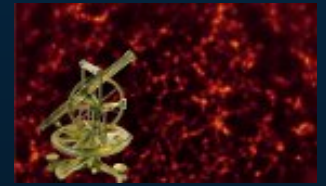


# Bayesian Large Scale Structure inference

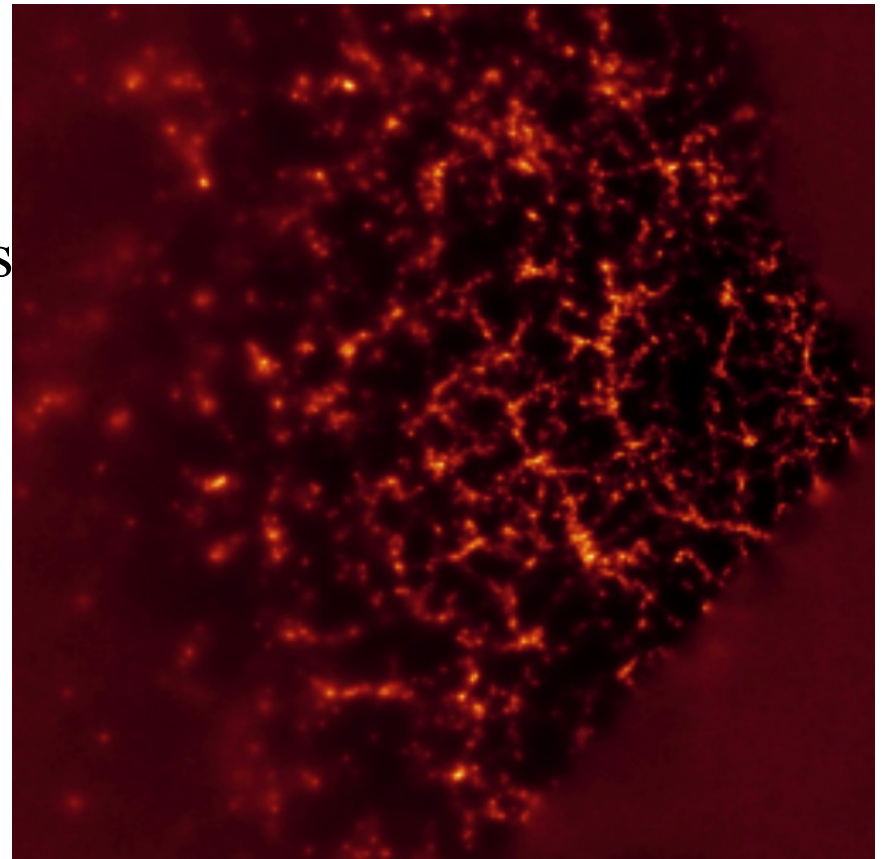
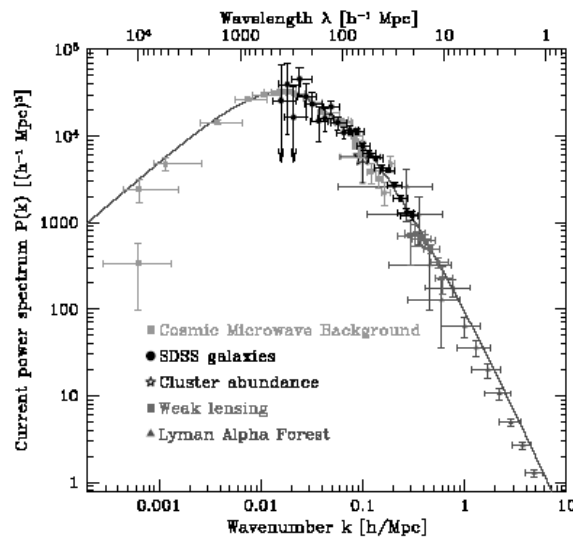
*Jens Jasche*

*Heidelberg, 20 September 2011*

# Motivation



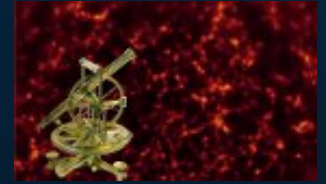
- Goal: 3D cosmography
  - Map matter distribution
  - Quantify statistical properties



Tegmark et al. (2004)

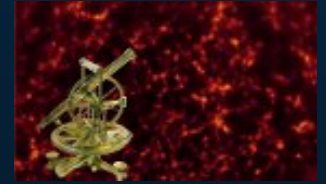
Jasche et al. (2010)

# Motivation



- No ideal Observations in reality

# Motivation

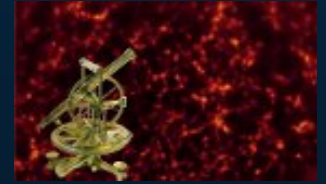


- No ideal Observations in reality
  - Statistical uncertainties



Noise, cosmic variance etc.

