Tatiana Grishina, Evgenia Kopatskaya, Liudmila Larionova, Elena Larionova, Anna Mokrushina, Ivan Troitsky, Yulia Troitskaya, George Borman

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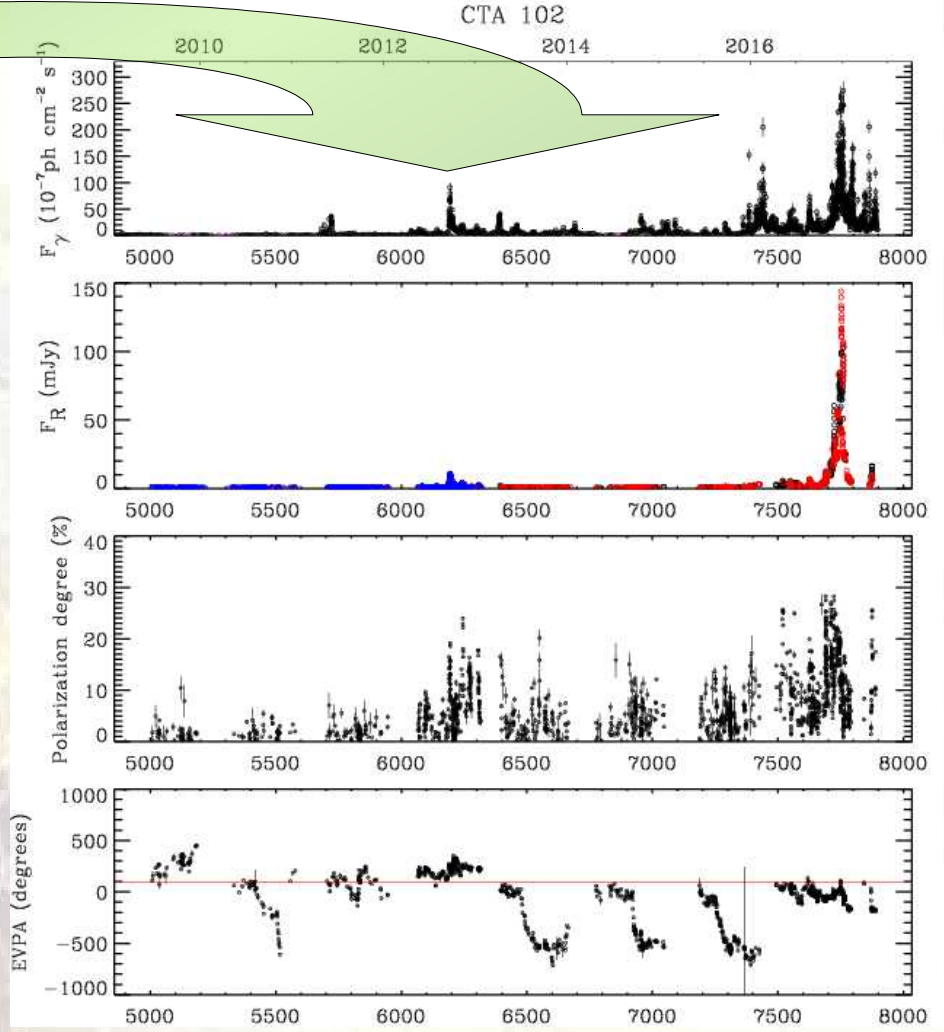
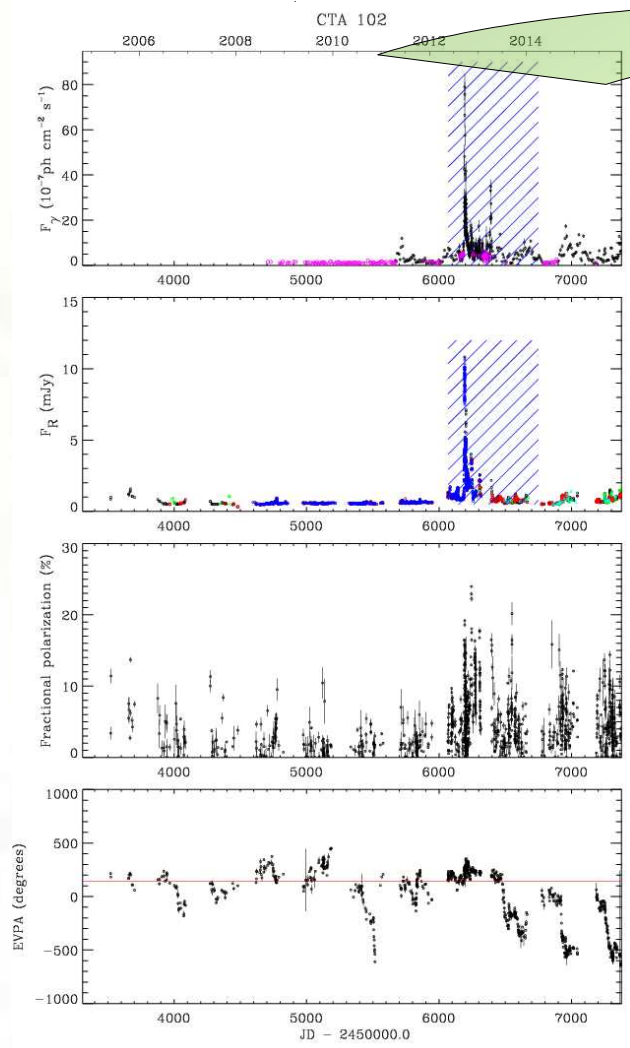
- **The importance of monitoring projects in blazar studies was clear from the very discovery due to their violent variability on all time scales, from hours and even minutes to years.**
- **The outbursts that may occur unexpectedly for any given source act like a magnifying glass that allow to penetrate deeper in most conspicuous regions of inner jet.**
- **Numerous successful campaigns were carried out that lead to systematic accumulation of observational data, either inside one observatory or as large international collaborations.**

