

Near-Field VLBI Delay Models

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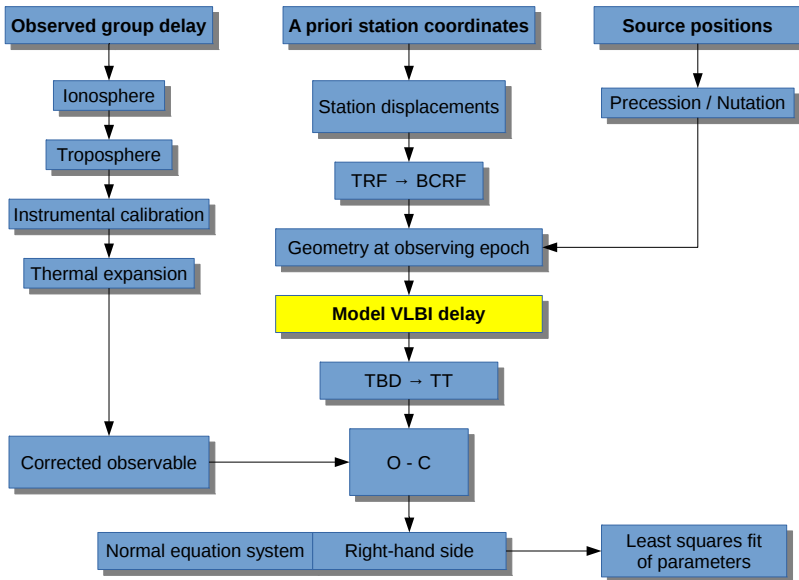
²University of Tasmania, School of Physical Sciences, Australia

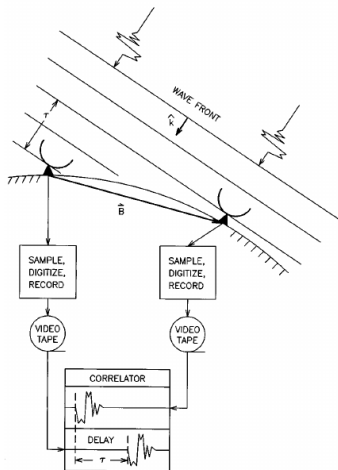
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VLBI Observations of Near-field Targets
Bonn, Germany

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- 2 Near-field VLBI delay models
 - Sekido & Fukushima (2006)
 - Duev et al. (2012)
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Geometric delay:

$$\tau \propto \vec{b} \cdot \hat{k}$$

plus corrections.

Assumption: Plane wavefronts.

⇒ **Not applicable for near-field sources.**

Figure : Schematic diagram of a VLBI experiment. Figure 1 in Sovers et al. (1998).

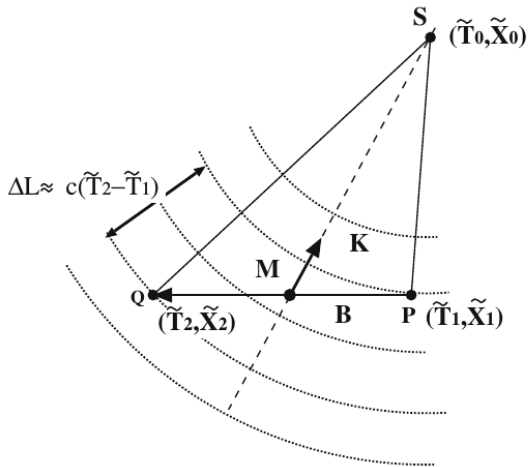


Figure : Configuration of a VLBI observation for a radio source at finite distance (Fig. 1 in Sekido & Fukushima 2006).

Light-time equation:

$$\tilde{T}_0 = \tilde{T}_1 - \frac{\left| \tilde{\mathbf{X}}_0(\tilde{T}_0) - \tilde{\mathbf{X}}_E(\tilde{T}_1) - \mathbf{R}_{1E}(\tilde{T}_1) \right|}{c} - \Delta T_{g,01}$$

Pseudo source vector:

$$\mathbf{K} \stackrel{\text{def}}{=} \frac{\mathbf{R}_1(\tilde{T}_1) + \mathbf{R}_2(\tilde{T}_1)}{R_1(\tilde{T}_1) + R_2(\tilde{T}_1)}$$

with

$$\begin{aligned} \mathbf{R}_i(\tilde{T}_1) &= \tilde{\mathbf{X}}_0(\tilde{T}_0) - \tilde{\mathbf{X}}_i(\tilde{T}_1) \\ &= \tilde{\mathbf{X}}_0(\tilde{T}_0) - \tilde{\mathbf{X}}_E(\tilde{T}_1) - \mathbf{R}_{iE}(\tilde{T}_1) \end{aligned}$$