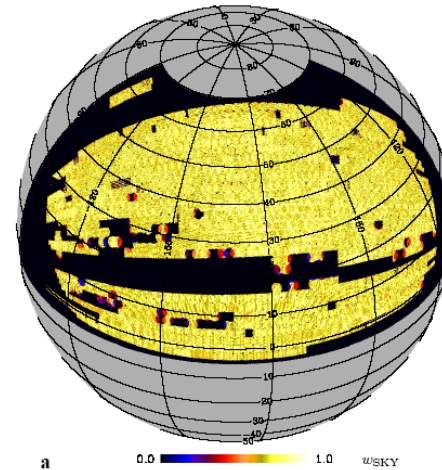
# Motivation



- No ideal Observations in reality
  - Statistical uncertainties
  - Systematics



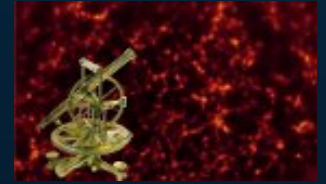
Noise, cosmic variance etc.



Kitaura et al. (2009)

Survey geometry,  
selection effects,  
biases

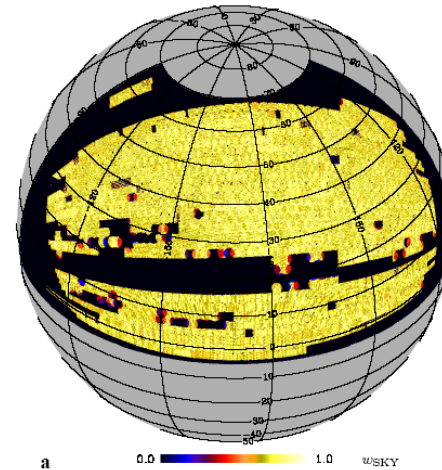
# Motivation



- No ideal Observations in reality
  - Statistical uncertainties
  - Systematics



Noise, cosmic variance etc.



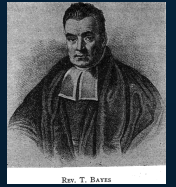
Kitaura et al. (2009)

Survey geometry,  
selection effects,  
biases

→ No unique recovery possible!!!

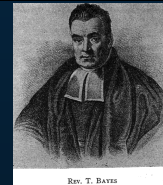


# Bayesian Approach



**“Which are the possible signals compatible with the observations?”**

# Bayesian Approach



**“Which are the possible signals compatible with the observations?”**

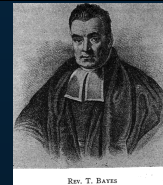
- Object of interest: Signal posterior distribution

$$\mathcal{P}(s|d) = \mathcal{P}(s) \frac{\mathcal{P}(d|s)}{\mathcal{P}(d)}$$





# Bayesian Approach



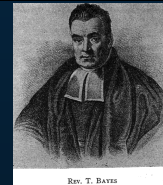
**“Which are the possible signals compatible with the observations?”**

- Object of interest: Signal posterior distribution

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# Bayesian Approach



**“Which are the possible signals compatible with the observations?”**

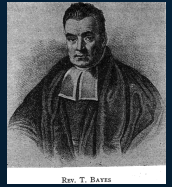
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$$\mathcal{P}(s|d) = \mathcal{P}(s) \frac{\mathcal{P}(d|s)}{\mathcal{P}(d)}$$

Likelihood



# Bayesian Approach



**“Which are the possible signals compatible with the observations?”**

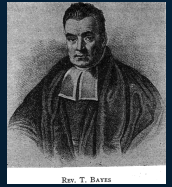
- Object of interest: Signal posterior distribution

$$\mathcal{P}(s|d) = \mathcal{P}(s) \frac{\mathcal{P}(d|s)}{\mathcal{P}(d)}$$

Evidence



# Bayesian Approach



**“Which are the possible signals compatible with the observations?”**

- Object of interest: Signal posterior distribution

$$\mathcal{P}(s|d) = \mathcal{P}(s) \frac{\mathcal{P}(d|s)}{\mathcal{P}(d)}$$

- We can do science!
  - Model comparison
  - Parameter studies
  - Report statistical summaries
  - Non-linear, Non-Gaussian error propagation



# LSS Posterior

- Aim: inference of non-linear density fields
  - non-linear density field
    - lognormal See e.g. Coles & Jones (1991), Kayo et al. (2001)
  - “correct” noise treatment
    - full Poissonian distribution

- Problem: Non-Gaussian sampling in high dimensions

- direct sampling not possible
- usual Metropolis-Hastings algorithm inefficient

➔ **HADES** (HAmiltonian Density Estimation and Sampling)

Jasche, Kitaura (2010)



# LSS inference with the SDSS

- Application of HADES to SDSS DR7
  - cubic, equidistant box with sidelength 750 Mpc
  - $\sim 3$  Mpc grid resolution
  - $\sim 10^7$  volume elements / parameters

Jasche, Kitaura, Li, Enßlin (2010)

Nasir, Bayesian LSS Inference

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  - to provide 3D cosmographic descriptions
  - to quantify uncertainties of the density distribution

Jasche, Kitaura, Li, Enßlin (2010)

Nasir, Bayesian LSS Inference



# LSS inference with the SDSS

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- Goal: provide a representation of the SDSS density posterior
  - to provide 3D cosmographic descriptions
  - to quantify uncertainties of the density distribution
    - 3 TB, 40,000 density samples

