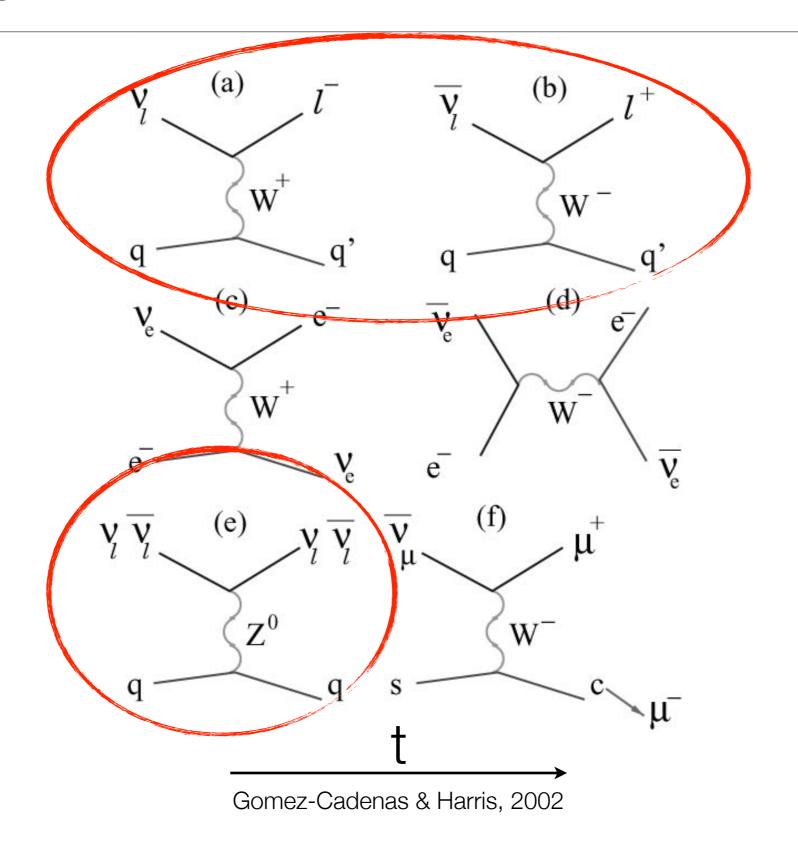
The IceCube Neutrino Observatory

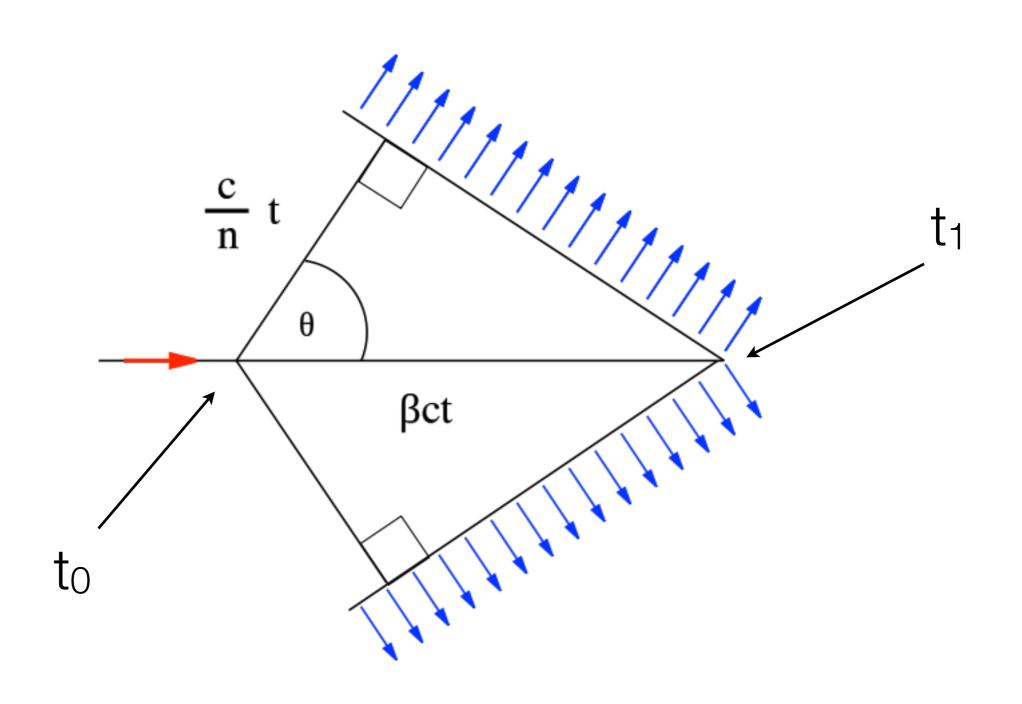
IMPRS Retreat 2011, Hamburg

Lars Flöer

Neutrinos produce secondary particles by scattering with a nucleus

