How cosmic rays start astrochemistry





Some basic reactions to get the chemistry started

$$H + H \xrightarrow{k_1} H_2$$

reaction I

$$H_2 \xrightarrow{k_2} H_2^+ + e^-$$

reaction II

$$H_2^+ + H_2 \xrightarrow{k_3} H_3^+$$

reaction III

Some basic reactions to get the chemistry started

$$H + H \xrightarrow{k_1} H_2$$

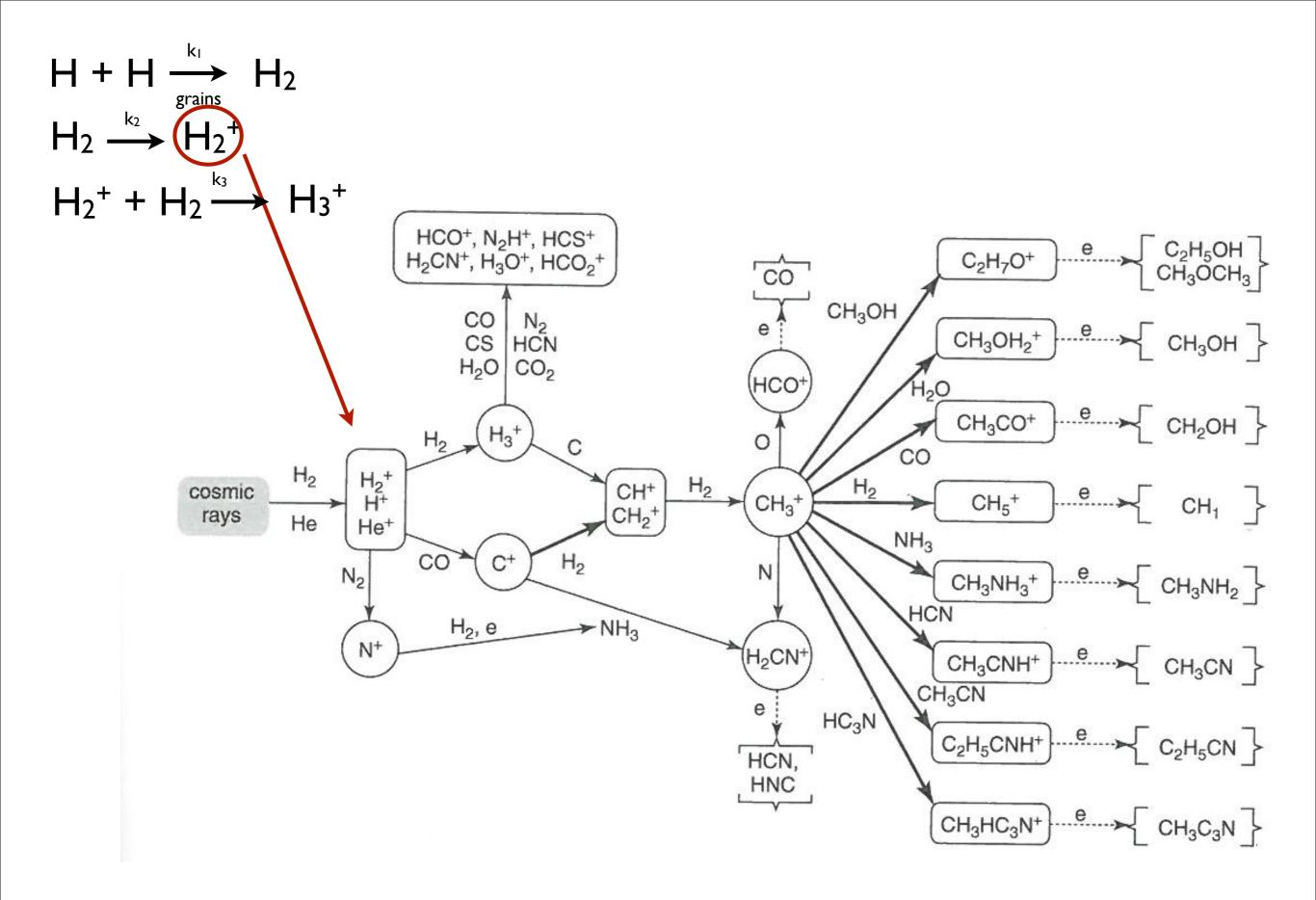
$$H_2 \xrightarrow{k_2} H_2^+ + H_2 \xrightarrow{k_3} H_3^+$$