Jasche, Kitaura, Li, Enßlin (2010)

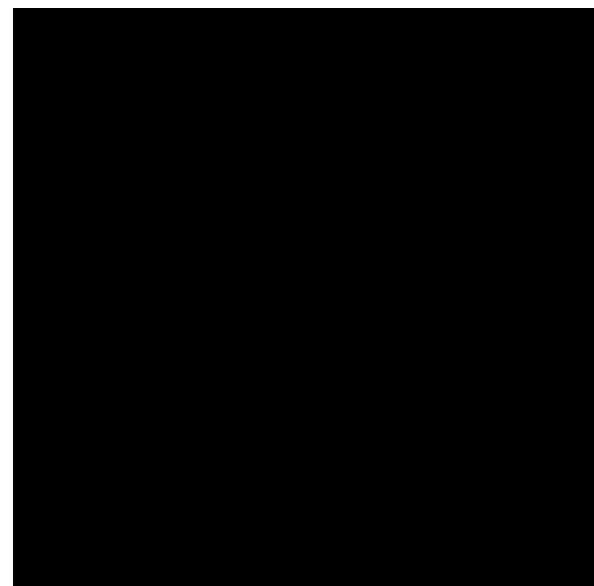
Galaxies, Bayesian LSS Inference

# LSS inference with the SDSS

What are the possible density fields  
Compatible with the observations ?