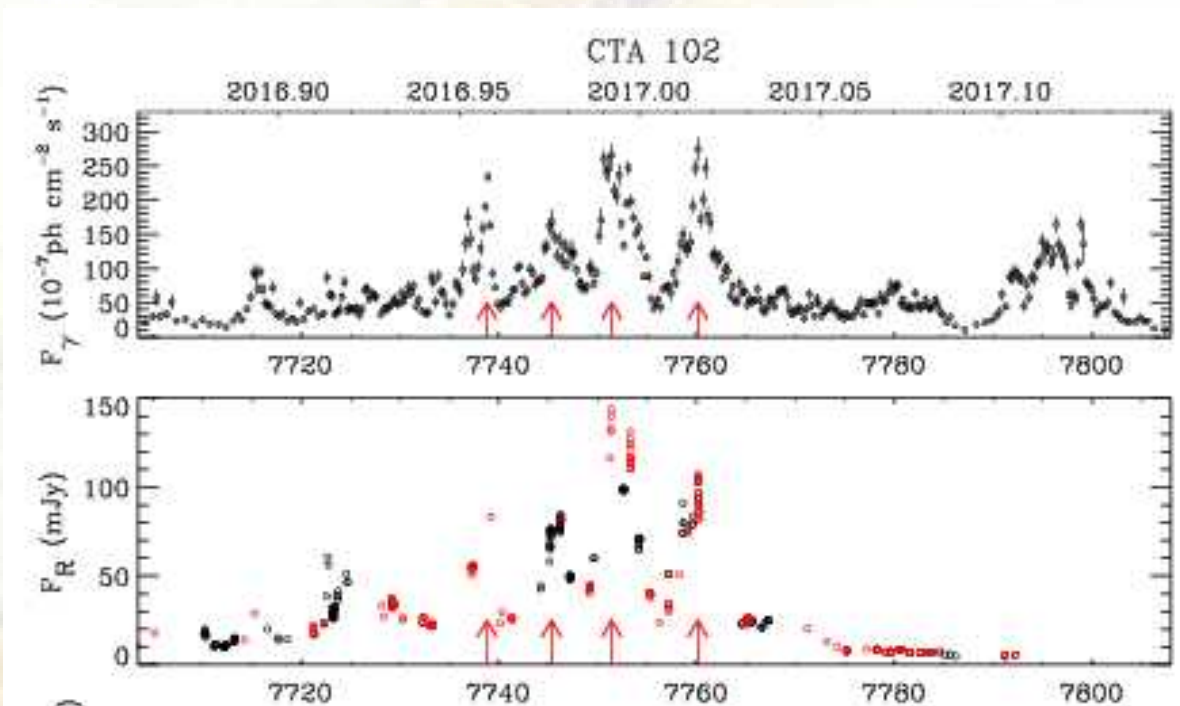
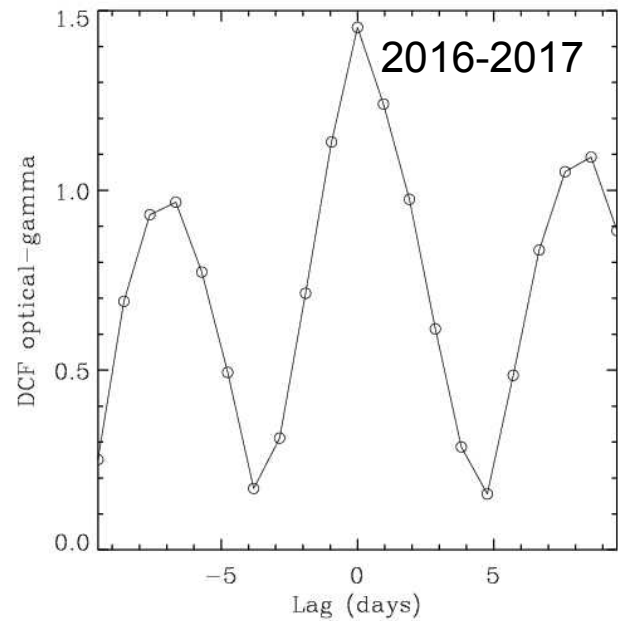
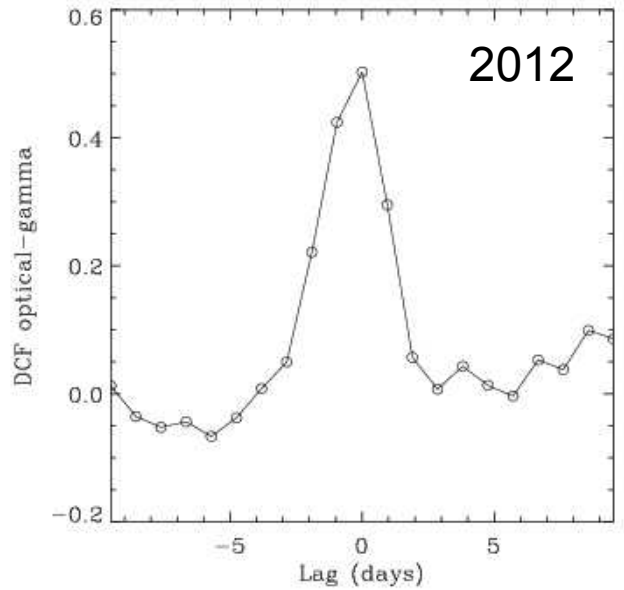
Exceptional outburst of the blazar CTA 102 in 2012: The GASP-WEBT campaign and its extension. *