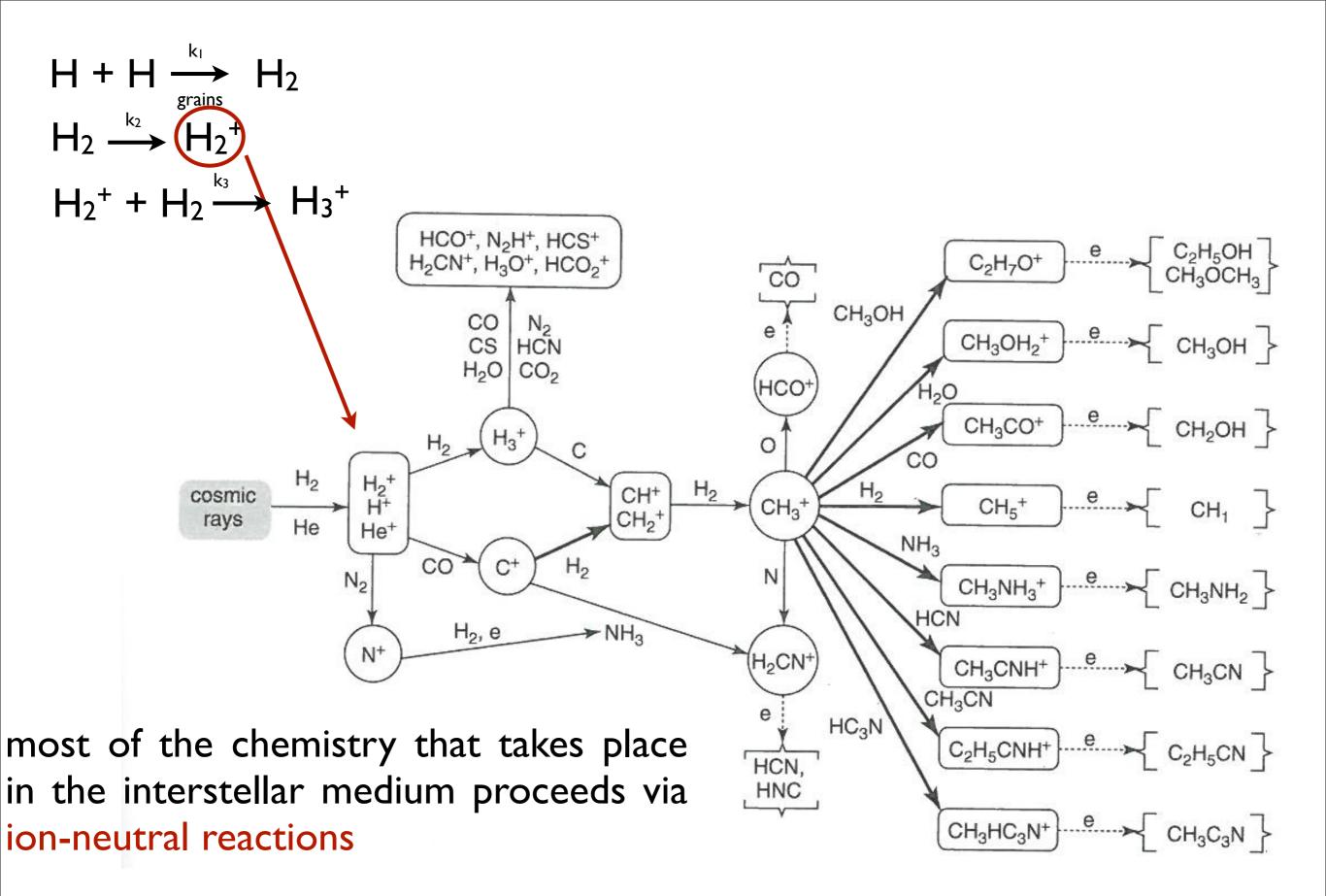
$$H_2^+ + H_2 \xrightarrow{k_3} H_3^+$$

reaction I

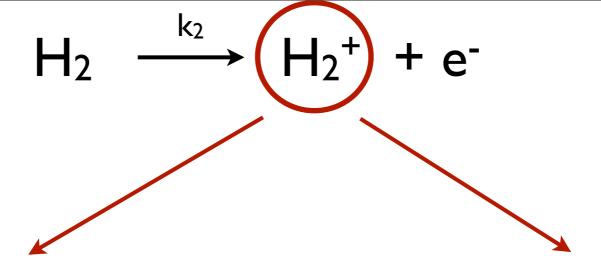
reaction II

reaction III





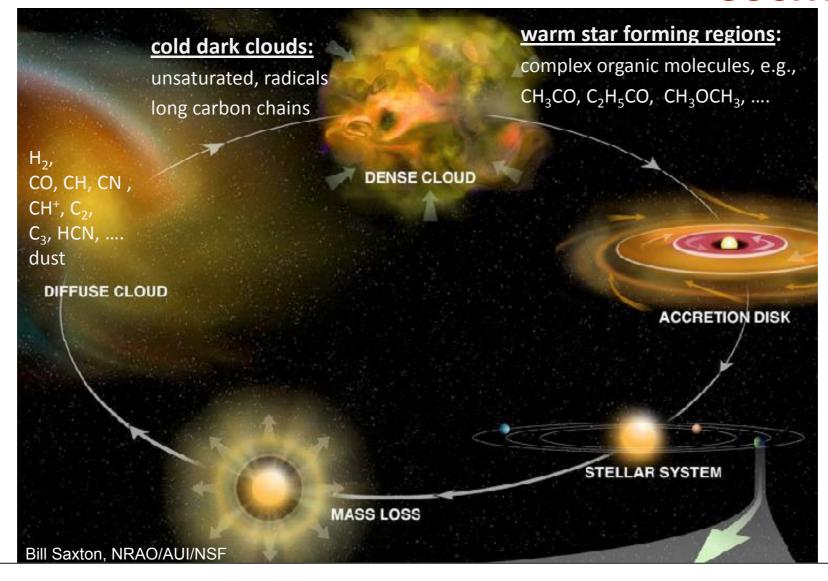