# LSS inference with the SDSS

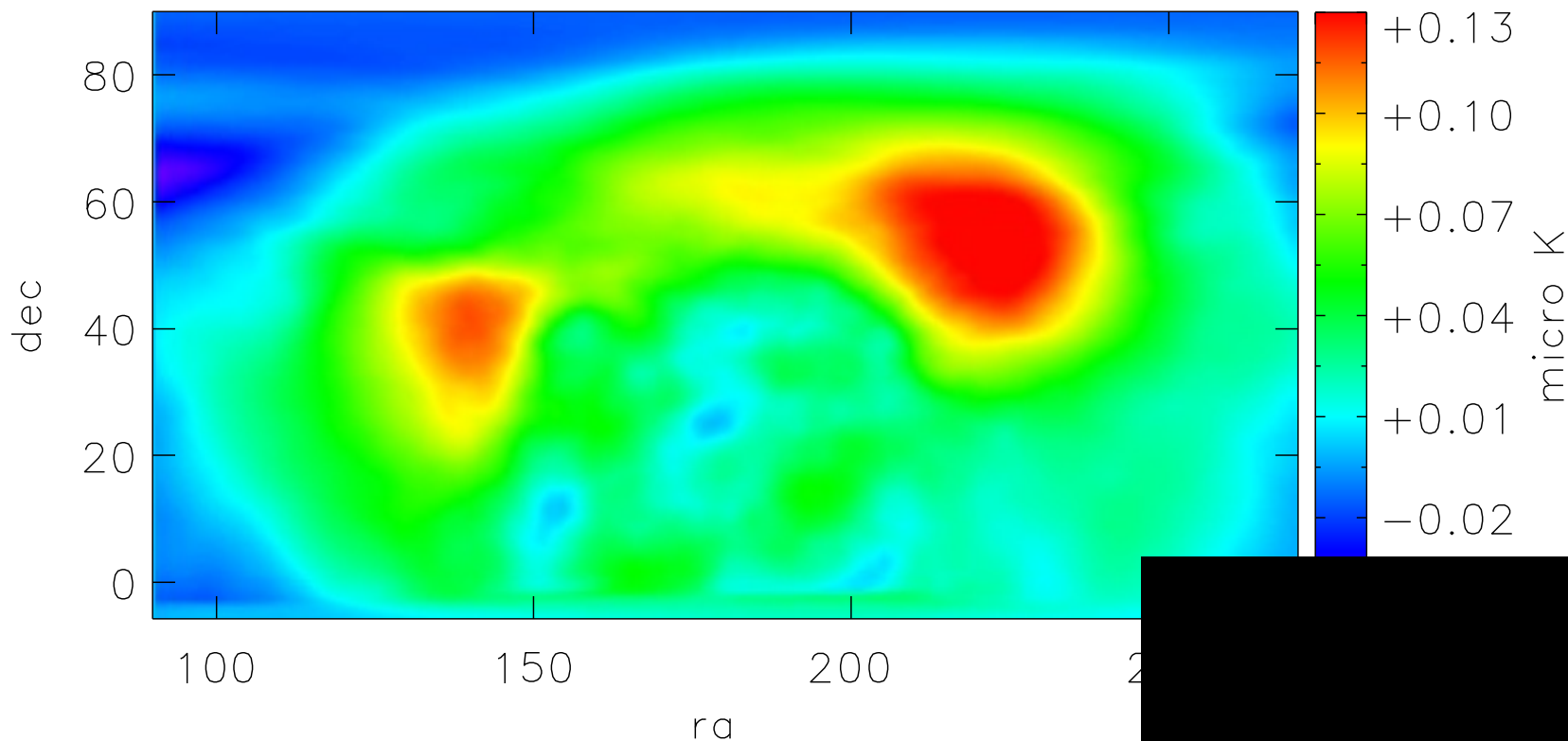
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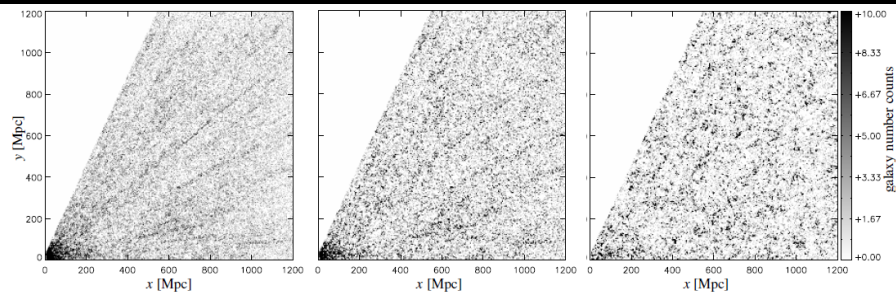
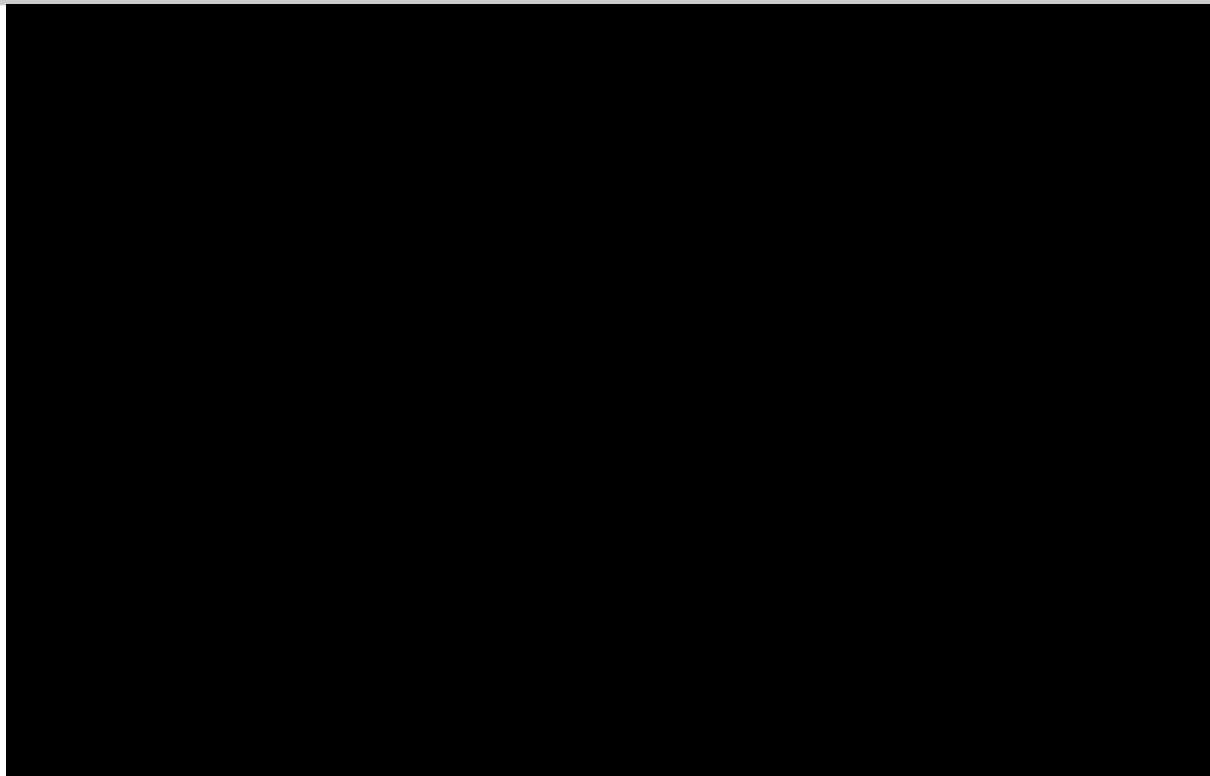


# LSS inference with the SDSS

## □ ISW-templates



# Photometric redshift sampling



Jasche, Wandelt (2011)

# Summary & Conclusion

- Bayesian statistics answers the question:
  - “What is the probability distribution of possible signals compatible with the data?”
  - signal posterior distribution
  
- LSS inference framework: HADES
  - Correction of systematic effects, survey geometry, selection effects
  - Precision non-linear density inference
  - accurate treatment of Poissonian shot noise
  - non-linear, non-Gaussian error propagation