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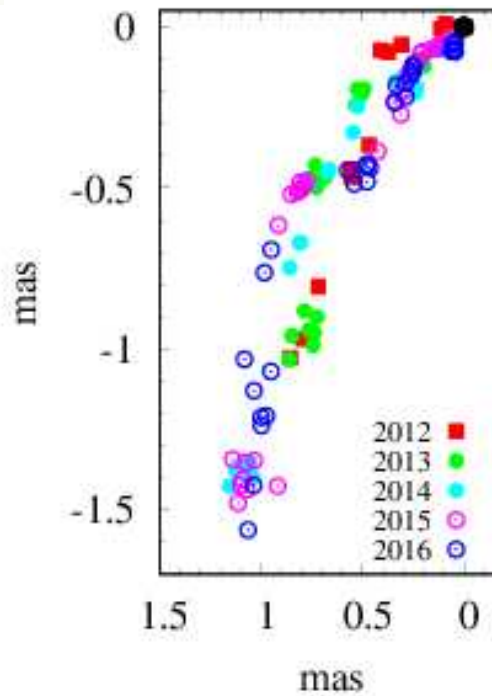
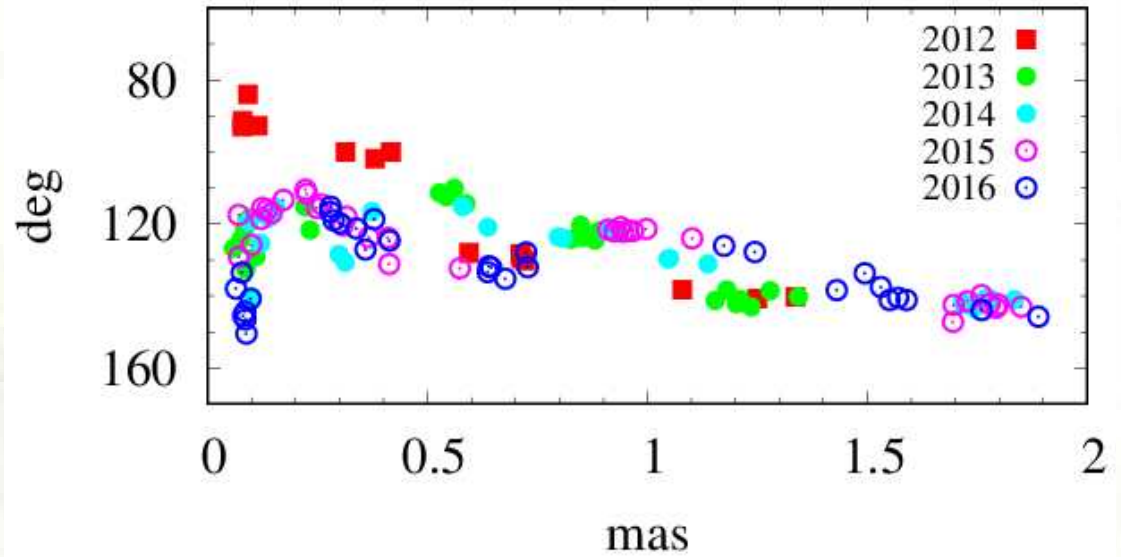
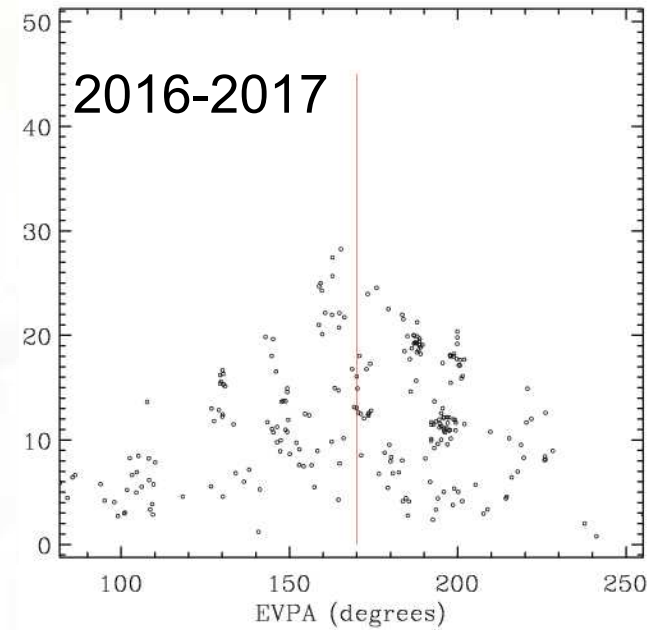
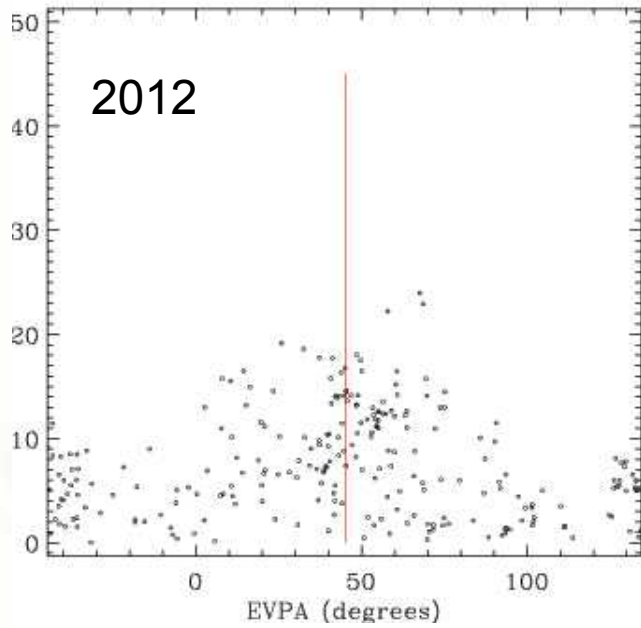