NO ENERGY BARRIER TO OVERCOME



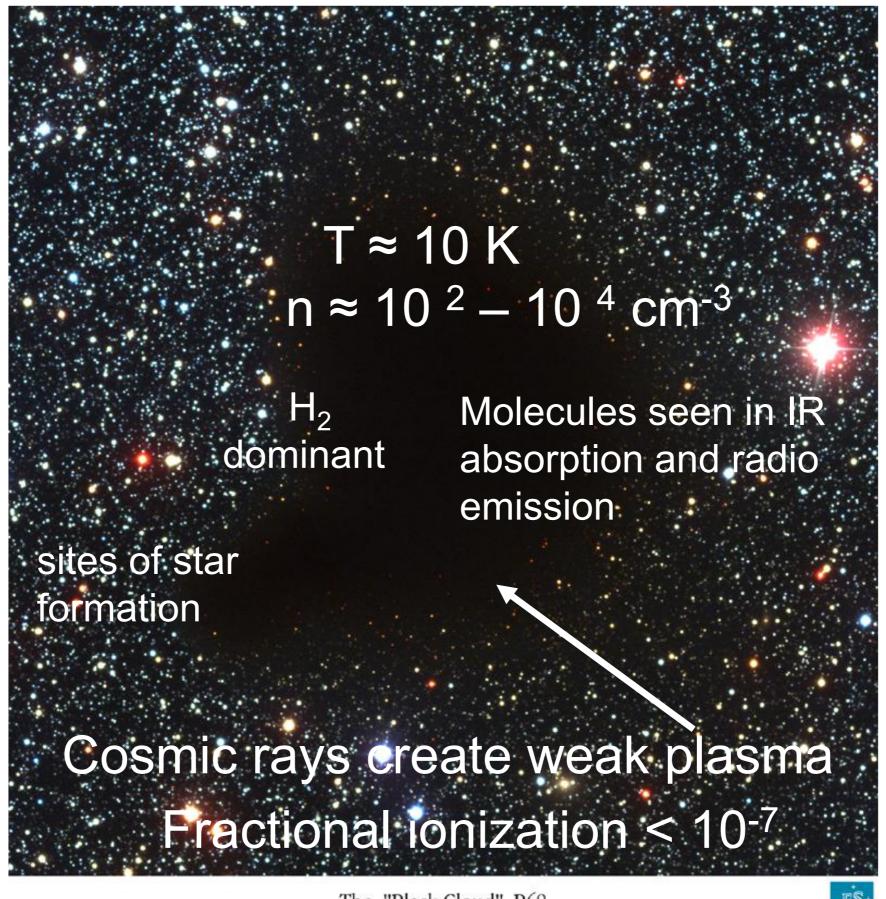
Diffuse clouds photons

Dense clouds

cosmic rays

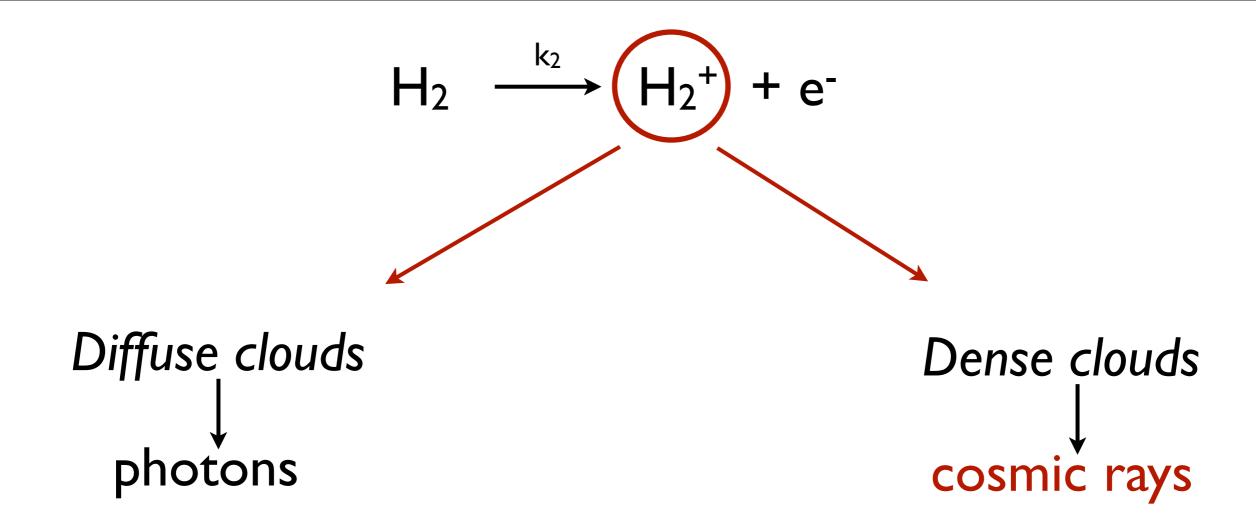


Dense interstellar clouds



The "Black Cloud" B68 (VLT ANTU + FORS1)





What cosmic rays "do" to molecules

$$AB + cr \longrightarrow AB^{+} + e^{-} + cr$$
 $AB + cr \longrightarrow A + B + cr$
 $AB + cr \longrightarrow A + B^{+} + e^{-} + cr$
 $AB + cr \longrightarrow AB^{*} + cr$