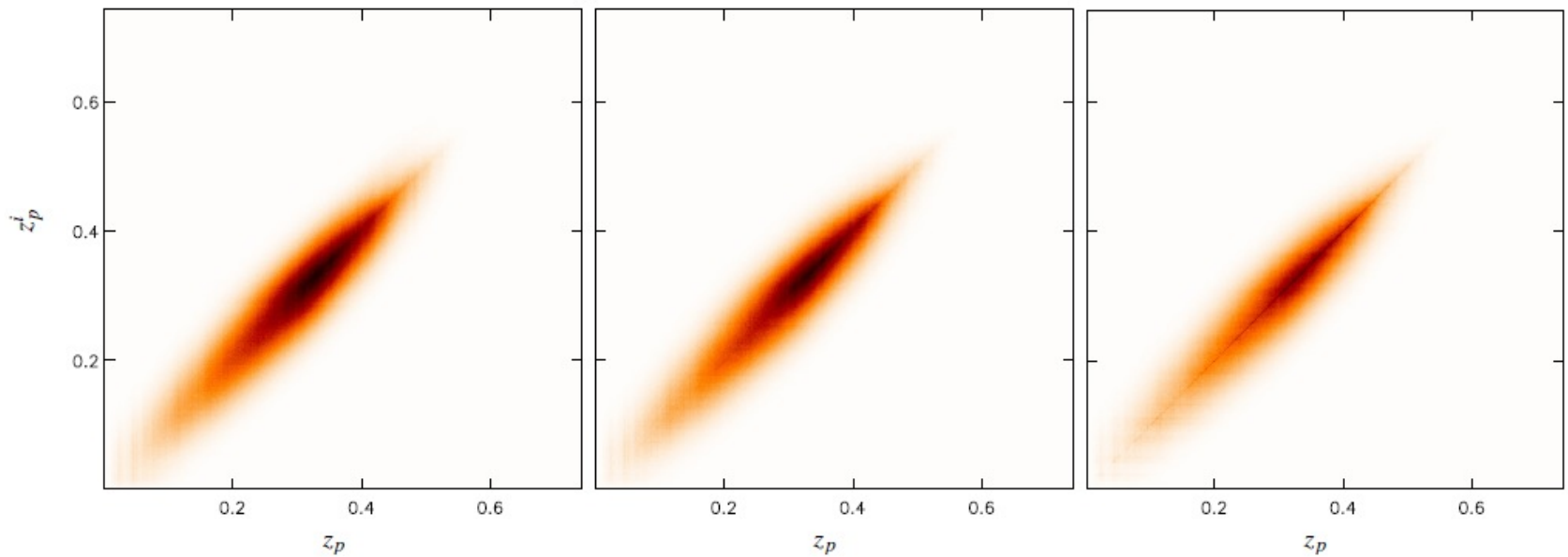




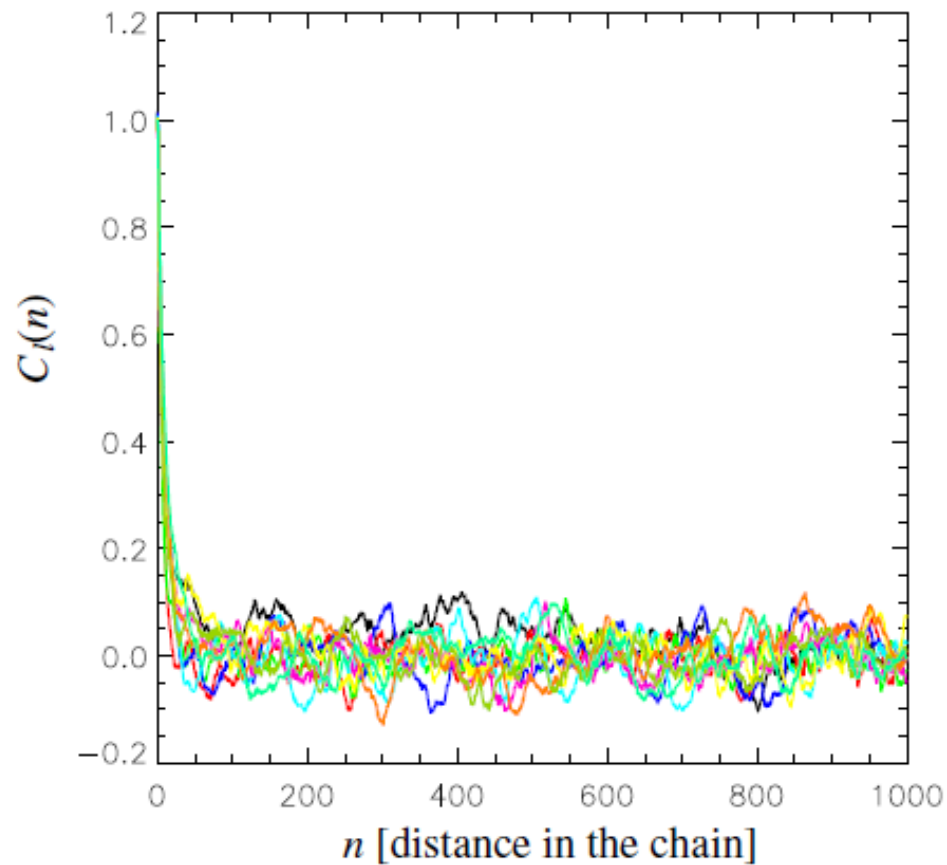
# The End ...