Sekido & Fukushima (2006)

VLBI delay for a radio source at finite distance in the TT frame:

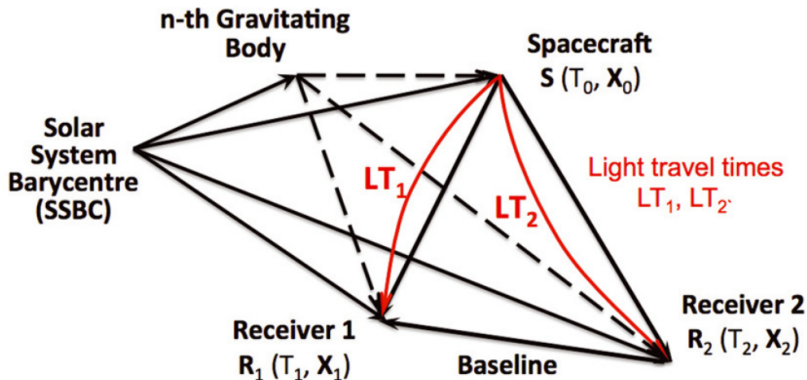
$$\tau = \frac{-\frac{\mathbf{K} \cdot \mathbf{b}}{c} \left(1 - \frac{2W_E}{c^2} - \frac{V_E^2 + 2\mathbf{V}_E \cdot \mathbf{v}_2}{2c^2}\right) - \frac{\mathbf{V}_E \cdot \mathbf{b}}{c^2} \left(1 + \frac{\hat{\mathbf{R}}_2 \cdot \mathbf{V}_2}{c} - \frac{(\mathbf{V}_E + 2\mathbf{v}_2) \cdot \mathbf{K}}{2c}\right) + \Delta T_{\text{grav}}}{\left(1 + \frac{\hat{\mathbf{R}}_2 \cdot \mathbf{V}_2}{c}\right) (1 + H)}$$

Eq. (20) in Sekido & Fukushima (2006)

Infinite distance delay:

$$\tau = \frac{-\frac{\hat{\mathbf{k}} \cdot \mathbf{b}}{c} \left(1 - \frac{2W_E}{c^2} - \frac{V_E^2 + 2\mathbf{V}_E \cdot \mathbf{v}_2}{2c^2}\right) - \frac{\mathbf{V}_E \cdot \mathbf{b}}{c^2} \left(1 + \frac{\hat{\mathbf{k}} \cdot \mathbf{V}_E}{2c}\right) + \Delta T_{\text{grav}}}{1 + \frac{\hat{\mathbf{k}} \cdot (\mathbf{V}_E + \mathbf{v}_2)}{c}}$$

IERS Conventions (2010, Chapt. 11)



Transformation from TBD to TT:

$$\tau = \left(\frac{T_2 - T_1}{1 - L_C} \cdot \left[1 - \frac{1}{c^2} \left(\frac{V_E^2}{2} + U_E \right) \right] - \frac{\mathbf{V}_E \cdot \mathbf{b}}{c^2} \right) \cdot \left(1 + \frac{\mathbf{V}_E \cdot \dot{\mathbf{r}}_{2,gc}}{c^2} \right)^{-1}$$

Duev *et al.* (2012)

- August 24, 2015, 12:00 – 16:00 UTC
- Baseline Hobart – Ceduna (~ 1700 km)
- Observed satellites: PRN02, 12, 24, 25
- SP3 final ephemerides

Plank *et al.* (in preparation)
Hellerschmied *et al.* IVS GM 2016

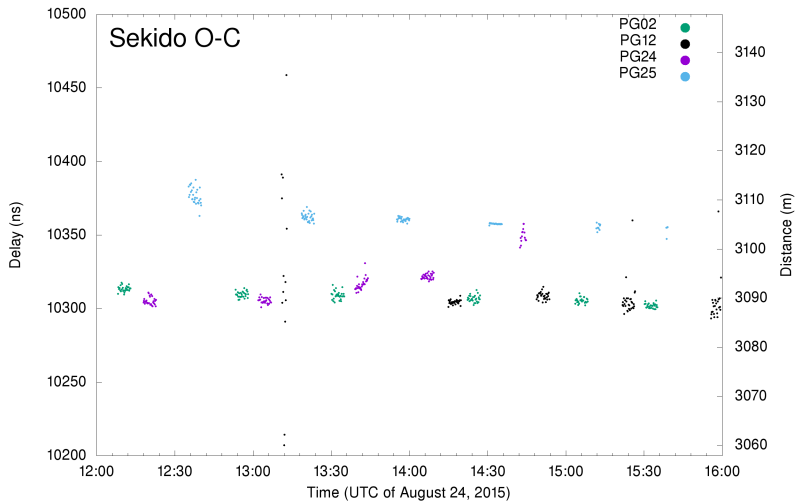


Figure : Observed delays minus computed with Sekido & Fukushima (2006).