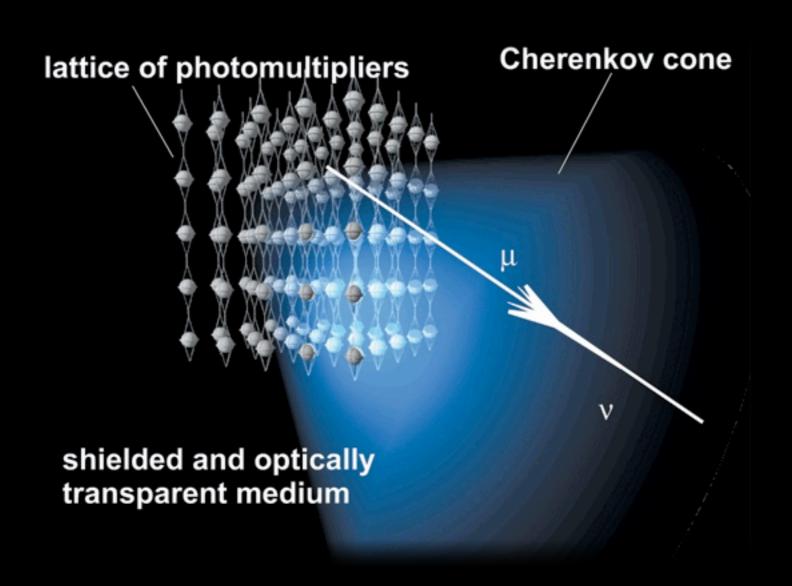


Particles emit Cherenkov radiation if they exceed the speed of light in the ambient medium



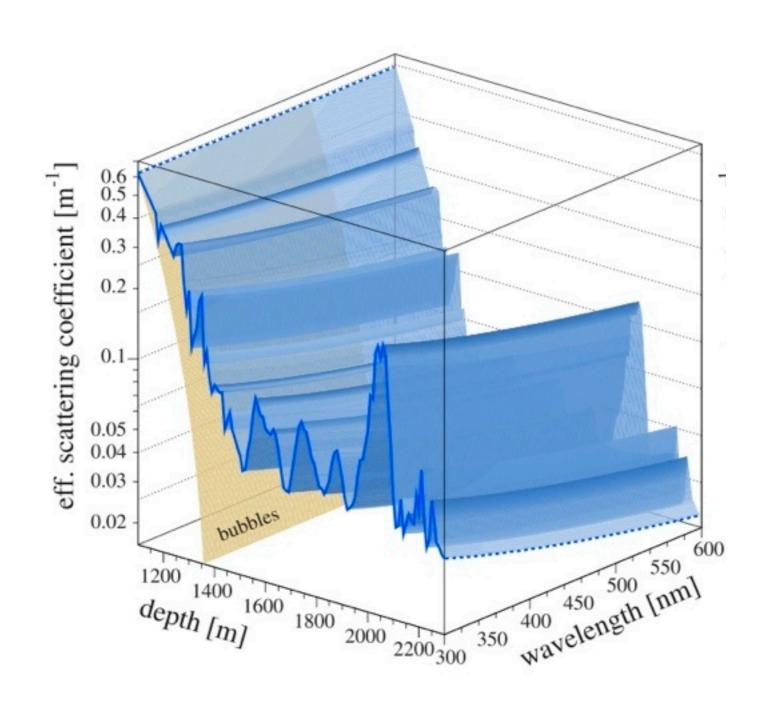


Neutrinos are detected by observation of Cherenkov radiation from secondary leptons