Dense clouds

$$H + H \xrightarrow{grains} H_2$$

$$H_2 \xrightarrow{k_2} H_2^+ + e^-$$

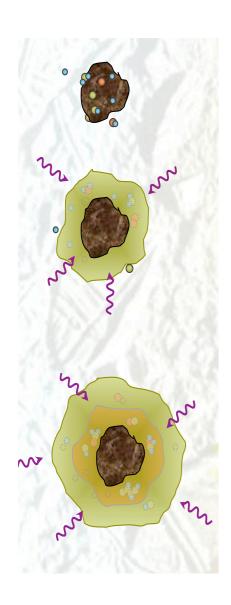
$$H_2^+ + H_2 \xrightarrow{k_3} H_3^+$$

cosmic rays ionization rate: I-5 10⁻¹⁷ s⁻¹ [Dalgarno 2006]

$$n(\mathrm{H}_3^+) = \zeta n(\mathrm{H}_2) / \left\{ \sum_{\mathrm{X}} n(\mathrm{X}) k(\mathrm{X}) + \sum_{\mathrm{M}} n(\mathrm{M}) k(\mathrm{M}) \right\}$$

$$+(\alpha n(e))+(k_0n(PAH)+k_1n(PAH^-))$$
dissociative recombination charge transfer and with electrons neutralization with PAHs

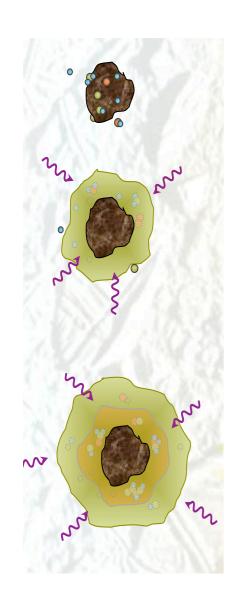
Chemistry on dust grains



Contribution from the surface need to be considered

Cosmic rays interact also with the grains, activating the chemistry on the surfaces

Chemistry on dust grains

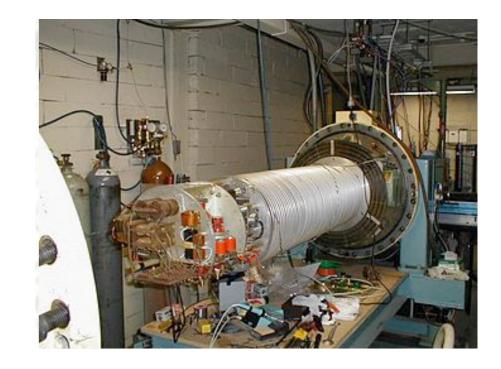


Contribution from the surface need to be considered

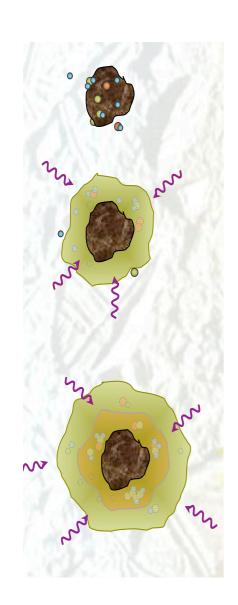
Cosmic rays interact also with the grains, activating the chemistry on the surfaces

Understanding those systems is quite complicated!

- produce a beam of cr and accelerate them to high energies (Van Der Graf accelerator)