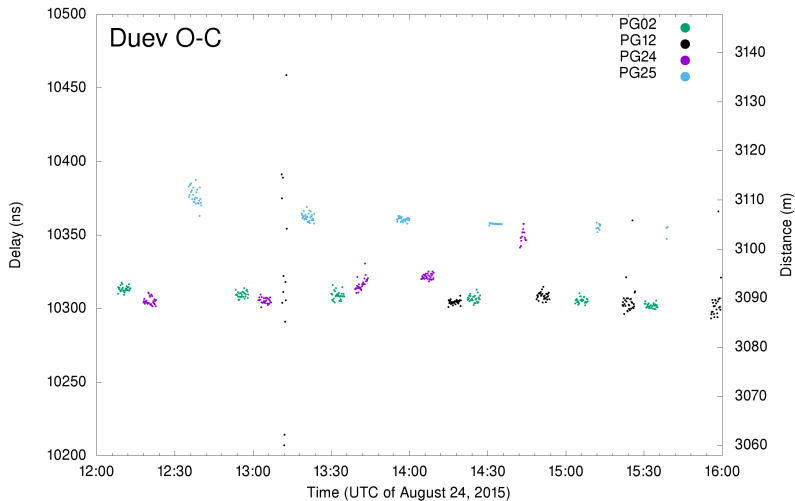


Figure : Observed delays minus computed with Duev *et al.* (2012).

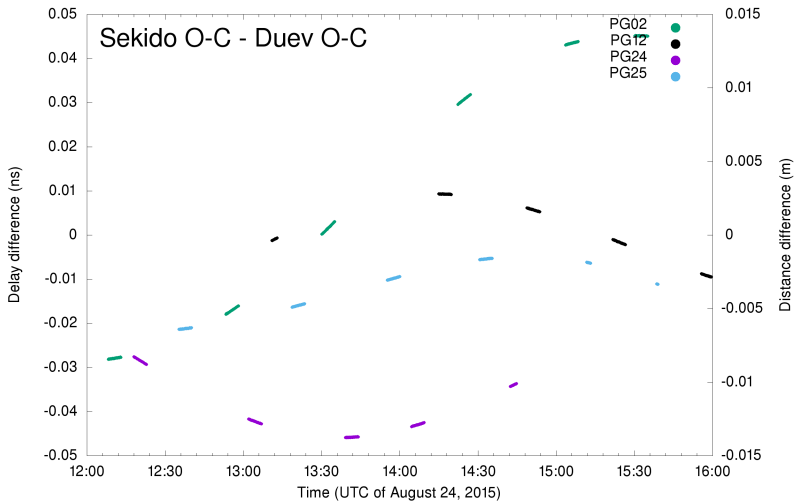


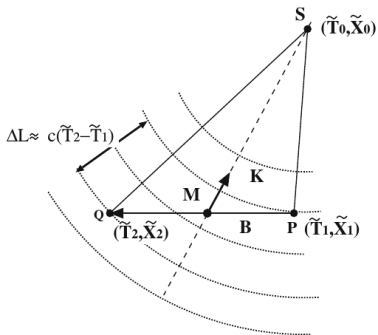
Figure : Difference between Sekido & Fukushima (2006) and Duev *et al.* (2012).



Two near-field VLBI delay models



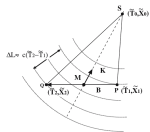
Two near-field VLBI delay models



Sekido & Fukushima (2006)



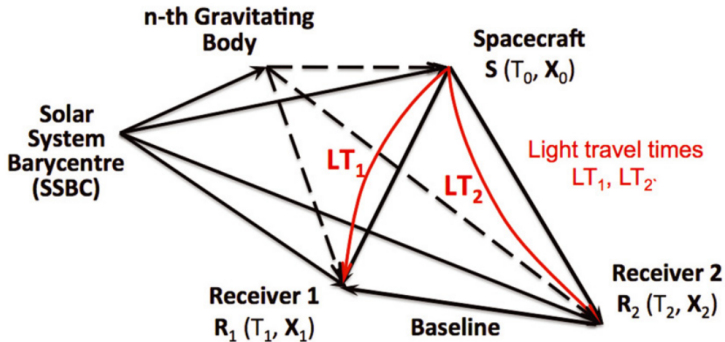
Two near-field VLBI delay models



Sekido & Fukushima (2006) (pseudo source vector)