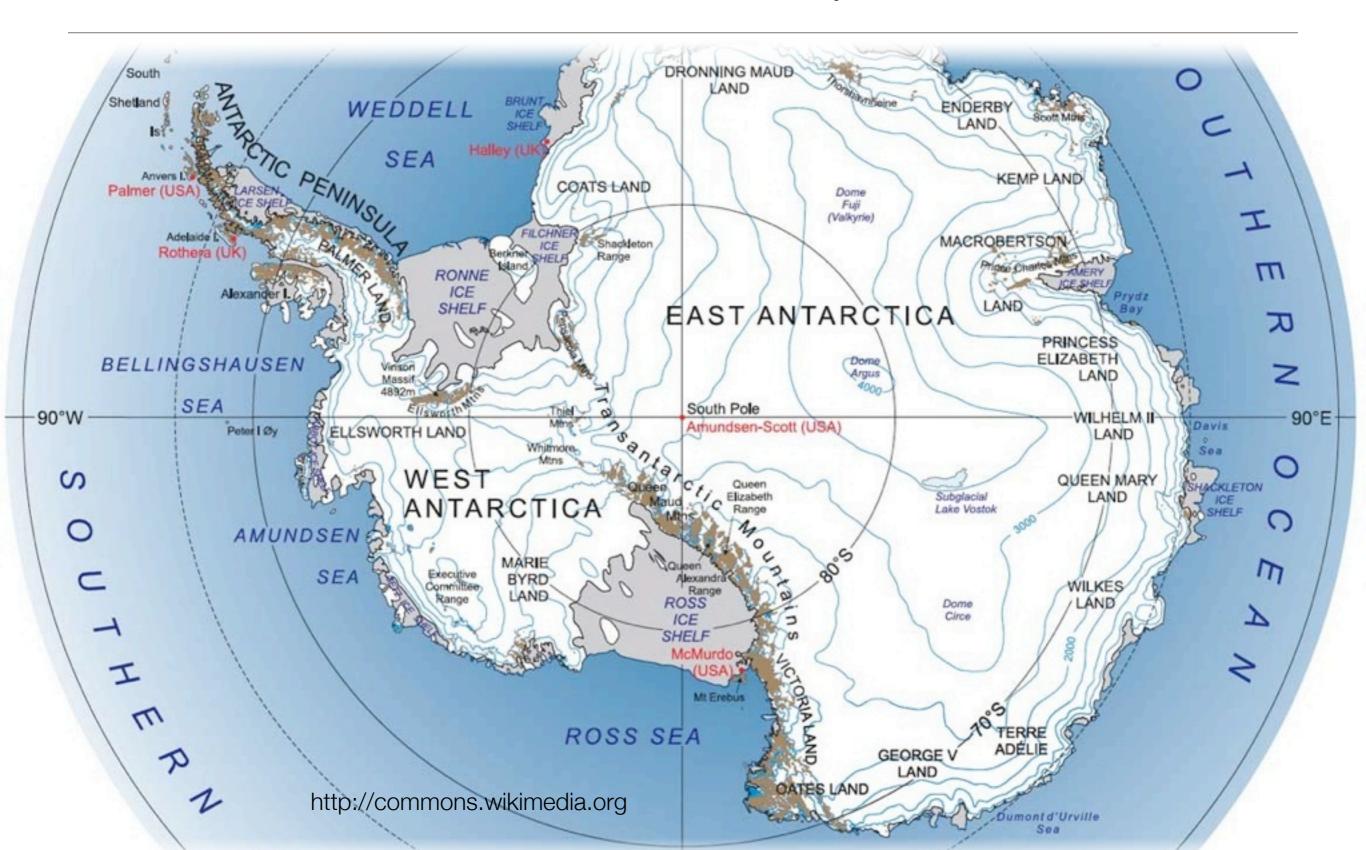


Francis Halzen, "IceCube neutrino observatory," in AccessScience, ©McGraw-Hill Companies, 