Thank you

# Preliminary results

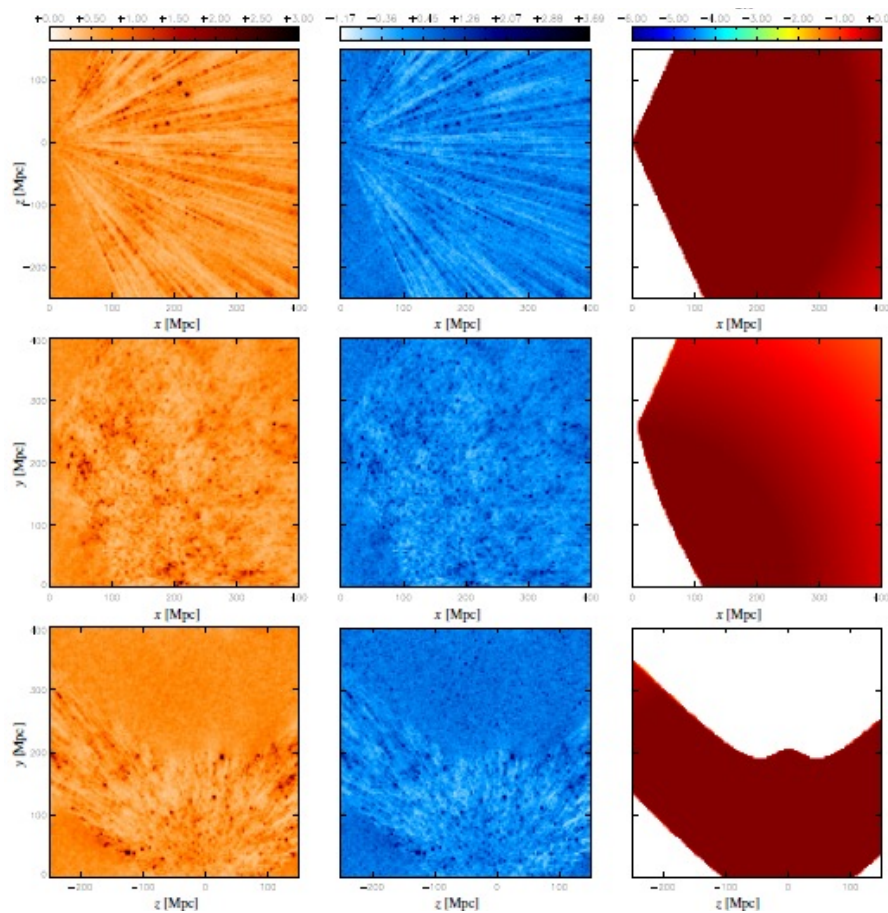


# Preliminary results

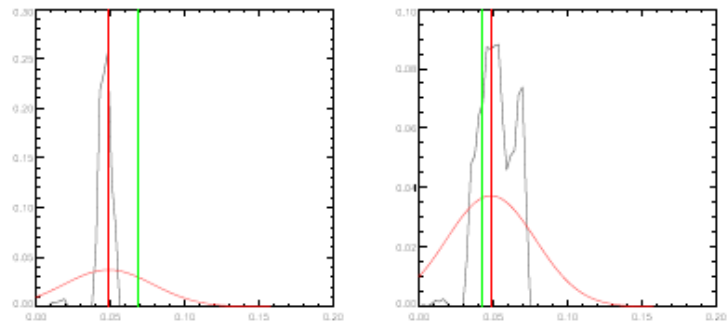


# Preliminary results

Density field



Marginalized redshift posteriors



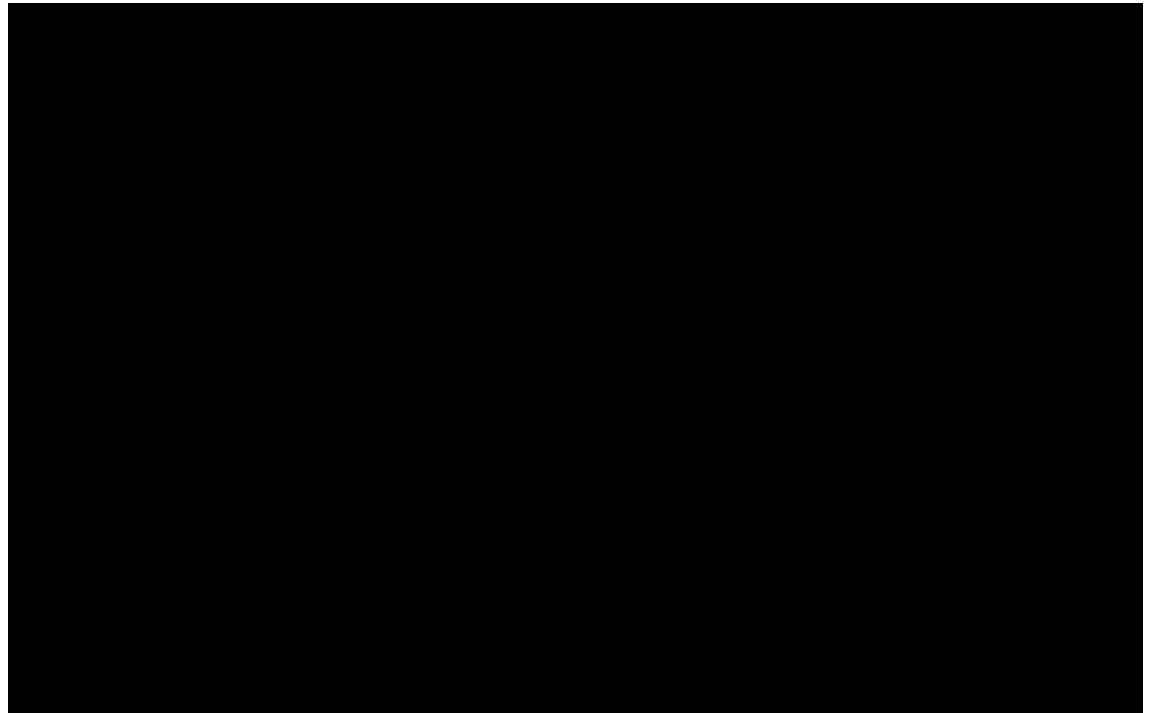
SS Inference

# The End ...

## LSS inference with the GAMA survey

By

Jens Jasche, Cristiano Porciani and Peder Norberg



# MCMC sampling

- Problem: How to evaluate the posterior distribution?
  - High dimensional parameter spaces ( $>10^6$  volume elements)
  - Difficult functional shapes
  - Non-linear density inference
- Approximate posterior distribution

$$\mathcal{P}(s|d) = \frac{1}{N} \sum_{i=1}^N \delta^D(s - s_i).$$

➔ We need efficient MCMC algorithms!