«We can expect that flaring activity in this blazar will be more pronounced than in previous years as long as the jet remains closely aligned with our line of sight.»
(Larionov+ 2016)

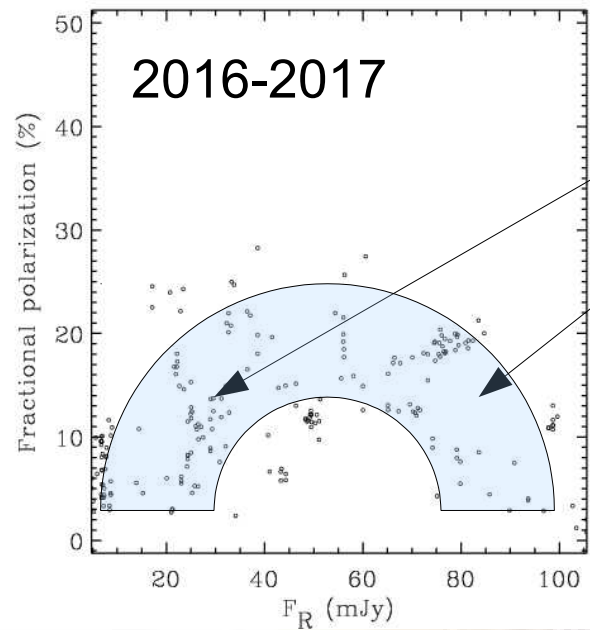
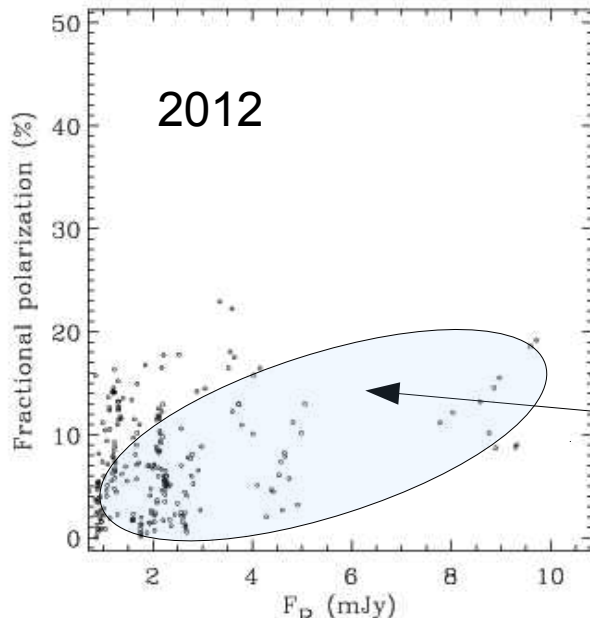