2011, http://www.accessscience.com

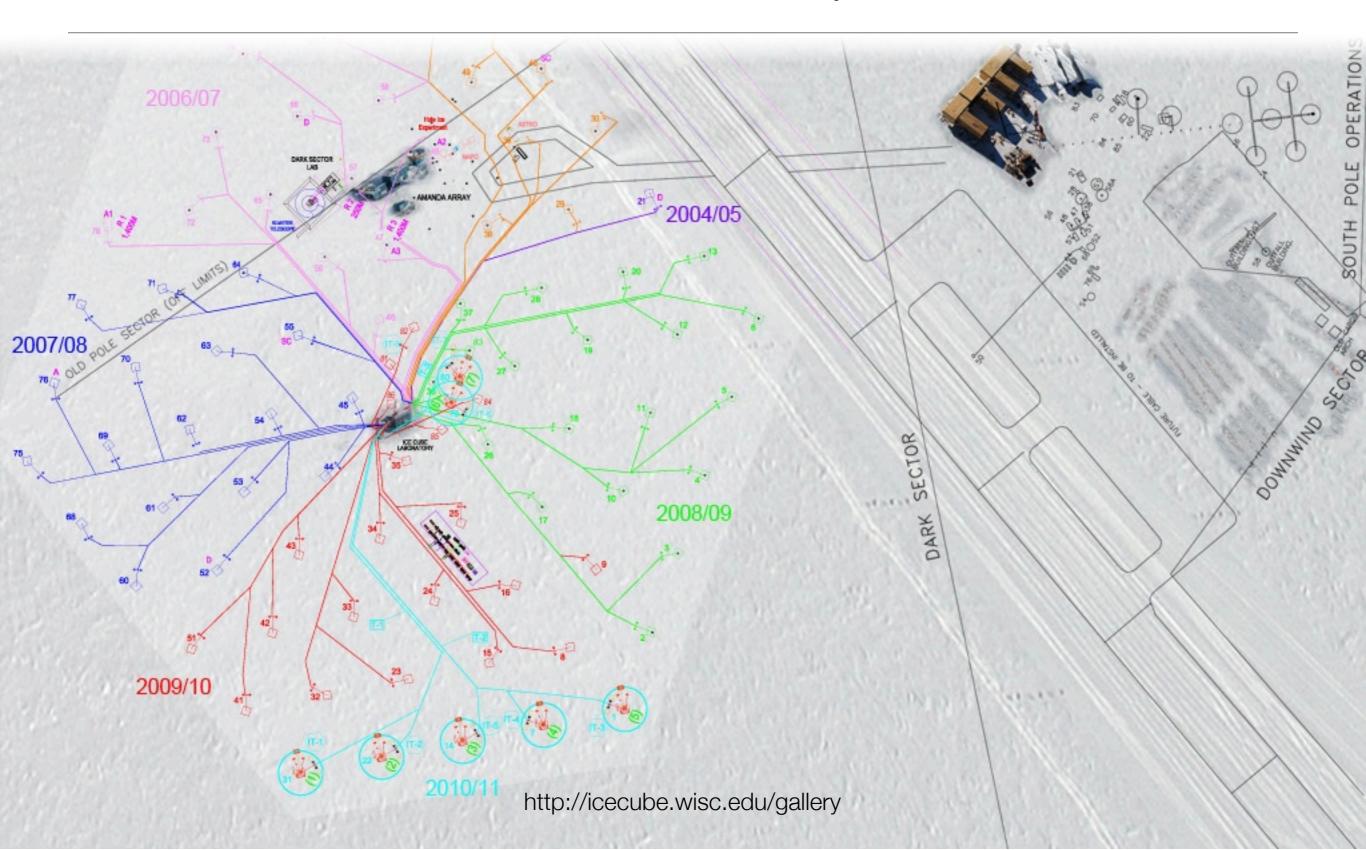
Antarctic ice at >1450m depth is the ideal detector medium