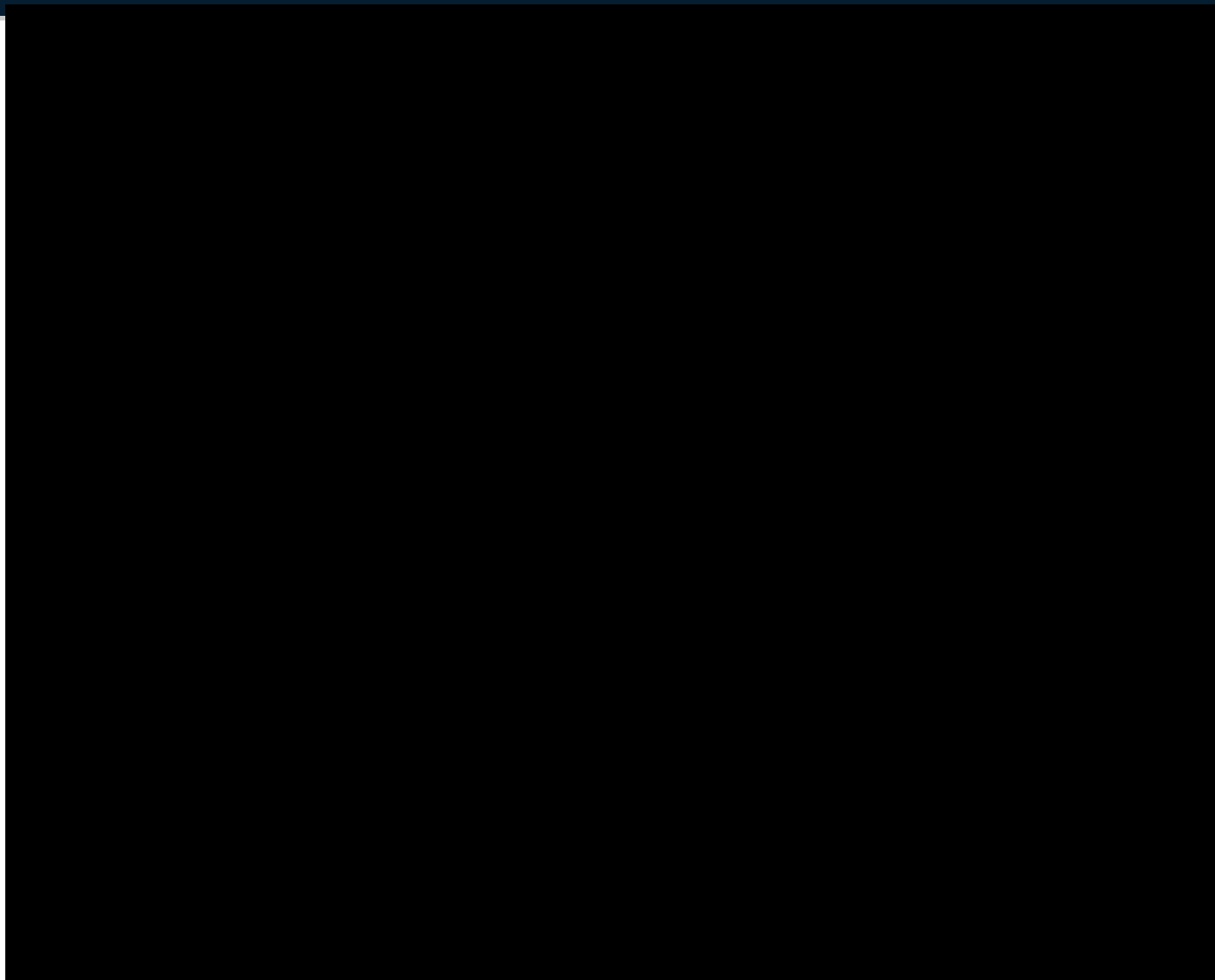


# The End ...

Thank you



# LSS inference with the GAMA survey



# Extensions of the sampler

## □ Multiple block Metropolis Hastings sampling

$$(\xi_1, \xi_2^j, \dots, \xi_M^j) \curvearrowright \mathcal{P}(\xi_1, \xi_2^j, \dots, \xi_M^j | d)$$

$$1) \quad \xi_1^{(j+1)} \curvearrowright \mathcal{P}(\xi_1 | \xi_2^j, \dots, \xi_M^j, d)$$

$$2) \quad \xi_2^{(j+1)} \curvearrowright \mathcal{P}(\xi_2 | \xi_1^{j+1}, \xi_3^j, \dots, \xi_M^j, d)$$

⋮

⋮

⋮

$$M) \quad \xi_M^{(j+1)} \curvearrowright \mathcal{P}(\xi_M | \xi_1^{j+1}, \xi_2^{j+1}, \dots, \xi_{M-1}^{j+1})$$