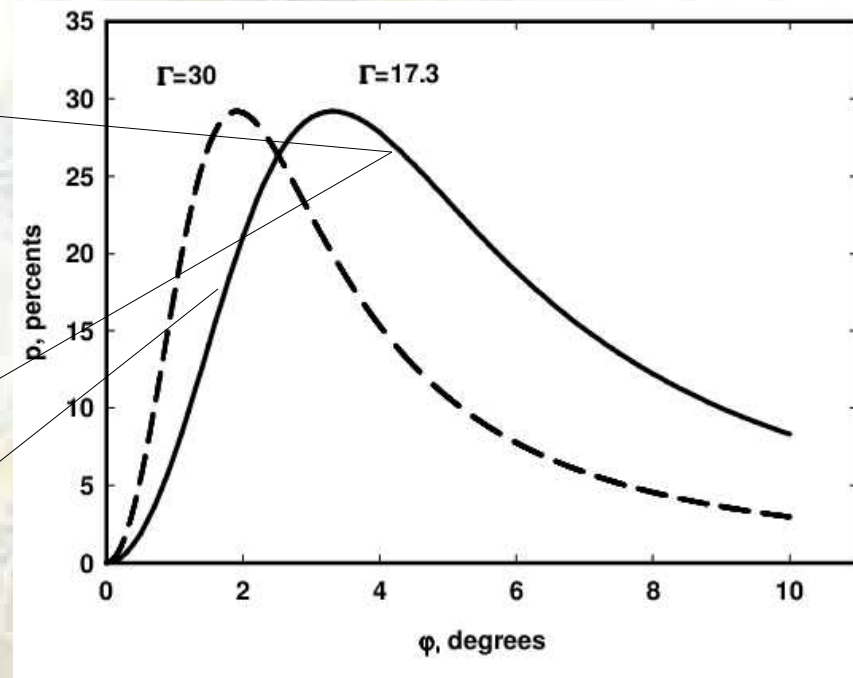
- During both outbursts there are no delays between γ -rays and optical => **the sites of synchrotron and IC radiation are co-spatial.**
- Triple DCF in 2016-2017 is a consequence of **~ 7 days recurrence of optical and γ -ray outbursts.**



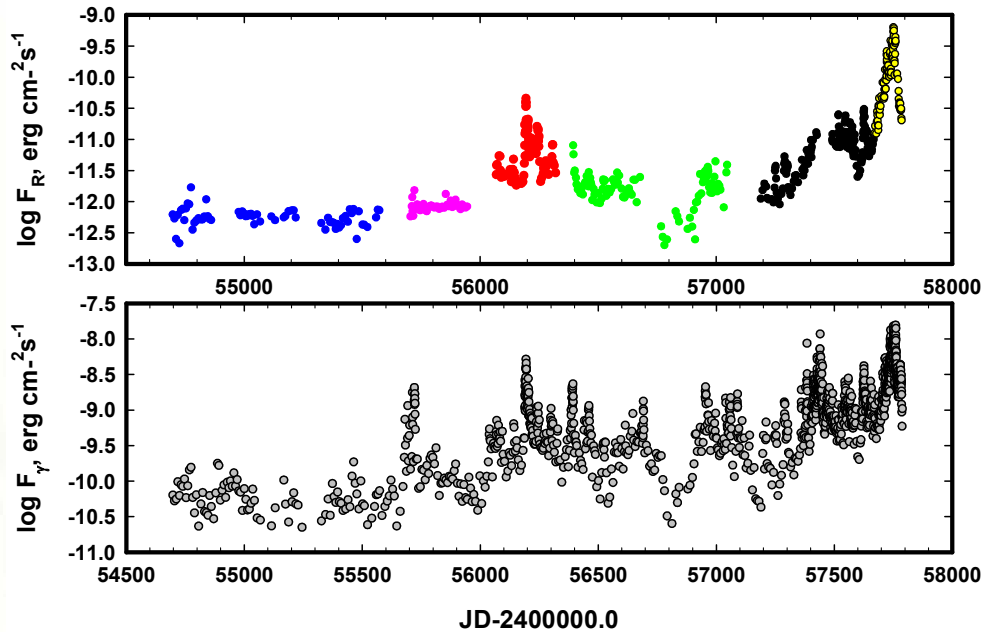
- 2012:
 $EVPA_{opt} - \theta_{43GHz} \sim -45^\circ$
- 2016:
 $EVPA_{opt} - \theta_{43GHz} \sim 30^\circ$



Behaviour of fractional polarization vs. viewing angle for plasma compression ratio $\eta = 1.5$, Lorentz-factor $\Gamma = 17.3$ (solid line), and $\Gamma = 30$ (dashed line), in the moving shock model. (Figure 12 in Larionov+2016).