IceCube is located at the south pole

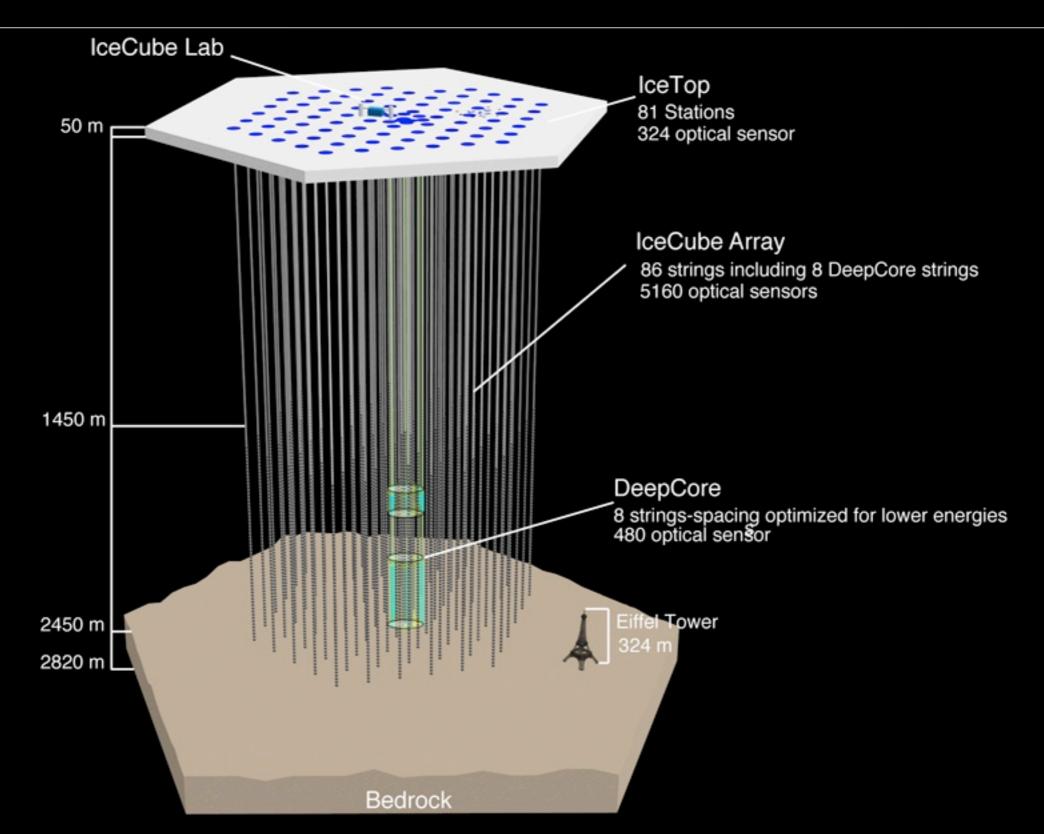


IceCube is located at the south pole

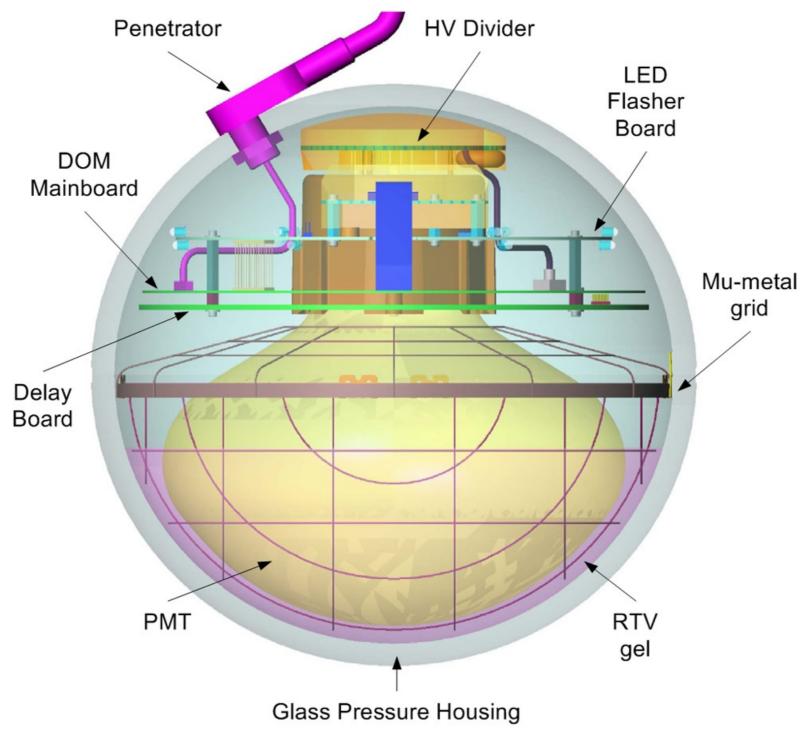


Assembly during the austral summer months with hot-water drills