Chemistry on dust grains



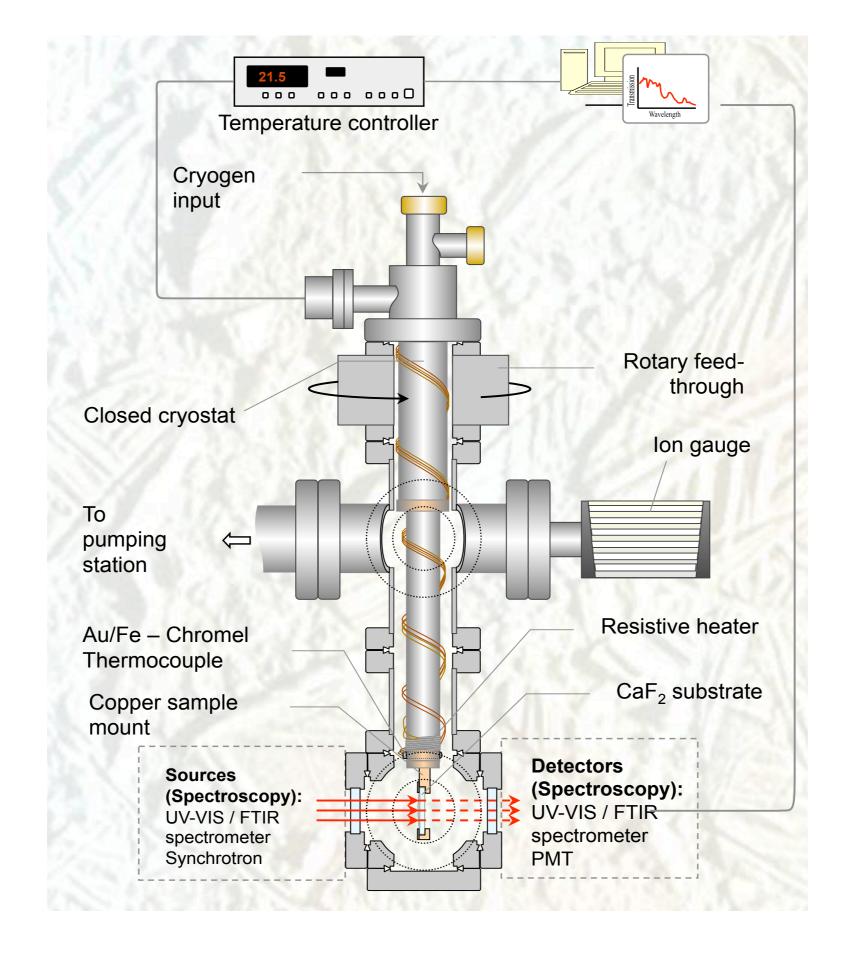
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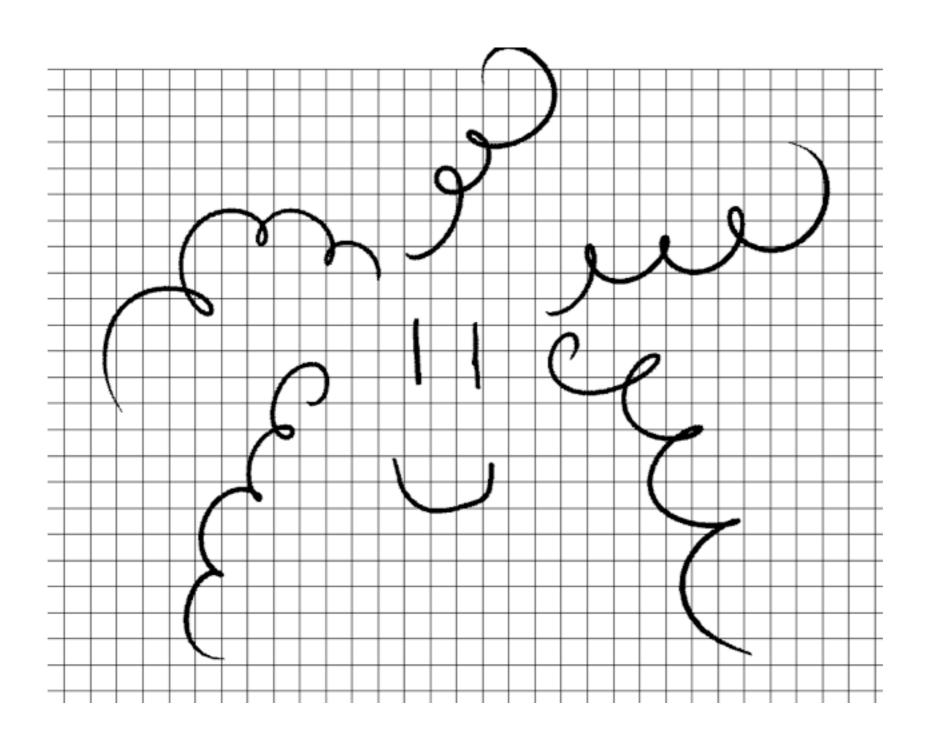
Understanding those systems is quite complicated!

- prepare the surface (target) for collision

morphology temperature mixture



credits. Nigel Mason group, Open University (UK)



Thanks!