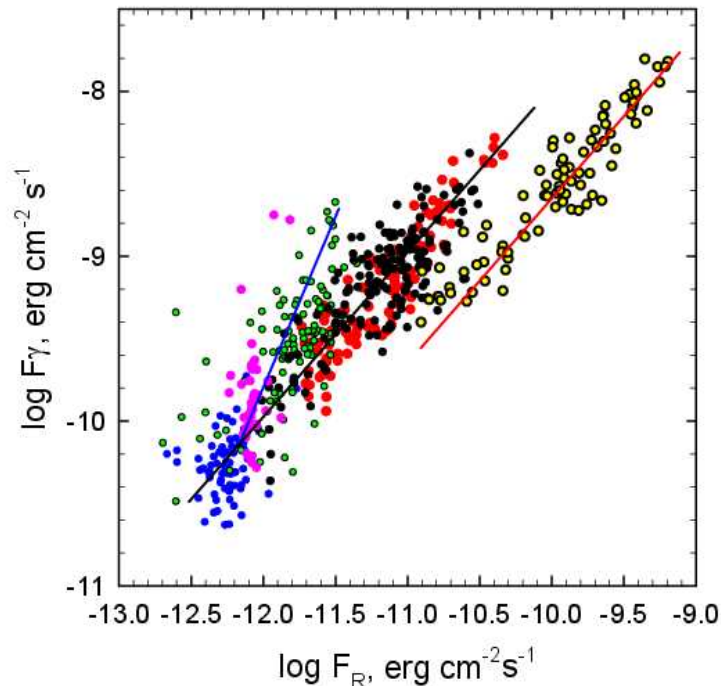


The dependence of PD vs. F_R during the 2016-2017 outburst reflects either smaller viewing angle or higher Lorentz-factor (or both)



An unexpected behaviour of dependence of F_γ vs F_{optical}

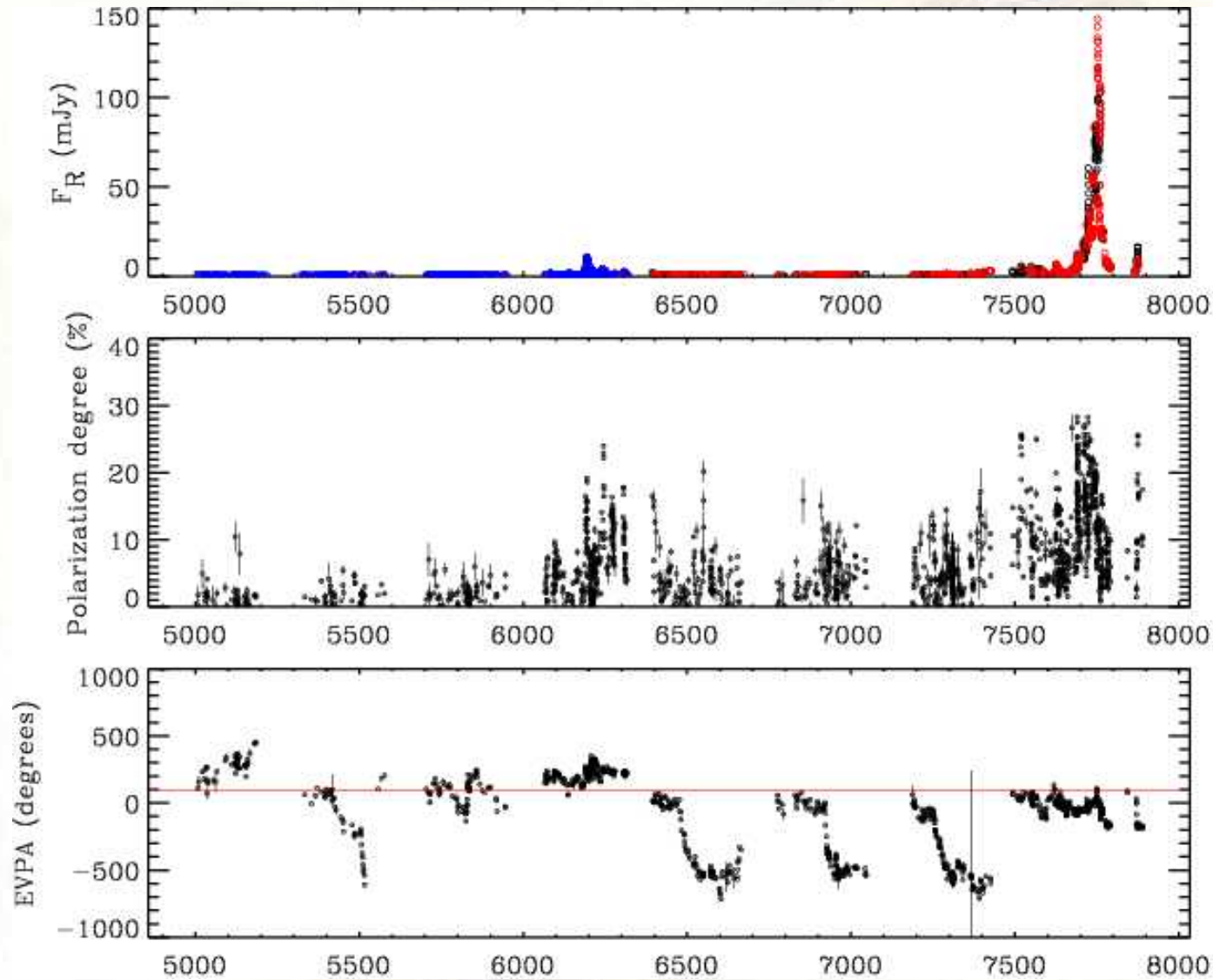
The 2016 outburst points location is distinctly different from 2012.