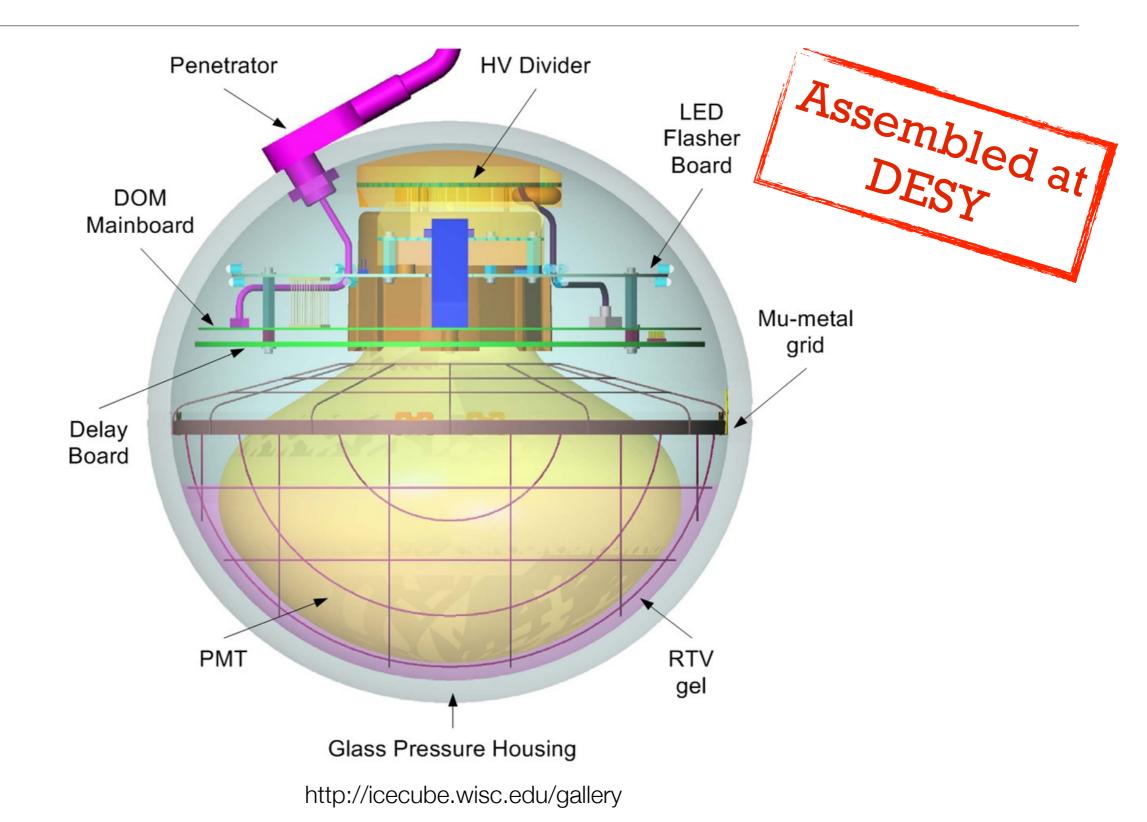


Digital-Optical-Modules send events to the surface

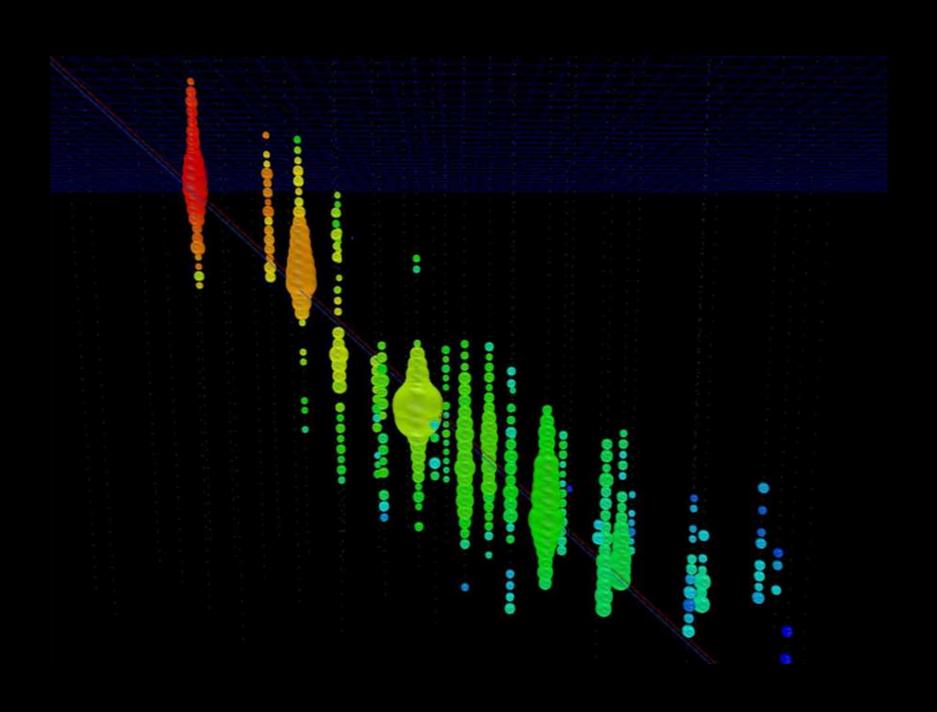


http://icecube.wisc.edu/gallery

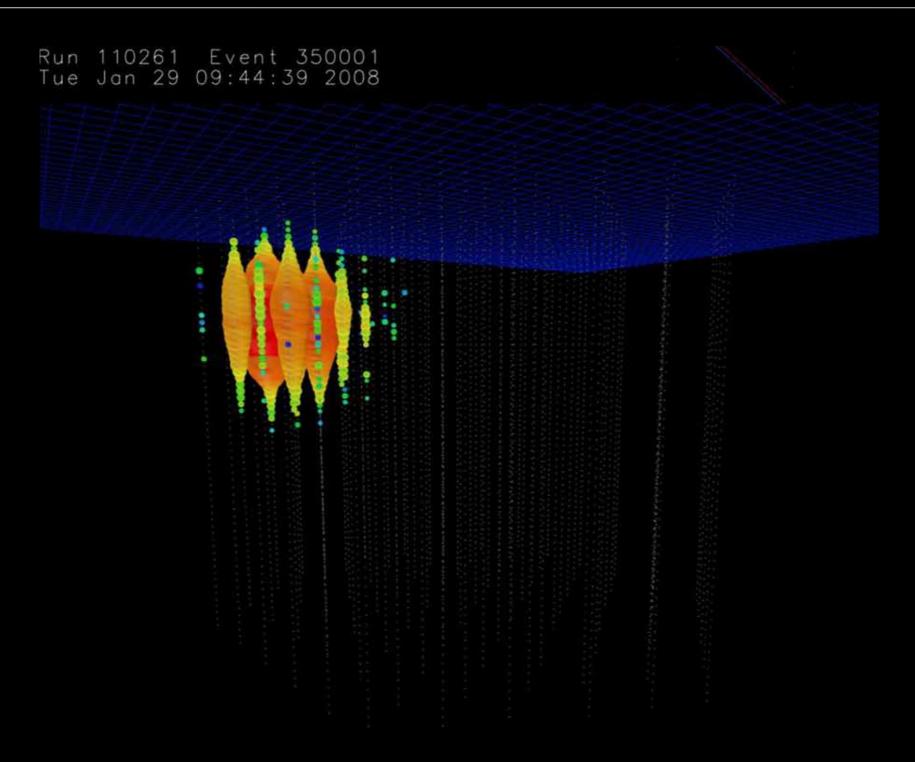
Digital-Optical-Modules send events to the surface