# Extensions of the sampler

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⋮

⋮

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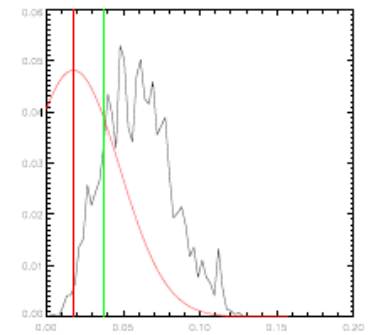
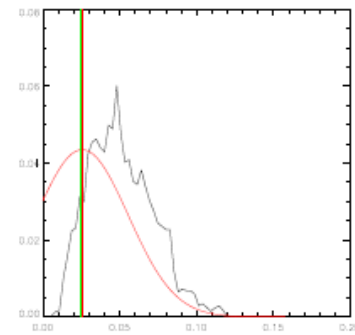
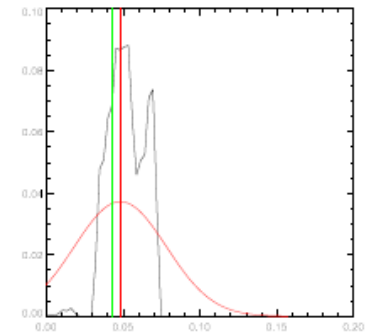
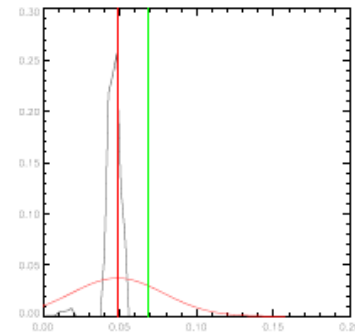
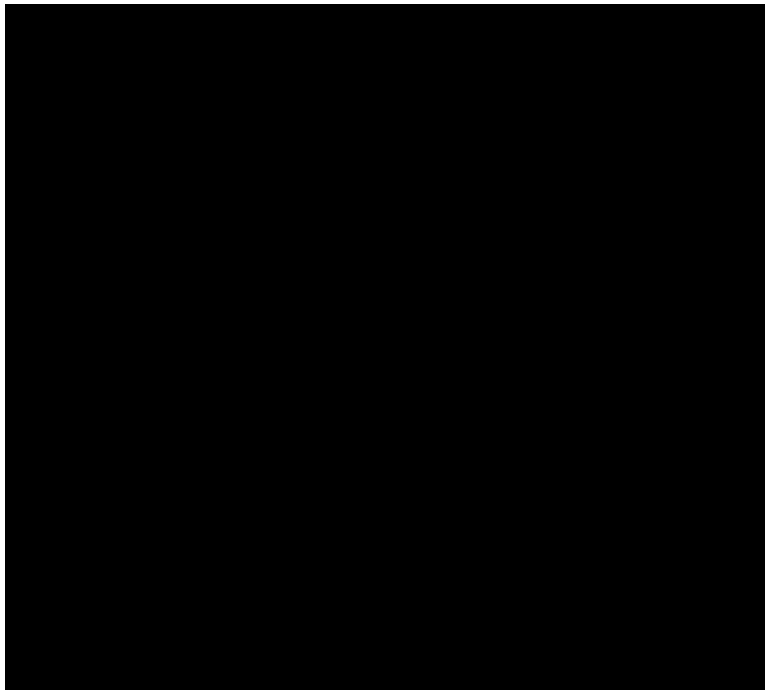


Example redshift sampling

J. Jasche, Bayesian LSS Inference

# Redshift sampling

## Preliminary Results



# Redshift sampling

## Preliminary Results