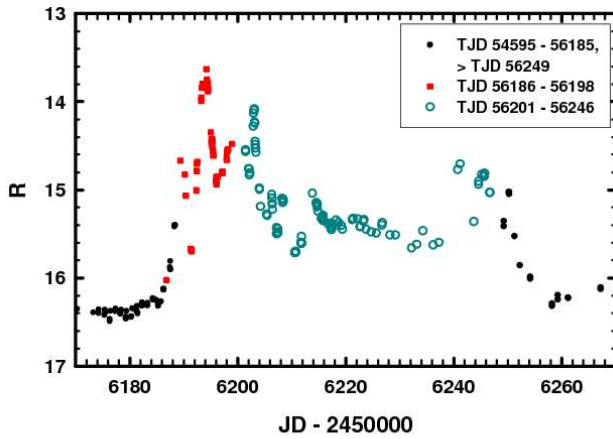
Does it mean that 2016 outburst is located downstream the jet, ~ 2 times farther than 2012 (supposing that

$$F(\text{IC}) \sim n_e * F_{\text{synchr}} \text{ and } n_e \sim r^{-2}?)$$

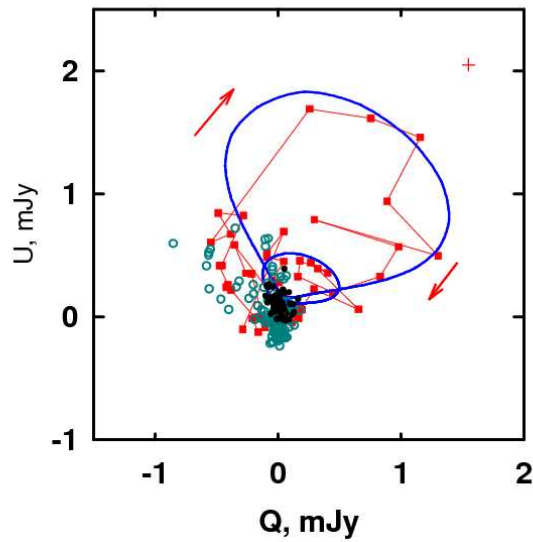
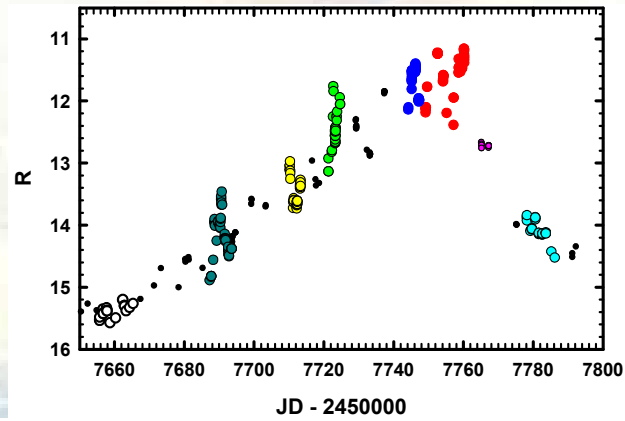
Rotations ?