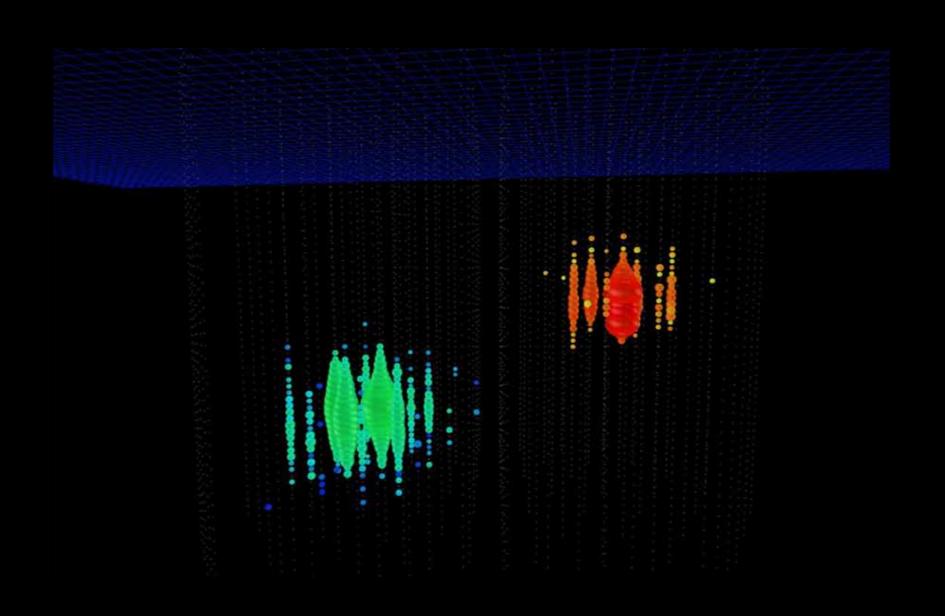
Muon neutrinos produce tracks



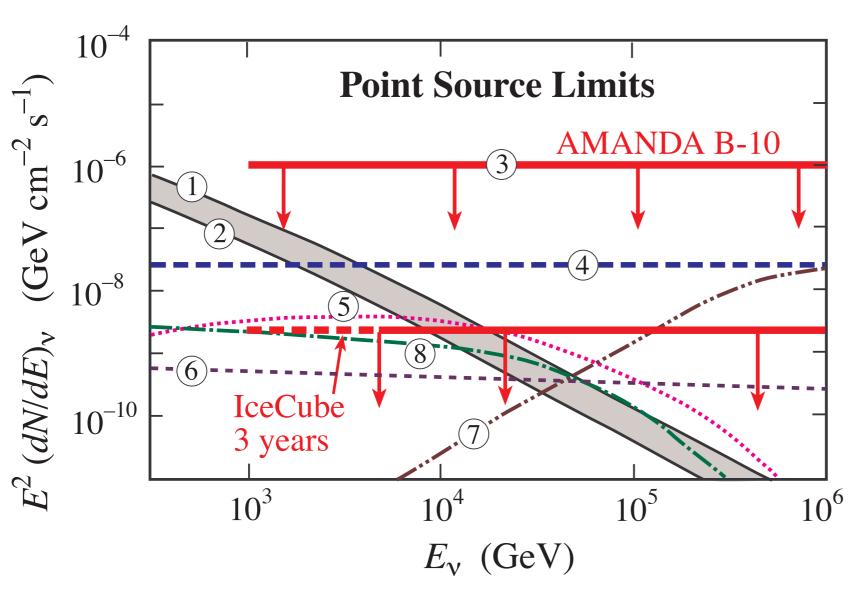
Electron neutrinos produce cascades



Tau neutrinos produce a cascade followed by a secondary cascade from tau lepton decay