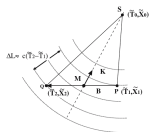
Two near-field VLBI delay models



Duev *et al.* (2012)



Two near-field VLBI delay models



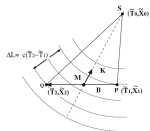
Sekido & Fukushima (2006) (pseudo source vector)



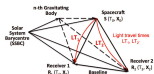
Duev *et al.* (2012) (solve light-time equation twice)



Two near-field VLBI delay models



Sekido & Fukushima (2006) (pseudo source vector)

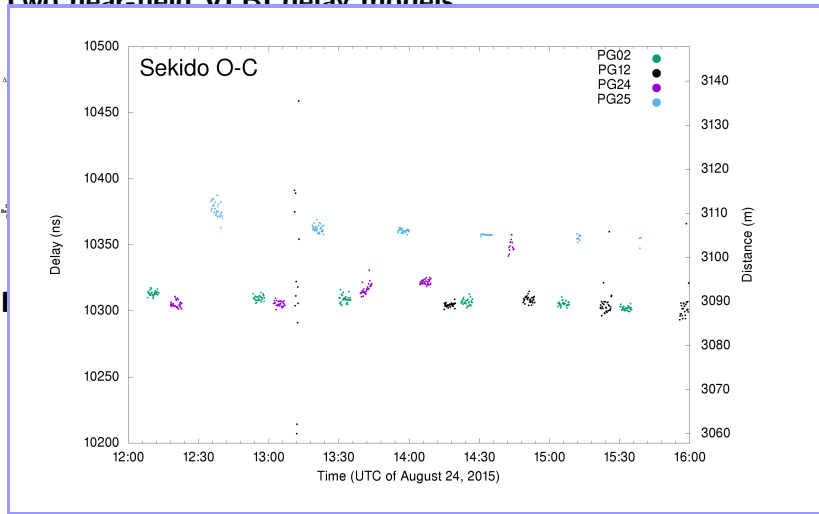


Duev *et al.* (2012) (solve light-time equation twice)

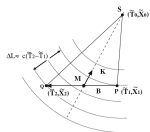
Results with our implementation



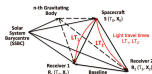
Two near-field VLBI delay models



Two near-field VLBI delay models

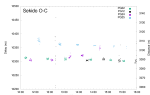


Sekido & Fukushima (2006) (pseudo source vector)



Duev *et al.* (2012) (solve light-time equation twice)

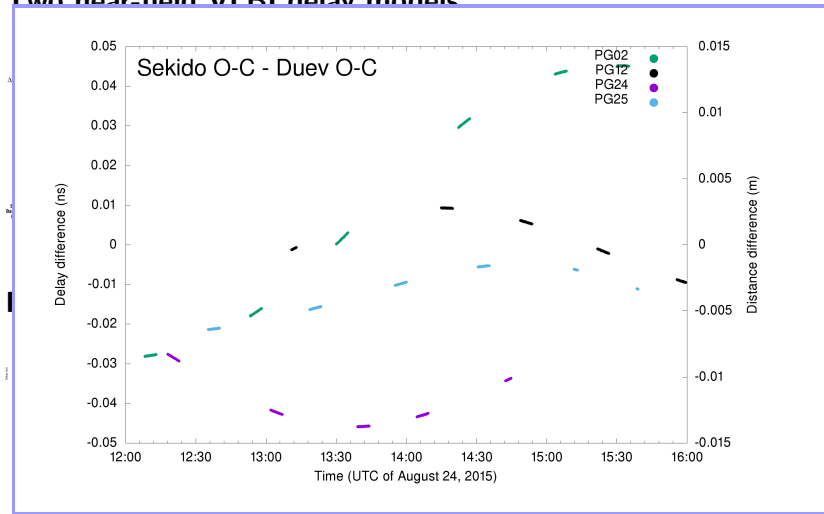
Results with our implementation



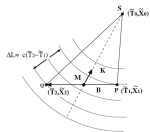
GPS satellites have individual systematics



Two near-field VLBI delay models



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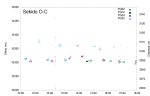


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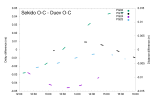


Duev *et al.* (2012) (solve light-time equation twice)

Results with our implementation



GPS satellites have individual systematics



Sekido – Duev $\approx \pm 50$ ps



Appendix



Figure : Artist's impression of a GPS IIF satellite in orbit (Source: Wikipedia).