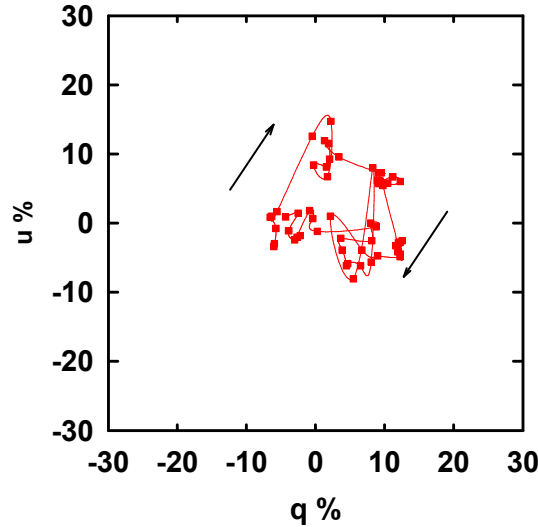
2012



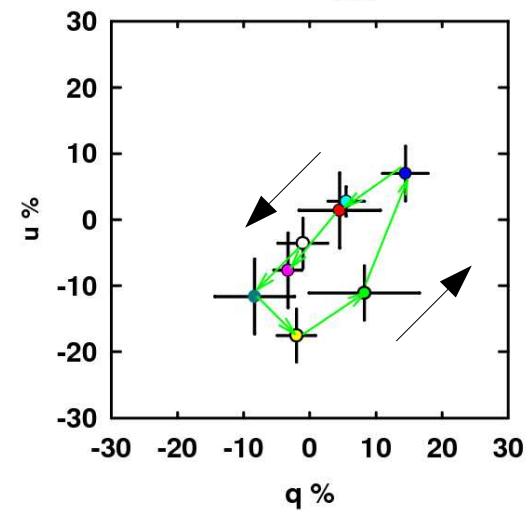
2016-2017



clockwise



clockwise



counterclockwise

Summary

- During both outbursts there are no delays between γ -rays and optical => **the sites of synchrotron and IC radiation are co-spatial**
- The dependence of PD vs. F_R during the 2016-2017 outburst reflects **either smaller viewing angle or higher Lorentz-factor (or both)**
- **EVPA behavior is substantially different between the 2012 & 2016-17 outbursts**
- **The shift of $\log F_{\text{opt}} - \log F_{\gamma}$ dependence** may be caused by the location of the 2016-17 outburst being 2 times farther downstream from the BH than during 2012 event. => Consistent with theoretical expectation that higher Lorentz factors are achieved farther from BH when ratio of magnetic to rest-mass energy at jet base increases
- As in the 2012 case, the 2016-17 outburst shows **clockwise rotation near the peak of the optical light curve on a time scale of several days**
- **On a time scale of months we notice counter-clockwise rotation** that might be a sign of two helices twisted in opposite directions

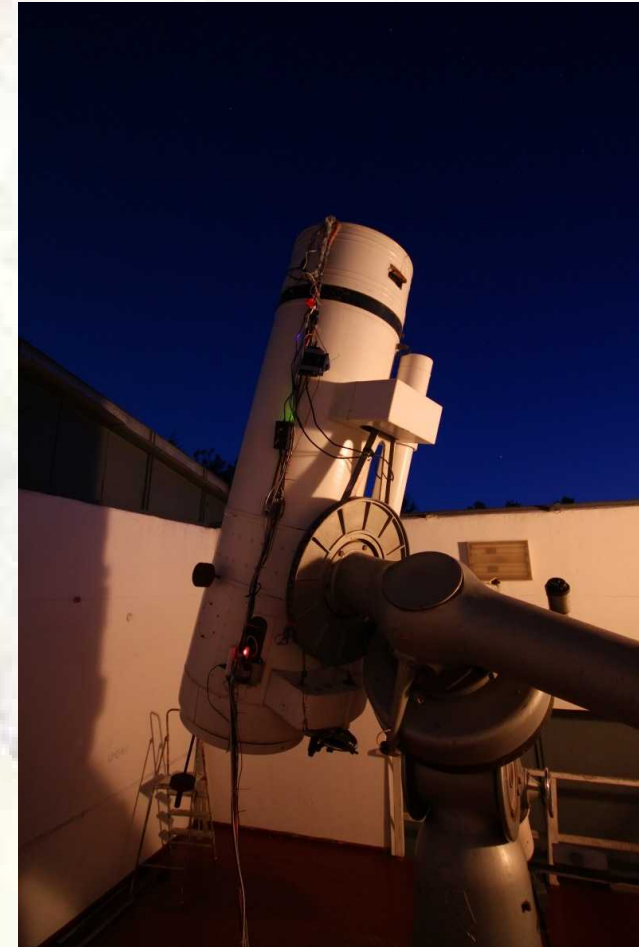
Perkins
Arizona (Lowell)



LX-200
St.Petersburg



AZT-8
Crimea





Ευχαριστώ!

Спасибо за внимание !

Thank you !