IceCube Science: High-energy neutrino point sources



4: 3C273 pp-Neutrinos

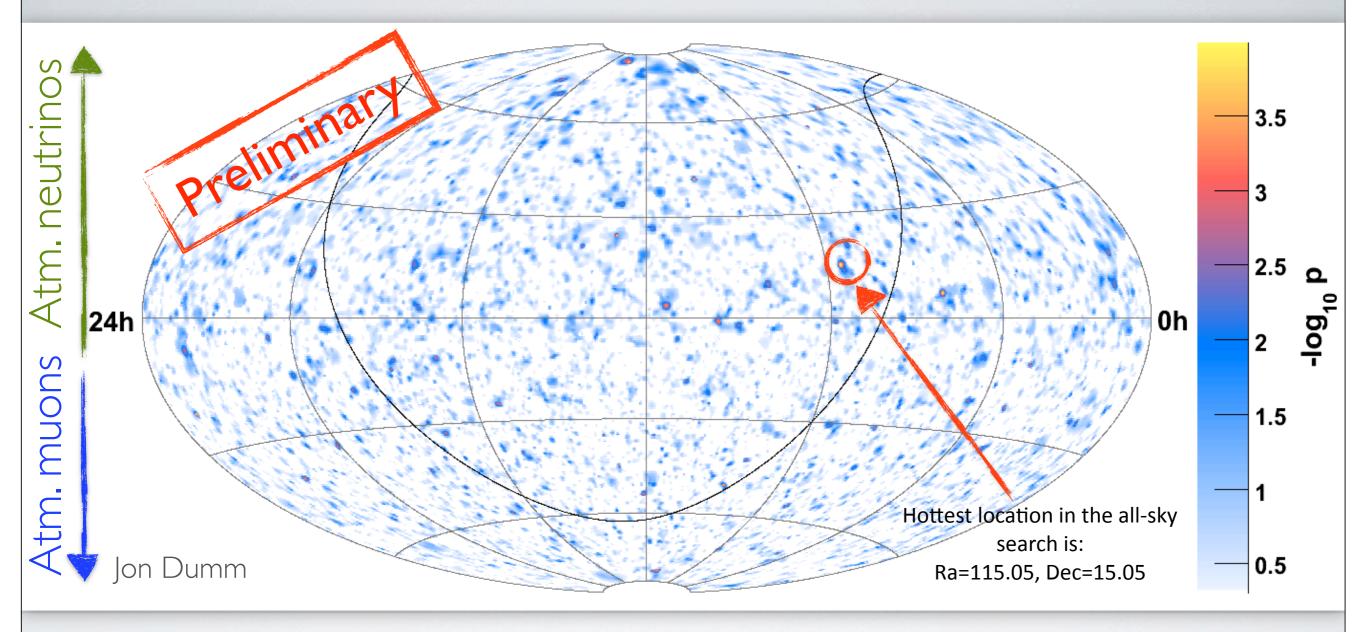
5: Crab Nebula

6: Coma Cluster

7: 3C273 py-Neutrinos

8: Supernovae

IC40 SIGNIFICANCE MAP

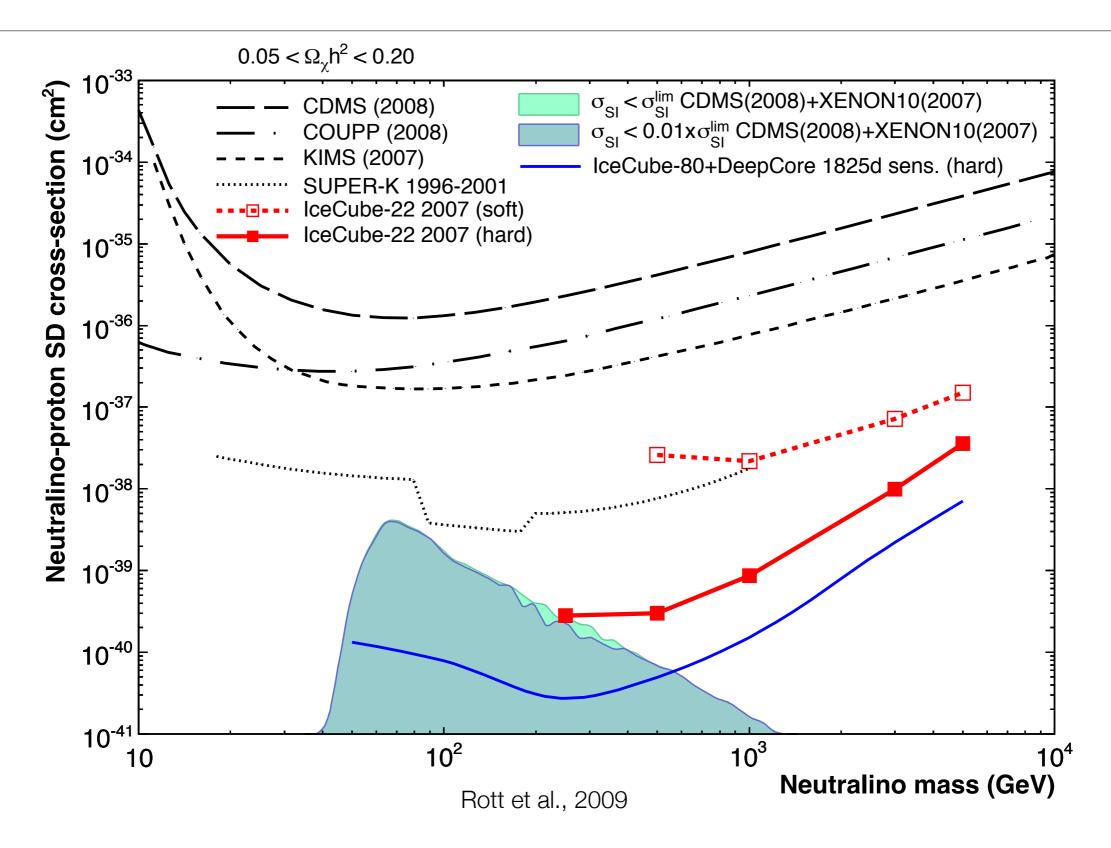


▶ Log-likelihood is calculated on a fine grid: 0.1° × 0.1°

Significance comes from the hottest single spot, calculated as the fraction of scrambled trials with equal or higher significance – robust result



IceCube Science: Search for WIMP decay



IceTop Science: Anisotropy of cosmic rays

